# 3-DIMENSIONAL ISOTROPIC SUBMANIFOLDS OF SPHERES 

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## 1. Introduction

In this paper, we study 3-dimensional isotropic submanifolds in spheres. The notion of an isotropic submanifold of an arbitrary Riemannian manifold was first introduced by B. O'Neill in [O] ${ }_{1}$. The basic equations for isotropic submanifolds are recalled in Section 2.

Isotropic immersions of submanifolds into spheres have been studied by, amongst others, T. Itoh, H. Nakagawa, K. Ogiue and K. Sakamoto in [I], [N-I], [I-O] and [S]. Here, we will prove the two following theorems.

Theorem 3.1. Let $x: M \rightarrow S^{n}$ be a constant isotropic immersion such that $\operatorname{dim}(\operatorname{im}(h)) \leqq 3$. Then, one of the following holds:
(a) $M$ is totally geodesic in $S^{n}$,
(b) There exists a totally geodesic $S^{4}$ in $S^{n}$, such that the image of $M$ is an open part of a small hypersphere of $S^{4}$,
(c) There exists a totally geodesic $S^{7}$ in $S^{n}$, such that the image of $M$ is congruent with an open part of $j\left(\boldsymbol{R} \times S^{2}\left(\frac{\sqrt{2}}{\sqrt{3}}\right)\right)$ in $S^{7}$, where $j$ is defined in Section 3.

Theorem 3.2. Let $M$ be a 3-dimensional, minimal, isotropic submanifold in $S^{n}$. Then, $M$ has constant sectional curvature.

## 2. Preliminaries

In this section $M$ will always denote a 3 -dimensional totally real submanifold of $S^{n}(1)$. We will denote the curvature tensor of $M$ by $R$. The formulas of Gauss and Weingarten are given by

$$
\begin{equation*}
D_{X} Y=\nabla_{X} Y+h(X, Y) \text { and } D_{X} \zeta=-A_{\zeta} X+\nabla_{\bar{X}}^{\frac{1}{x}} \zeta, \tag{2.1}
\end{equation*}
$$

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where $X$ and $Y$ are tangent vector fields and $\zeta$ is a normal vector field on $M$. The space spanned by the image of $h$, will be called the first normal space. The equations of Gauss, Codazzi and Ricci for a submanifold of $S^{n}(1)$ are given by

$$
\begin{align*}
R(X, Y) Z= & \langle Y, Z\rangle X-\langle X, Z\rangle Y+A_{h(Y, z)} X-A_{h(x, z)} Y  \tag{2.2}\\
& (\nabla h)(X, Y, Z)=(\nabla h)(Y, X, Z)  \tag{2.3}\\
& \left\langle R^{\perp}(X, Y) \zeta, \eta\right\rangle=\left\langle\left[A_{\zeta}, A_{\eta}\right] X, Y\right\rangle \tag{2.4}
\end{align*}
$$

for tangent (resp. normal) vector fields $X, Y$ and $Z$ (resp. $\zeta$ and $\eta$ ) and $R^{\perp}$ (resp. $\widetilde{R}$ ) denotes the curvature tensor of $\nabla^{\perp}$ (resp. $D$ ).

From now on, we will also assume that $M$ is an isotropic submanifold, i.e. in each point $p$ of $M,\|h(v, v)\|$ is independent of the unit vector $v$. Hence, we obtain a function $\lambda$ on $M$ by

$$
\begin{equation*}
\lambda(p)=\|h(v, v)\| \tag{2.5}
\end{equation*}
$$

where $v \in U M_{p}$. If the function $\lambda$ is also independent of the point $p$, we say that $M$ is constant isotropic. In that case, we obtain from [ O$]_{1}$ the following conditions for orthonormal tangent vectors $X, Y, Z$ and $W$ :

$$
\begin{gather*}
\langle h(X, Y), h(X, X)\rangle=0,  \tag{2.6}\\
\lambda^{2}-\langle h(X, X), h(Y, Y)\rangle-2\langle h(X, Y), h(X, Y)\rangle=0,  \tag{2.7}\\
\langle h(Y, Z), h(X, X)\rangle+2\langle h(X, Y), h(X, Z)\rangle=0,  \tag{2.8}\\
\langle h(X, Y), h(Z, W)\rangle+\langle h(X, Z), h(W, Y)\rangle+\langle h(X, W), h(Y, Z)\rangle=0 . \tag{2.9}
\end{gather*}
$$

## 3. Proof of the theorems

Let $M$ be a 3 -dimensional, isotropic submanifold of $S^{n}(1)$ and let $p \in M$. Then, we choose an orthonormal basis $\left\{e_{1}, e_{2}, e_{3}\right\}$ of $T_{p} M$ in the following way. Let $S=\left\{(u, v) \mid u, v \in T_{p} M\right.$ with $\langle u, v\rangle=0$ and $\left.\|u\|=\|v\|=1\right\}$. We define a function $f$ on $S$ by

$$
f((u, v))=\|h(u, v)\|^{2} .
$$

Since $S$ is compact, we can choose ( $e_{1}, e_{2}$ ) as a point in which the function $f$ attains a maximum. To conclude the choice of our basis, we choose $e_{3}$ such that $e_{3}$ is orthogonal to both $e_{1}$ and $e_{2}$. Since $\left(e_{1}, e_{2}\right)$ is an absolute maximum we obtain that

$$
\begin{aligned}
& \left\langle h\left(e_{1}, e_{2}\right), h\left(e_{1}, e_{3}\right)\right\rangle=0, \\
& \left\langle h\left(e_{1}, e_{2}\right), h\left(e_{2}, e_{3}\right)\right\rangle=0 .
\end{aligned}
$$

Lemma 3.1. Let $M$ be a 3-dimensional isotropic submanifold with $\operatorname{dim}(\operatorname{im}(h))$ $\leq 3$ and let $p \in M$. Then there exists an orthonormal basis $\left\{e_{2}, e_{2}, e_{3}\right\}$ of $T_{p} M$ such that one of the following holds.

$$
\begin{align*}
& h\left(e_{1}, e_{1}\right)=h\left(e_{2}, e_{2}\right)=h\left(e_{3}, e_{3}\right)=0 .  \tag{3.1}\\
& h\left(e_{1}, e_{3}\right)=h\left(e_{1}, e_{3}\right)=h\left(e_{2}, e_{3}\right)=0,
\end{align*}
$$

or

$$
\begin{align*}
& h\left(e_{1}, e_{1}\right)=h\left(e_{2}, e_{2}\right)=h\left(e_{3}, e_{3}\right)=0,  \tag{3.2}\\
& h\left(e_{1}, e_{2}\right)=h\left(e_{1}, e_{3}\right)=h\left(e_{2}, e_{3}\right)=0,
\end{align*}
$$

or

$$
\begin{align*}
& h\left(e_{1}, e_{1}\right)=-h\left(e_{2}, e_{2}\right)=h\left(e_{3}, e_{3}\right)=\lambda g_{1},  \tag{3.3}\\
& h\left(e_{1}, e_{2}\right)=\lambda g_{2}, \\
& h\left(e_{1}, e_{2}\right)=0, \\
& h\left(e_{2}, e_{3}\right)=\lambda g_{3},
\end{align*}
$$

where $g_{1}, g_{2}, g_{3}$ are unit normal vectors at the point $p$ and $\lambda \neq 0$.
Proof. First, we assume that $\operatorname{dim}\left(\operatorname{im}\left(h_{p}\right)\right)=0$. This means that $p$ is a totally geodesic point. Therefore, we obtain (a).

Next, we assume that $\operatorname{dim}\left(\operatorname{im}\left(h_{p}\right)\right)=1$. Since $h$ is symmetric, this implies that $\lambda(p) \neq 0$. We choose an orthonormal basis of $T_{p} M$ as shown above. Then, it follows from the first isotropy condition (2.6) that $h\left(e_{1}, e_{2}\right)$ is orthogonal to $h\left(e_{1}, e_{1}\right)$. Since $\operatorname{dim}\left(\operatorname{im}\left(h_{p}\right)\right)=1$ and $\lambda(p) \neq 0$ this implies that $h\left(e_{1}, e_{2}\right)=0$. Similarly, we also obtain that $h\left(e_{1}, e_{3}\right)=h\left(e_{2}, e_{3}\right)=0$.

From the second isotropy condition it then follows that

$$
0=\lambda^{2}-\left\langle h\left(e_{1}, e_{1}\right), h\left(e_{2}, e_{2}\right)\right\rangle .
$$

Hence, since $M$ is isotropic and $\lambda \neq 0$, we obtain by applying the Cauchy Schwartz inequality that $h\left(e_{1}, e_{1}\right)=h\left(e_{2}, e_{2}\right)$. Similarly, we also obtain that $h\left(e_{3}, e_{3}\right)=h\left(e_{1}, e_{1}\right)$. This proves (b).

Next, we assume that $\operatorname{dim}\left(\operatorname{im}\left(h_{p}\right)\right)=2$. First, we assume that the function $f$ defined above is identically zero. In this case, we obtain similar as in the previous case that (b) holds. This is in contradiction with the assumption that $\operatorname{dim}\left(\operatorname{im}\left(h_{p}\right)\right)=2$. Therefore $f$ is not identically zero. Thus, if we choose an orthonormal basis indicated above, we obtain that $\left\|h\left(e_{1}, e_{2}\right)\right\|=\mu \neq 0$. Therefore $h\left(e_{1}, e_{1}\right)$ and $h\left(e_{1}, e_{2}\right)$ span the first normal space at the point $p$. By our choice of orthonormal basis and by the isotropy conditions, we know that $h\left(e_{1}, e_{3}\right)$ and $h\left(e_{2}, e_{3}\right)$ are orthogonal to $h\left(e_{1}, e_{1}\right)$ and $h\left(e_{1}, e_{2}\right)$. Thus $h\left(e_{1}, e_{3}\right)=h\left(e_{2}, e_{3}\right)=0$.

From the isotropy conditions and the Cauchy-Schwartz ineqality it then follows that $h\left(e_{1}, e_{1}\right)=h\left(e_{3}, e_{3}\right)=h\left(e_{2}, e_{2}\right)$. By applying then again (2.7), we obtain that $\mu=0$, which is again a contradiction.

Finally, we assume that $\operatorname{dim}(\operatorname{im}(h))=3$. By a similar argument as in the previous case, we obtain that the function $f$ is not identically zero. Therefore, if we choose an orthonormal basis in the same way as in the previous case, we have that $\left\|h\left(e_{1}, e_{2}\right)\right\|=\mu \neq 0$. Then, we obtain from the isotropy conditions that

$$
\left\langle h\left(e_{1}, e_{3}\right), h\left(e_{2}, e_{2}\right)\right\rangle=-2\left\langle h\left(e_{1}, e_{2}\right), h\left(e_{2}, e_{3}\right)\right\rangle=0 .
$$

Thus we see that $h\left(e_{1}, e_{3}\right)$ and $h\left(e_{2}, e_{3}\right)$ are orthogonal to $h\left(e_{1}, e_{1}\right), h\left(e_{2}, e_{2}\right)$ and $h\left(e_{1}, e_{2}\right)$. If $h\left(e_{1}, e_{1}\right), h\left(e_{2}, e_{2}\right)$ and $h\left(e_{1}, e_{2}\right)$ span the first normal space, we obtain that $h\left(e_{1}, e_{3}\right)=h\left(e_{2}, e_{3}\right)=0$. From this, we obtain in the same way as in the previous case that $\mu=0$. Therefore, we may assume that $h\left(e_{1}, e_{1}\right), h\left(e_{1}, e_{2}\right)$ and $h\left(e_{2}, e_{2}\right)$ are linearly dependent. The first isotropy condition then implies that $h\left(e_{2}, e_{2}\right)$ only has a component in the direction of $h\left(e_{1}, e_{1}\right)$ and the second isotropy condition then implies that $\mu=\lambda$ and that $h\left(e_{1}, e_{1}\right)=-h\left(e_{2}, e_{2}\right)$. These formulas imply that there exist orthonormal normal vectors $f_{1}, f_{2}$ and $f_{3}$ such that

$$
\begin{aligned}
& h\left(e_{1}, e_{1}\right)=\lambda f_{1}, \\
& h\left(e_{2}, e_{2}\right)=-\lambda f_{1}, \\
& h\left(e_{1}, e_{2}\right)=\lambda f_{2}, \\
& h\left(e_{1}, e_{3}\right)=\nu_{1} f_{3}, \\
& h\left(e_{2}, e_{3}\right)=\nu_{2} f_{3}, \\
& h\left(e_{3}, e_{3}\right)=\alpha f_{1}+\beta f_{2} .
\end{aligned}
$$

Then, the isotropy conditions are equivalent with

$$
\begin{aligned}
& 2 \nu_{1}^{2}=\lambda^{2}-\alpha \lambda, \\
& 2 \nu_{2}^{2}=\lambda^{2}+\alpha \lambda, \\
& \beta \lambda+2 \nu_{1} \nu_{2}=0 .
\end{aligned}
$$

From the first two equations, we see that we can put $\nu_{1}=\sin (\theta) \lambda$ and $\nu_{2}=\cos (\theta) \lambda$. But then it is clear from the last two equations that $\alpha=\left(2 \cos ^{2}(\theta)-1\right) \lambda$ and $\beta=-2 \sin (\theta) \cos (\theta) \lambda$. But then if we put $u_{3}=e_{3}, g_{3}=f_{3}, g_{1}=\cos (2 \theta) f_{1}-\sin (2 \theta) f_{2}$, $g_{2}=\cos (2 \theta) f_{2}+\sin (2 \theta) f_{1}, \quad u_{1}=\cos (\theta) e_{1}-\sin (\theta) e_{2} \quad$ and $\quad u_{2}=\sin (\theta) e_{1}+\cos (\theta) e_{2} \quad$ we obtain (c). This completes the proof of the lemma.

## Lemma 3.2. Let $M$ be as in Lemma 3.1.

(a) If (3.1) holds at a point of $M$, then $K_{p} \equiv 1$,
(b) If (3.2) holds at a point $p$ of $M$, then $K_{p} \equiv 1+\lambda^{2}$,
(c) If (3.3) holds at a point $p$ of $M$, then the sectional curvatures $K$ of $M$ at the point $p$ satisfy $1-2 \lambda^{2} \leqq K \leqq 1+\lambda^{2}$. Furthermore, $K_{p}=1-2 \lambda^{2}$ for every plane through $e_{2}$ and $K_{p}=1+\lambda^{2}$ only for the plane through $e_{1}$ and $e_{3}$.

Proof. (a) and (b) immediately follow from (3.1) and (3.2). To prove (c), we take an arbitrary tangent plane $\sigma$ at $p$. Then, we can find an orthonormal basis $\{X, Y\}$ of $\sigma$ such that $X=\cos \theta e_{1}+\sin \theta e_{3}$ and $Y=-\cos \phi \sin \theta e_{1}+\sin \phi e_{2}$ $+\cos \phi \cos \theta e_{3}$, where $\theta, \phi \in \boldsymbol{R}$. Then

$$
\begin{aligned}
\langle R(X, Y) Y, X\rangle= & \cos ^{2} \theta\left\langle R\left(e_{1}, Y\right) Y, e_{1}\right\rangle+2 \cos \theta \sin \theta\left\langle R\left(e_{1}, Y\right) Y, e_{3}\right\rangle \\
& +\sin ^{2} \theta\left\langle R\left(e_{3}, Y\right) Y, e_{3}\right\rangle \\
= & \left(1+\lambda^{2}\right) \cos ^{2} \phi-3 \lambda^{2} \sin ^{2} \phi .
\end{aligned}
$$

From this formula, (c) follows immediately.
Let us now assume that $M$ is constant isotropic, i. e. $\lambda$ is a constant on $M$. Then it follows from Lemma 3.1, Lemma 3.2 and the connectedness of $M$ that either
(a) (3.1) holds everywhere on $M$, i.e. $M$ is totally geodesic, or
(b) (3.2) holds everywhere on $M$, i.e. $M$ is totally umbilical, or
(c) (3.3) holds everywhere on $M$.

Totally geodesic and totally umbilical submanifolds of spheres are well known ([C], $[\mathrm{O}]_{2}$ ). So the only case we still have to consider is the case that (3.3) holds everywhere on $M$ with $\lambda \neq 0$. Let $p \in M$. Since in that case the sectional curvature equals $1+\lambda^{2}$ only for the plane through $e_{1}$ and $e_{3}$, we see that at each point $p$ the vector $e_{2}$ is uniquely determined, namely $e_{2}$ is the vector orthogonal to the unique plane with sectional curvature $1+\lambda^{2}$. From this it follows that we can choose differentiable vector fields $E_{1}, E_{2}, E_{3}$, defined on a neighbourhood $U$ of $p$, such that $\left\{E_{1}(q), E_{2}(q), E_{3}(q)\right\}$ satisfies (3.3) for every $q \in U$. Therefore, we also obtain orthonormal normal vector fields $g_{1}, g_{2}$ and $g_{3}$ such that

$$
\begin{aligned}
& h\left(E_{1}, E_{1}\right)=-h\left(E_{2}, E_{2}\right)=h\left(E_{3}, E_{3}\right)=\lambda g_{1}, \\
& h\left(E_{1}, E_{2}\right)=\lambda g_{2},
\end{aligned}
$$

$$
\begin{aligned}
& h\left(E_{1}, E_{3}\right)=0, \\
& h\left(E_{2}, E_{3}\right)=\lambda g_{3} .
\end{aligned}
$$

Then, we have the following lemma.
Lemma 3.3. Let us assume that (3.3) holds on $M$, where $\lambda$ is a constant different from zero. Then $\lambda=\frac{1}{\sqrt{2}}, M$ is locally isometric with $\boldsymbol{R} \times S^{2}\left(\frac{\sqrt{2}}{\sqrt{3}}\right)$ and after identification $E_{2}$ is tangent to $\boldsymbol{R}$. Further, $\nabla_{X} E_{2}=0$, for every tangent vector field $X$ and we can still choose locally $E_{1}$ and $E_{3}$ in such a way that they satisfy (3.3) and such that $\nabla_{E_{2}} E_{1}=\nabla_{E_{2}} E_{3}=0$. Finally, $(\nabla h)$ is orthogonal to $\mathrm{im}(h)$.

Proof. Let $p \in M$. First, we will prove that $(\nabla h) \perp h$. In order to do so, it is sufficient to prove that $(\nabla h)$ is orthogonal to $g_{1}, g_{2}$ and $g_{3}$. In order to do so, we first extend $e_{1}, e_{2}, e_{3}$ to local orthonormal vector fields $U_{1}, U_{2}, U_{3}$ such that $U_{i}(p)=e_{i}$ and $\nabla_{e_{j}} U_{i}=0$, where $i, j \in M$. Since $M$ is constant isotropic, we know that

$$
\left\langle h\left(U_{i}, U_{i}\right), h\left(U_{i}, U_{i}\right)\right\rangle=\lambda^{2}
$$

where $i \in\{1,2,3\}$. By differentiating this we obtain that

$$
\begin{equation*}
\left\langle(\nabla h)\left(e_{j}, e_{i}, e_{i}\right), h\left(e_{i}, e_{i}\right)\right\rangle=0 . \tag{3.4}
\end{equation*}
$$

Similarly, using the previous equations, we find from $\left\langle h\left(U_{i}, U_{j}\right), h\left(U_{i}, U_{i}\right)\right\rangle=0$, for different $i$ and $j$, that

$$
\begin{equation*}
\left\langle(\nabla h)\left(e_{k}, e_{i}, e_{j}\right), h\left(e_{i}, e_{i}\right)\right\rangle+\left\langle(\nabla h)\left(e_{k}, e_{i}, e_{i}\right), h\left(e_{i}, e_{j}\right)\right\rangle=0 . \tag{3.5}
\end{equation*}
$$

So, if we take $i=1, j=3$ and $k=2$, we find that

$$
\begin{equation*}
\left\langle(\nabla h)\left(e_{2}, e_{1}, e_{3}\right), g_{1}\right\rangle=0 . \tag{3.6}
\end{equation*}
$$

Then (3.4) together with (3.6) implies that $(\nabla h) \perp g_{1}$. Therefore, (3.5) becomes

$$
\left\langle(\nabla h)\left(e_{k}, e_{i}, e_{i}\right), h\left(e_{i}, e_{j}\right)\right\rangle=0,
$$

where $i, j, k \in\{1,2,3\}, i \neq j$. Since (3.3) holds at the point $p$ this implies that

$$
\begin{aligned}
& \left\langle(\nabla h)\left(e_{k}, e_{2}, e_{2}\right), g_{2}\right\rangle=\left\langle(\nabla h)\left(e_{k}, e_{1}, e_{1}\right), g_{2}\right\rangle=0, \\
& \left\langle(\nabla h)\left(e_{k}, e_{2}, e_{2}\right), g_{3}\right\rangle=\left\langle(\nabla h)\left(e_{k}, e_{3}, e_{3}\right), g_{3}\right\rangle=0 .
\end{aligned}
$$

Then, if we take the local orthonormal frame $\left\{E_{1}, E_{2}, E_{3}\right\}$ previously defined, we find that

$$
\left\langle h\left(E_{1}, E_{2}\right), h\left(E_{1}, E_{2}\right)\right\rangle=\lambda^{2} .
$$

Hence by deriving this with respect to $E_{3}$, we obtain that

$$
\begin{aligned}
0= & \left\langle(\nabla h)\left(E_{3}, E_{1}, E_{2}\right), h\left(E_{1}, E_{2}\right)\right\rangle+\left\langle h\left(\nabla_{E_{3}} E_{1}, E_{2}\right), h\left(E_{1}, E_{2}\right)\right\rangle \\
& +\left\langle h\left(E_{1}, \nabla_{E_{3}} E_{2}\right), h\left(E_{1}, E_{2}\right)\right\rangle .
\end{aligned}
$$

Thus $(\nabla h)\left(E_{3}, E_{1}, E_{2}\right)$ is also orthogonal to $g_{2}$. Similarly, starting from $\left\langle h\left(E_{2}, E_{3}\right), h\left(E_{2}, E_{3}\right)\right\rangle=\lambda^{2}$, we obtain that $(\nabla h)\left(E_{1}, E_{2}, E_{3}\right)$ is also orthogonal to $g_{3}$. Using similar arguments, we also can prove that

$$
\begin{aligned}
& \left\langle(\nabla h)\left(E_{1}, E_{3}, E_{3}\right), g_{2}\right\rangle=\left\langle(\nabla h)\left(E_{2}, E_{3}, E_{3}\right), g_{2}\right\rangle=0, \\
& \left\langle(\nabla h)\left(E_{1}, E_{1}, E_{3}\right), g_{3}\right\rangle=\left\langle(\nabla h)\left(E_{2}, E_{1}, E_{1}\right), g_{3}\right\rangle=0 .
\end{aligned}
$$

Therefore in order to prove that $(\nabla h) \perp h$, it only remains to prove that $(\nabla h)\left(E_{3}, E_{3}, E_{3}\right)$ is orthogonal to $g_{2}$ and that $(\nabla h)\left(E_{1}, E_{1}, E_{1}\right)$ is orthogonal to $g_{3}$. In particular, we already know that $(\nabla h)\left(E_{k}, E_{1}, E_{3}\right)$ is orthogonal to im $(h)$ for every $k$. But on the other hand, we have that

$$
\begin{aligned}
(\nabla h)\left(E_{k}, E_{1}, E_{3}\right)= & -h\left(\nabla_{E_{k}} E_{1}, E_{3}\right)-h\left(E_{1}, \nabla_{E_{k}} E_{3}\right) \\
= & -\left\langle\nabla_{E_{k}} E_{1}, E_{2}\right\rangle \lambda g_{3}-\left\langle E_{2}, \nabla_{E_{k}} E_{3}\right\rangle \lambda g_{2} \\
& -\left(\left\langle\nabla_{E_{k}} E_{1}, E_{3}\right\rangle+\left\langle\nabla_{E_{k}} E_{3}, E_{1}\right\rangle\right) \lambda g_{1}
\end{aligned}
$$

Thus $\nabla_{E_{k}} E_{2}=0$, where $k=1,2,3$. But then if follows by deriving $\left\langle h\left(E_{3}, E_{3}\right), h\left(E_{1}, E_{2}\right)\right\rangle=0$ and $\left\langle h\left(E_{1}, E_{1}\right), h\left(E_{2}, E_{3}\right)\right\rangle=0$ that also $(\nabla h)\left(E_{3}, E_{3}, E_{3}\right)$ is orthogonal to $g_{2}$ and that $(\nabla h)\left(E_{1}, E_{1}, E_{1}\right)$ is orthogonal to $g_{3}$. So $(\nabla h) \perp h$.

Since $\nabla_{E_{k}} E_{2}=0, k=1,2,3$, it follows that $R\left(E_{1}, E_{2}\right) E_{2}=0$. Hence from the Gauss equation we find that $0=1-2 \lambda^{2}$. Thus $\lambda=\frac{1}{\sqrt{2}}$.

Now, we can define two orthogonal distributions $T_{1}$ and $T_{2}$ by

$$
\begin{aligned}
& T_{1}: p \longmapsto T_{1}(p)=\operatorname{vect}\left\{E^{z}(p)\right\}, \\
& T_{2}: p \longmapsto T_{2}(p)=\operatorname{vect}\left\{E_{1}(p), E_{3}(p)\right\}
\end{aligned}
$$

Since $\nabla_{E_{k}} E_{2}=0, k=1,2,3$, we find that $\nabla_{r_{2}} T_{1} \subset T_{1}, \nabla_{T_{1}} T_{1} \subset T_{1}$. Since $T_{1}$ and $T_{2}$ are orthogonal distributions, we find from the de Rham decomposition theorem $([\mathrm{K}-\mathrm{N}])$ that $M$ is locally isometric with $\boldsymbol{R} \times M^{\prime}$, where $T_{1}$ is tangent to $\boldsymbol{R}$ and $T_{2}$ is tangent to $M^{\prime}$. Since $M^{\prime}$ has constant Gaussian curvature $\frac{3}{2}$, we also have that $M^{\prime}$ is locally isometric with a sphere of radius $\frac{\sqrt{2}}{\sqrt{3}}$.

Finally, since $M$ is locally isometric with a product, it is clear that we can choose locally vector fields $E_{1}$ and $E_{3}$, orthogonal to $T_{1}$, such that $\nabla_{E_{2}} E_{1}=\nabla_{E_{2}} E_{3}$ $=0$.

In the following lemmas, we will compute the normal connection on $M$ and
prove that $M$ lies linearly full in a 7-dimensional sphere. Let $p \in M$ and let $\left\{E_{1}, E_{2}, E_{3}\right\}$ the local orthonormal basis given by Lemma 3.2. Then, if we define local functions $\alpha$ and $\beta$ on $M$, by

$$
\begin{aligned}
& \nabla_{E_{1}} E_{3}=\alpha E_{1}, \\
& \nabla_{E_{3}} E_{1}=\beta E_{3},
\end{aligned}
$$

we have the following lemmas.
Lemma 3.4. If we denote the corresponding orthonormal normal vector fields by $g_{1}, g_{2}$ and $g_{3}$ we obtain that

$$
\begin{aligned}
& \nabla_{E_{1}}^{\perp} g_{1}=\nabla_{E_{\mathrm{t}}}^{\perp} g_{1}=\nabla_{E_{2}}^{\perp} g_{2}=\nabla_{E_{2}}^{\perp} g_{3}=0, \\
& \nabla_{E_{1}}^{1} g_{3}=\alpha g_{2}, \\
& \nabla_{E_{3}} g_{2}=\beta g_{3} \text {, } \\
& \nabla_{\frac{1}{E_{2}}} g_{1}=f \text {, } \\
& \nabla_{E_{1}} g_{2}=-\alpha g_{3}+f, \\
& \nabla_{⿺_{E_{3}}} g_{3}=-\beta g_{2}+f,
\end{aligned}
$$

where $f$ is a normal vector field to $M$ which is also normal to $g_{1}, g_{2}$ and $g_{3}$.
Proof. First, we notice that

$$
(\nabla h)\left(E_{k}, E_{1}, E_{3}\right)=-h\left(\nabla_{E_{k}} E_{1}, E_{3}\right)-h\left(E_{1}, \nabla_{E_{k}} E_{3}\right)=0
$$

Therefore, if we put $k=1$ and apply the Codazzi equation, we obtain that

$$
0=\lambda \nabla_{\mathbb{E}_{3}}^{1} g_{1}-2 h\left(\nabla_{E_{3}} E_{1}, E_{1}\right)=\lambda \nabla_{\bar{E}_{3}}^{1} g_{1} .
$$

Similarly, if we put $k=3$, we obtain that $\nabla_{E_{1}}^{\perp} g_{1}=0$. Finally, if we put $k=2$, we find from the Codazzi equations that

$$
\begin{aligned}
& 0=\lambda \nabla_{E_{1}} g_{3}-\alpha h\left(E_{1}, E_{2}\right), \\
& 0=\lambda \nabla \frac{1}{E_{3}} g_{2}-\beta h\left(E_{2}, E_{3}\right) .
\end{aligned}
$$

From the Codazzi equation $(\nabla h)\left(E_{2}, E_{1}, E_{2}\right)=(\nabla h)\left(E^{\mathrm{I}}, E_{2}, E_{2}\right)$, we then obtain that

$$
\lambda \nabla_{E_{2}}^{\perp} g_{2}=-\lambda \nabla_{E_{1}}^{\perp} g_{1}=0 .
$$

Similarly, we obtain from the Codazzi equation $(\nabla h)\left(E_{2}, E_{3}, E_{2}\right)=(\nabla h)\left(E_{3}, E_{2}, E_{2}\right)$ that $\nabla_{E_{2}}^{\frac{1}{2}} g_{3}=0$. Then, we define a normal vector field $f$ by

$$
f=\nabla_{E_{2}}^{1} g_{1} .
$$

It follows from the fact that $(\nabla h)\left(E_{2}, E_{1}, E_{1}\right)$ is orthogonal to $\operatorname{im}(h)$ that the normal vector field $f$ is orthogonal to $g_{1}, g_{2}$ and $g_{3}$. Then, it follows from the Codazzi equations $(\nabla h)\left(E_{2}, E_{1}, E_{1}\right)=(\nabla h)\left(E_{1}, E_{2}, E_{1}\right)$ and $(\nabla h)\left(E_{2}, E_{3}, E_{3}\right)=$ $(\nabla h)\left(E_{3}, E_{2}, E_{3}\right)$ that

$$
\begin{aligned}
& \nabla_{⿺_{1}} g_{2}=-\alpha g_{3}+f, \\
& \nabla_{\frac{E_{3}}{}} g_{3}=-\beta g_{2}+f .
\end{aligned}
$$

This completes the proof of this lemma.
Lemma 3.5. Let $p \in M$ and let $E_{1}, E_{2}, E_{3}, g_{1}, g_{2}, g_{3}$ and $f$ be local vector fields (normal or tangent) as defined above. Then
(i) $\langle f, f\rangle=1$
(ii) $\nabla_{\frac{1}{E_{1}}} f=-g_{2}$,

$$
\nabla_{E_{3}}^{\perp} f=-g_{3} .
$$

Proof. First, we take a local normal vector field $\eta$ which is orthogonal to $\operatorname{im}(h)$, i.e. which is orthogonal to $g_{1}, g_{2}$ and $g_{3}$. Since $\eta$ is orthogonal to $\operatorname{im}(h)$, it follows from the Ricci equation that

$$
\begin{align*}
& \left\langle R^{\perp}\left(E_{1}, E_{2}\right) g_{1}, \eta\right\rangle=0,  \tag{3.7}\\
& \left\langle R^{\perp}\left(E_{2}, E_{3}\right) g_{1}, \eta\right\rangle=0,  \tag{3.8}\\
& \left\langle R^{\perp}\left(E_{2}, E_{1}\right) g_{2}, \eta\right\rangle=0 . \tag{3.9}
\end{align*}
$$

On the other hand, using Lemma 3.4, we find that

$$
\begin{aligned}
R^{\perp}\left(E_{1}, E_{2}\right) g_{1} & =\nabla_{E_{1}}^{\frac{1}{E_{2}}} \nabla_{1}^{\perp} g_{1}-\nabla_{E_{2}}^{\perp} \nabla_{E_{1}}^{\perp} g_{1}-\nabla_{\left[E_{1}, E_{2}\right]}^{\perp} g_{1} \\
& =\nabla_{\stackrel{L}{1}^{\prime}}^{\perp} f .
\end{aligned}
$$

Combining this with (3.7) and Lemma 3.4, we find that

$$
\nabla_{E_{1}}^{1} f=-\langle f, f\rangle g_{2} .
$$

Similarly, we find from (3.8) and (3.9) that

$$
\begin{aligned}
& \nabla_{E_{2}} f=-\langle f, f\rangle g_{1}, \\
& \nabla_{\frac{⿺_{3}}{3}} f=-\langle f, f\rangle g_{3} .
\end{aligned}
$$

From these formulas, it is immediately clear that $f$ has constant length. But, we have first that

$$
\left\langle R^{\perp}\left(E_{1}, E_{2}\right) g_{1}, g_{2}\right\rangle=\left\langle\nabla_{E_{1}}^{\perp} f, g_{2}\right\rangle=-\langle f, f\rangle .
$$

On the other hand, by applying the Ricci identity, we obtain that

$$
\begin{aligned}
\left\langle R^{\perp}\left(E_{1}, E_{2}\right) g_{1}, g_{2}\right\rangle & =\left\langle\left[A_{g_{1}}, A_{g_{2}}\right] E, E_{2}\right\rangle \\
& =\left\langle A_{g_{2}} E_{1}, A_{g_{1}} E_{2}\right\rangle-\left\langle A_{g_{1}} E_{1}, A_{g_{2}} E_{2}\right\rangle \\
& =-\lambda^{2}-\lambda^{2}=-1 .
\end{aligned}
$$

This completes the proof of this lemma.
From these lemmas, it is clear that $(\nabla h)$ only has a component in the direction of $f$. Furthermore, by applying all these lemmas, we see that the space spanned by $\operatorname{im}(h)$ and $\operatorname{im}(\nabla h)$ is parallel with respect to the normal connection and has constant dimension 4. Therefore, by the reduction theorem of J. Erbacher [E], there exists a totally geodesic $S^{7}$ of $S^{n}$, such that $M$ is contained in $S^{7}$. The following example then shows that this case is possible.

Example 3.1. Let us consider $S^{7}$ as a hypersurface in $C^{4}$. Then, it is well-known that starting from the complex structure on $S^{7}$, one can induce a Sasakian structure $\phi$ on $S^{7}$ with structure vector field $\zeta$. Then, we consider the following immersion $j$ from $\boldsymbol{R} \times S^{2}\left(\frac{\sqrt{2}}{\sqrt{3}}\right)$ into $S^{7}$ :

$$
j\left(u, y_{1}, y_{2}, y_{3}\right)=\left(j_{1}, j_{2}, j_{3}, j_{4}, j_{5}, j_{6}, j_{7}, j_{8}\right),
$$

where

$$
\begin{aligned}
& j_{1}=\frac{\sqrt{2}}{\sqrt{3}} \cos \left(\frac{1}{\sqrt{2}} u\right) y_{1}+\frac{1}{3} \cos (\sqrt{2} u) \\
& j_{2}=-\frac{\sqrt{2}}{\sqrt{3}} \sin \left(\frac{1}{\sqrt{2}} u\right) y_{1}+\frac{1}{3} \sin (\sqrt{2} u) \\
& j_{3}=\frac{1}{\sqrt{3}} \sin \left(\frac{1}{\sqrt{2}} u\right) y_{1}+\frac{\sqrt{2}}{3} \sin (\sqrt{2} u) \\
& j_{4}=\frac{1}{\sqrt{3}} \cos \left(\frac{1}{\sqrt{2}} u\right) y_{1}-\frac{\sqrt{2}}{3} \cos (\sqrt{2} u) \\
& j_{5}=\cos \left(\frac{1}{\sqrt{2}} u\right) y_{2} \\
& j_{6}=-\sin \left(\frac{1}{\sqrt{2}} u\right) y_{2} \\
& j_{7}=\cos \left(\frac{1}{\sqrt{2}} u\right) y_{3} \\
& j_{8}=-\sin \left(\frac{1}{\sqrt{2}} u\right) y_{3}
\end{aligned}
$$

where $y_{1}^{2}+y_{2}^{2}+y_{3}^{2}=\frac{2}{3}$. Then, a straightforward computation shows that $\zeta$ is a normal vector field on $\boldsymbol{R} \times S^{2}\left(\frac{\sqrt{2}}{\sqrt{3}}\right)$. Hence the immersion is a totally real immersion. Furthermore, if we choose coordinates in such a way that

$$
\begin{aligned}
& y_{1}=\frac{\sqrt{2}}{\sqrt{3}} \cos \left(\frac{\sqrt{3}}{\sqrt{2}} v\right) \cos \left(\frac{\sqrt{3}}{\sqrt{2}} w\right) \\
& y_{2}=\frac{\sqrt{2}}{\sqrt{3}} \sin \left(\frac{\sqrt{3}}{\sqrt{2}} v\right) \cos \left(\frac{\sqrt{3}}{\sqrt{2}} w\right) \\
& y_{3}=\frac{\sqrt{2}}{\sqrt{3}} \sin \left(\frac{\sqrt{3}}{\sqrt{2}} w\right)
\end{aligned}
$$

it is easy to check that $j$ is an isotropic immersion.
ThEORFM 3.1. Let $x: M \rightarrow S^{n}$ be a constant isotropic immersion such that $\operatorname{dim}(\operatorname{im}(h)) \leqq 3$. Then, one of the following holds:
(a) $M$ is totally geodesic in $S^{n}$,
(b) There exists a totally geodesic $S^{4}$ in $S^{n}$, such that the image of $M$ is an open part of a small hypersphere of $S^{4}$,
(c) There exists a totally geodesic $S^{7}$ in $S^{n}$, such that the image of $M$ congruent with an open part of $j\left(\boldsymbol{R} \times S^{2}\left(\frac{\sqrt{2}}{\sqrt{3}}\right)\right)$ in $S^{7}$.

Proof. From Lemma 3.1 and Lemma 3.2 it follows that either (3.1) or (3.2) or (3.3) holds on $M$. If (3.1) or (3.2) holds on $M, M$ is totally umbilical. Therefore, by the classification of totally umbilical submanifolds of spheres ([C]), we obtain (a) and (b). Therefore, we may assume that (3.3) holds on $M$. Then, we know from Lemma 3.3 that $M$ is locally isometric with $\boldsymbol{R} \times S^{2}\left(\frac{\sqrt{2}}{\sqrt{3}}\right)$. Further from the remark following Lemma 3.5, we know that there exists a totally geodesic $S^{7}$ in $S^{n}$ such that the image of $M$ is contained in $S^{7}$. Now, let $p \in M$ and let $U$ be a neighbourhood of $p$ on which $M$ is isometric with $\boldsymbol{R} \times S^{2}\left(\frac{\sqrt{2}}{\sqrt{3}}\right)$. Then, on $U$, we can consider the two following immersions


From Lemma 3.3, Lemma 3. 4 and Lemma 3.5, it then follows that we can apply the uniqueness theorem ( $[\mathrm{Sp}]$, volume 4). Thus there exists an isometry
$A$ of $S^{7}$ such that on $U$

$$
A \circ x=j .
$$

Since the immersion $j$ is an analytic immersion, it follows immediately that the isometry $A$ is indepenent of the point $p$.

Theorem 3.2. Let $M$ be a 3-dimensional, minimal, isotropic submanifold in $S^{n}(1)$. Then, $M$ has constant sectional curvature.

Proof. Let $p \in M$ and let us assume that $p$ is not a totally geodesic point. Then, it follows by combining Lemma 3.1 with the minimality of $M$ that $\operatorname{dim}(\operatorname{im}(h))$ is greater than or equal to 4 . On the other hand, it follows from the minimality of $M$ that $\operatorname{dim}(\operatorname{im}(h))$ is less than or equal to 5 . Let us choose, as indicated at the beginning of this section, an orthonormal basis $\left\{e_{1}, e_{2}, e_{3}\right\}$ of $T_{p} M$. Clearly $\operatorname{im}(h)$ is spanned by $h\left(e_{1}, e_{1}\right), h\left(e_{1}, e_{2}\right), h\left(e_{2}, e_{2}\right), h\left(e_{1}, e_{3}\right)$ and $h\left(e_{2}, e_{3}\right)$.

Since $h\left(e_{1}, e_{3}\right)$ is orthogonal to $h\left(e_{1}, e_{1}\right)$ and $h\left(e_{3}, e_{3}\right)$, it follows from the minimality $M$ that $h\left(e_{1}, e_{3}\right)$ is orthogonal to $h\left(e_{2}, e_{2}\right)$. Furthermore, by the choice of our basis, we also know that $h\left(e_{1}, e_{3}\right)$ is orthogonal to $h\left(e_{1}, e_{2}\right)$. Similarly, we also obtain that $h\left(e_{2}, e_{3}\right)$ is orthogonal to $h\left(e_{1}, e_{1}\right), h\left(e_{2}, e_{2}\right)$ and $h\left(e_{1}, e_{2}\right)$.

From (2.8) and the minimality of $M$, it than follows that

$$
\begin{aligned}
\left\langle h\left(e_{1}, e_{3}\right), h\left(e_{2}, e_{3}\right)\right\rangle & =-\frac{1}{2}\left\langle h\left(e_{1}, e_{2}\right), h\left(e_{3}, e_{3}\right)\right\rangle \\
& =\frac{1}{2}\left\langle h\left(e_{1}, e_{2}\right), h\left(e_{1}, e_{1}\right)\right\rangle+\frac{1}{2}\left\langle h\left(e_{1}, e_{2}\right), h\left(e_{2}, e_{2}\right)\right\rangle \\
& =0 .
\end{aligned}
$$

On the other hand, it follows from (2.7) and the Cauchy Schwartz inequality that $h\left(e_{1}, e_{3}\right) \neq 0 \neq h\left(e_{2}, e_{3}\right)$. By the choice of our basis, this implies that $h\left(e_{1}, e_{2}\right) \neq 0$.

By combining these information, we see that there orthonormal normal vector fields $g_{1}, g_{2}, g_{3}, g_{4}$ and a normal vector field $g_{5}$, which is orthogonal to $g_{1}, g_{2}, g_{3}$ and $g_{4}$ such that

$$
\begin{aligned}
& h\left(e_{1}, e_{1}\right)=\lambda g_{1}, \\
& h\left(e_{1}, e_{2}\right)=\mu_{1} g_{2}, \\
& h\left(e_{1}, e_{3}\right)=\mu_{2} g_{3}, \\
& h\left(e_{2}, e_{3}\right)=\mu_{3} g_{4}, \\
& h\left(e_{2}, e_{2}\right)=\mu_{4} g_{1}+g_{5} .
\end{aligned}
$$

Then, we find from (2.7) that

$$
\begin{aligned}
2 \mu_{2}^{2} & =\lambda^{2}-\left\langle h\left(e_{1}, e_{1}\right), h\left(e_{3}, e_{3}\right)\right\rangle \\
& =2 \lambda^{2}+\left\langle h\left(e_{1}, e_{1}\right), h\left(e_{2}, e_{2}\right)\right\rangle \\
& =2 \lambda^{2}+\lambda^{2}-2 \mu_{1}^{2} .
\end{aligned}
$$

Similarly, we find that

$$
2 \mu_{3}^{2}=3 \lambda^{2}-2 \mu_{1}^{2}
$$

The minimality and (2.7) also imply that

$$
\begin{aligned}
\lambda^{2} & =\left\langle h\left(e_{3}, e_{3}\right), h\left(e_{3}, e_{3}\right)\right\rangle \\
& =2 \lambda^{2}+2\left\langle h\left(e_{1}, e_{1}\right), h\left(e_{2}, e_{2}\right)\right\rangle \\
& =4 \lambda^{2}-4 \mu_{1}^{2} .
\end{aligned}
$$

Thus, we may assume that $\mu_{1}=\frac{\sqrt{3}}{2} \lambda$. Hence, we may assume that $\mu_{2}=\mu_{3}=\frac{\sqrt{3}}{2}$. From (2.7) it then follows that $\mu_{4}=-\frac{1}{2} \lambda$. Finally, it follows from $\left\langle h\left(e_{2}, e_{2}\right)\right.$, $\left.h\left(e_{2}, e_{2}\right)\right\rangle=\lambda^{2}$ that $\left\langle g_{5}, g_{5}\right\rangle=\frac{3}{4} \lambda^{2}$. We can summarize this as follows. There exists orthonormal normal vector fields $f_{1}, f_{2}, f_{3}, f_{4}$ and $f_{5}$ such that:

$$
\begin{aligned}
& h\left(e_{1}, e_{1}\right)=\lambda f_{1}, \\
& h\left(e_{1}, e_{2}\right)=\frac{\sqrt{3}}{2} \lambda f_{2}, \\
& h\left(e_{2}, e_{2}\right)=-\frac{\lambda}{2} f_{1}+\frac{\sqrt{3}}{2} \lambda f_{3}, \\
& h\left(e_{1}, e_{3}\right)=\frac{\sqrt{3}}{2} \lambda f_{4}, \\
& h\left(e_{2}, e_{3}\right)=\frac{\sqrt{3}}{2} \lambda f_{5}, \\
& h\left(e_{3}, e_{3}\right)=-\frac{\lambda}{2} f_{1}-\frac{\sqrt{3}}{2} \lambda f_{3} .
\end{aligned}
$$

Using the Gauss equation, we find from these formulas that

$$
\begin{aligned}
R\left(e_{1}, e_{2}\right) e_{3}=R\left(e_{2}, e_{3}\right) e_{1} & =R\left(e_{3}, e_{1}\right) e_{2}=0, \\
\left\langle R\left(e_{1}, e_{2}\right) e_{2}, e_{1}\right\rangle=\left\langle R\left(e_{1}, e_{3}\right) e_{3}, e_{1}\right\rangle & =\left\langle R\left(e_{2}, e_{3}\right) e_{3}, e_{2}\right\rangle=1-\frac{5}{4} \lambda^{2} .
\end{aligned}
$$

Hence, $M$ has constant sectional curvatures.

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