BEST CHEBYSHEV QUADRATURES

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Introduction. The Chebyshev quadrature

(1)
$$\int_{-1}^{1} x(s) ds \cong u \sum_{k=1}^{m} x(a_{k-1}),$$

will be considered. Following the method of Sard [1], it will be assumed that $d^{n-1}x/ds^{n-1}$ is absolutely continuous, and that the approximation is precise for degree $\leq n-1$. Under these assumptions, the remainder, or error term,

(2)
$$Rx = \int_{-1}^{1} x(s) dx - u \sum_{k=1}^{m} x(a_{k-1})$$

may be written in the form (see [1, p. 25])

(3)
$$Rx = \int_{-1}^{1} K(t) \frac{d^{n}x}{ds^{n}}(t)dt.$$

One possible appraisal of the magnitude of the error term is obtained by applying the Schwarz inequality to (3), obtaining

$$|Rx|^2 \leq \int_{-1}^1 \left[\frac{d^n x}{ds^n}(t) \right]^2 dt$$

where

(4)
$$J = \int_{-1}^{1} [K(t)]^2 dt.$$

Any L_p norm $(p \ge 1)$ of K(t) could be considered. The L_2 norm was chosen because of the resulting simplicity of the calculations.

This paper will obtain "best" Chebyshev quadratures in the sense of Sard, i.e., those which minimize J. We will require that $-1 \leq a_0 < a_1 < \cdots < a_{m-1} \leq 1$, and that the a_k be symmetric, i.e., $a_{i-1} = -a_{m-i}$, $i = 1, \dots, m$. Precision zero will be required in all cases, thus $n \geq 1$ and u = 2/m.

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The cases to be considered numerically are n = 1, all m; and n = 2, $2 \leq m \leq 11$. These choices of n allow minimization of J with no constraints on the a_k , under the above assumptions. The values of J obtained in each case will be compared with the corresponding value for the classical Chebyshev approximations, where the highest possible precision is required. For m = 8, 10, and 11, the classical Chebyshev approximations do not exist under our restrictions. For those values of m, Barnhill, Dennis, and Nielson [4] have obtained Chebyshev type quadratures, and the best approximations for n = 2 will be compared with them.

The functions K(t) and $J(a_0, \dots, a_{m-1})$. We define $r^{(i)} = r^i/i!$ and

$$\theta(t,s) = 1 - \varphi(t,s) = 1, \quad t < s,$$
$$= 0, \quad t \ge s.$$

Then, by Sard's Kernel Theorem [1, p. 25],

(5)
$$K(t) = -R[(s - t)^{(n-1)}\theta(t, s)] = R[(s - t)^{(n-1)}\varphi(t, s)],$$

for $t \neq a_0, \dots, a_{m-1}$. Here R acts on the argument only as a function of s.

We consider the two equivalent forms of the kernel K(t). We have

$$K(t) = -R[(s-t)^{(n-1)}\theta(t,s)]$$

= $-\int_{-1}^{1} (s-t)^{(n-1)}\theta(t,s)ds$
+ $u \sum_{k=1}^{m} (a_{k-1}-t)^{(n-1)}\theta(t,a_{k-1}),$

which, when simplified, yields

$$K(t) = (-1 - t)^{(n)}; \quad -1 \leq t < a_0,$$

$$= (-1 - t)^{(n)} + u \sum_{i=1}^{k} (a_{i-1} - t)^{(n-1)};$$

 $a_{k-1} < t < a_k, \quad k = 1, \cdots, m-1,$

(6)

$$= (-1 - t)^{(n)} + u \sum_{i=1}^{m} (a_{i-1} - t)^{(n-1)};$$

 $a_{m-1} < t \leq 1$.

Similarly,

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$$K(t) = R[(s-t)^{(n-1)}\varphi(t,s)]$$

= $\int_{-1}^{1} (s-t)^{(n-1)}\varphi(t,s)ds - u \sum_{k=1}^{m} (a_{k-1}-t)^{(n-1)}\varphi(t,a_{k-1}),$

which, when simplified, yields

$$K(t) = (1 - t)^{(n)} - u \sum_{i=1}^{m} (a_{i-1} - t)^{(n-1)}; \quad -1 \leq t < a_0,$$

$$(7) \qquad = (1 - t)^{(n)} - u \sum_{i=k+1}^{m} (a_{i-1} - t)^{(n-1)};$$

$$a_{k-1} < t < a_k, \quad k = 1, \dots, m-1$$

$$= (1 - t)^{(n)}; \quad a_{m-1} < t \leq 1.$$

Under the assumption of symmetry of the nodes it is easily verified, using equations (6) and (7), that $K(-t) = (-1)^n K(t)$. Thus

$$J = \int_{-1}^{1} [K(t)]^{2} dt = 2 \int_{-1}^{0} [K(t)]^{2} dt,$$

and we need only to minimize

$$J_1 = \int_{-1}^0 [K(t)]^2 dt.$$

Let m^* be the number of nodes in [-1, 0), i.e.,

$$m^* = m/2$$
, m even
= $(m - 1)/2$, m odd.

Then make the following definitions:

$$I_{0} = \int_{-1}^{a_{0}} [K(t)]^{2} dt,$$

$$I_{k} = \int_{a_{k-1}}^{a_{k}} [K(t)]^{2} dt, \quad k = 1, 2, \cdots, m^{*} - 1,$$

$$I_{m^{*}} = \int_{a_{m^{*}-1}}^{0} [K(t)]^{2} dt.$$

We have

$$J_1 = \sum_{k=0}^{m^*} I_k.$$

TABLE 1

		$J_{ m best} \ { m and} \ J_{ m cla}$	assical	
	n =	1	n =	2
m	$J_{ m b}$	Jc	$J_{ m b}$	Jc
2	1.67×10^{-1}	1.79×10^{-1}	4.08×10^{-3}	4.75×10^{-3}
3	7.41×10^{-2}	7.63×10^{-2}	$7.25 imes 10^{-4}$	7.80×10^{-4}
4	4.17×10^{-2}	4.76×10^{-2}	2.15×10^{-4}	4.19×10^{-4}
5	2.67×10^{-2}	2.80×10^{-2}	8.48×10^{-5}	1.15×10^{-4}
6	1.85×10^{-2}	2.99×10^{-2}	3.98×10^{-5}	2.28×10^{-4}
7	1.36×10^{-2}	1.58×10^{-2}	2.11×10^{-5}	5.29×10^{-5}
9	8.23×10^{-3}	1.49×10^{-2}	7.50×10^{-6}	6.11×10^{-5}

We must obtain a simplified expression for J_1 . Rewriting the I_k , we have

$$\begin{split} I_{0} &= \int_{-1}^{a_{0}} \left[(-1-t)^{(n)} \right]^{2} dt. \\ I_{k} &= \int_{a_{k-1}}^{a_{k}} \left[(-1-t)^{(n)} + u \sum_{i=1}^{k} (a_{i-1}-t)^{(n-1)} \right]^{2} dt \\ &= \int_{a_{k-1}}^{a_{k}} \left[(-1-t)^{(n)} \right]^{2} dt \\ &+ 2u \sum_{i=1}^{k} \int_{a_{k-1}}^{a_{k}} (-1-t)^{(n)} (a_{i-1}-t)^{(n-1)} dt \\ &+ u^{2} \sum_{i,j=1}^{k} \int_{a_{k-1}}^{a_{k}} (a_{i-1}-t)^{(n-1)} (a_{j-1}-t)^{(n-1)} dt, \\ &\quad k = 1, \cdots, m^{*} - 1. \\ I_{m^{*}} &= \int_{a_{m^{*}-1}}^{0} \left[(-1-t)^{(n)} + u \sum_{i=1}^{m^{*}} (a_{i-1}-t)^{(n-1)} \right]^{2} dt \\ &= \int_{a_{m^{*}-1}}^{0} \left[(-1-t)^{(n)} \right]^{2} dt \\ &+ 2u \sum_{i=1}^{m^{*}} \int_{a_{m^{*}-1}}^{0} (-1-t)^{(n)} (a_{i-1}-t)^{(n-1)} dt \\ &+ u^{2} \sum_{i,j=1}^{m^{*}} \int_{a_{m^{*}-1}}^{0} (a_{i-1}-t)^{(n-1)} dt. \end{split}$$

Then,

$$\begin{split} J_{1} &= \sum_{k=0}^{m^{*}} I_{k} = \int_{-1}^{0} \left[(-1-t)^{(n)} \right]^{2} dt \\ &+ 2u \left\{ \sum_{k=1}^{m^{*}-1} \sum_{i=1}^{k} \int_{a_{k-1}}^{a_{k}} (-1-t)^{(n)} (a_{i-1}-t)^{(n-1)} dt \right. \\ &+ \sum_{i=1}^{m^{*}} \int_{a_{m^{*}-1}}^{0} (-1-t)^{(n)} (a_{i-1}-t)^{(n-1)} dt \right\} \\ &+ u^{2} \left\{ \sum_{k=1}^{m^{*}-1} \sum_{i, j=1}^{k} \int_{a_{k-1}}^{a_{k}} (a_{i-1}-t)^{(n-1)} (a_{j-1}-t)^{(n-1)} dt \right. \\ &+ \sum_{i, j=1}^{m^{*}} \int_{a_{m^{*}-1}}^{0} (a_{i-1}-t)^{(n-1)} (a_{j-1}-t)^{(n-1)} dt \right\} \\ &= \int_{-1}^{0} \left[(-1-t)^{(n)} \right]^{2} dt \\ &+ 2u \sum_{k=1}^{m^{*}} \int_{a_{k-1}}^{0} (-1-t)^{(n)} (a_{k-1}-t)^{(n-1)} dt \\ &+ u^{2} \sum_{k=1}^{m^{*}} \int_{a_{k-1}}^{0} \left\{ \left[(a_{k-1}-t)^{(n-1)} \right]^{2} \\ &+ 2(a_{k-1}-t)^{(n-1)} \sum_{i=1}^{k-1} (a_{i-1}-t)^{(n-1)} \right\} dt. \end{split}$$

The sum in the last term is to be interpreted as zero when k - 1 = 0, as are any sums in the following work where the upper limit is smaller than the lower limit. Integrating the above expressions, by parts several times in the last two terms, one obtains

$$J_{1} = \frac{1}{(n!)^{2}(2n+1)} + 2u(-1)^{n+1} \sum_{k=1}^{m^{*}} \sum_{\ell=0}^{n} (1)^{(n-\ell)} a_{k-1}^{(n+\ell)}$$

$$(8) \qquad + u^{2} \sum_{k=1}^{m^{*}} \left[-\frac{a_{k-1}^{2n-1}}{[(n-1)!]^{2}(2n-1)} + 2\sum_{i=1}^{k-1} \sum_{\ell=1}^{n} (-1)^{\ell} a_{k-1}^{(n+\ell-1)} a_{i-1}^{(n-\ell)} \right]$$

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	TABLE 2						
	Nodes: $n = 2$, and classical case						
m	n=2		Classical [3, p. 920]				
2	$\pm .55051$	02572	$\pm .57735$	02692			
3	$\pm .69722$	43623	$\pm.70710$	67812			
	.0		.0				
4	$\pm .77315$	45618	$\pm .79465$	44723			
	$\pm .24721$	76225	$\pm .18759$	24741			
5	$\pm .81850$	45407	$\pm .83249$	74870			
	$\pm .39799$	68704	$\pm .37454$	14096			
	.0		.0				
6	$\pm .84875$	53923	$\pm .86624$	68181			
	$\pm .49831$	19555	$\pm .42251$	86538			
	$\pm .16685$	25367	$\pm .26663$	54015			
7	$\pm .87036$	16256	$\pm.88386$	17008			
	$\pm .56998$	33013	$\pm .52965$	67753			
	$\pm .28585$	74954	$\pm .32391$	18105			
	.0		.0				
9	$\pm .89917$	01631	$\pm .91158$	93077			
	$\pm .66554$	24529	$\pm .60101$	86554			
	$\pm.44455$	69674	$\pm .52876$	17831			
	$\pm .22221$	09597	$\pm.16790$	61842			
	.0		.0				

The case n = 1. Making the appropriate simplifications to (8), we find

(9)
$$J_1 = \frac{1}{3} + \frac{4}{m} \sum_{k=1}^{m^*} [a_{k-1}^{(2)} + a_{k-1}] + \frac{4}{m^2} \sum_{k=1}^{m^*} (-2k+1)a_{k-1}.$$

The a_{k-1} are independent in this case, so setting the partial derivatives to zero we obtain the equations

(10)
$$\frac{\partial J_1}{\partial a_{k-1}} = \frac{4}{m} (a_{k-1} + 1) + \frac{4}{m^2} (-2k+1) = 0,$$
$$a_{k-1} = \frac{2(k-1) - m + 1}{m}, \qquad k = 1, 2, \cdots, m^*.$$

This is the repeated midpoint rule, a result previously noted by Krylov [2, pp. 138-140]. The value of J_1 can be obtained by substituting (10) into (9), and after laborious simplification, it is found that

TABLE 3

 $J_{
m best}$ and $J_{
m Barnhill, \ Dennis, \ Nielson}$

m	$J_{ m b}$	$J_{ m B,D,N}$
8	1.22×10^{-5}	1.06×10^{-4}
10	4.87×10^{-6}	2.88×10^{-4}
11	3.30×10^{-6}	1.00×10^{-4}

TABLE 4

Nodes: m = 8, 10, 11; n = 2, and Barnhill, Dennis, Nielson

			Barnhill, Dennis,	
m	n=2		Nielson [4]	
8	$\pm .88656$	64347	$\pm .90044$	55323
	$\pm .62373$	52450	$\pm .55898$	89280
	$\pm .37512$	67322	$\pm.45850$	78272
	$\pm .12498$	59060	±.0	
10	$\pm .90925$	31469	$\pm .92106$	80558
	$\pm .69898$	82064	$\pm .63181$	13569
	$\pm .50010$	12823	$\pm .58191$	74110
	$\pm .29998$	97486	$\pm .38366$	17341
	$\pm .10000$	11389	0.±	
11	$\pm .91750$	28608	$\pm .92676$	50132
	$\pm .72635$	29150	$\pm .70492$	41194
	$\pm .54554$	66192	$\pm .51792$	29707
	$\pm .36362$	70548	$\pm.45740$	29197
	$\pm .18181$	91126	$\pm .0$	
	.0		.0	

 $J_1 = 1/3m^2$. Table 1 compares the values of J corresponding to (10) and to the classical Chebyshev approximations.

The case n = 2. Equation (8) reduces to

(11)
$$J_{1} = \frac{1}{20} - \frac{4}{m} \sum_{k=1}^{m^{*}} \left[\frac{a_{k-1}^{2}}{4} + \frac{a_{k-1}^{3}}{6} + \frac{a_{k-1}^{4}}{24} \right] + \frac{4}{m^{2}} \sum_{k=1}^{m^{*}} \left[\frac{k-2}{3} a_{k-1}^{3} - a_{k-1}^{2} \sum_{i=1}^{k-1} a_{i-1} \right].$$

There are no constraints on the a_{k-1} since symmetry assures precision n-1=1. Equating the partial derivatives to zero, we have the system of equations

$$\frac{\partial J_1}{\partial a_{k-1}} = -\frac{4}{m} \left[\frac{a_{k-1}}{2} + \frac{a_{k-1}^2}{2} + \frac{a_{k-1}^3}{6} \right] \\ + \frac{4}{m^2} \left[(k-2)a_{k-1}^2 - 2a_{k-1}\sum_{i=1}^{k-1} a_{i-1} - \sum_{i=k+1}^{m^*} a_{i-1}^2 \right] = 0$$

for
$$k = 1, \dots, m^*$$
.

Simplifying, we obtain the system

(12)
$$-\frac{6}{m}\left[(k-2)a_{k-1}^{2}-2a_{k-1}\sum_{i=1}^{k-1}a_{i-1}-\sum_{i=k+1}^{m^{*}}a_{i-1}^{2}\right]=0,$$
$$k=1,\cdots,m^{*}.$$

Using Newton's Method for systems, (12) was solved for the a_{k-1} for $2 \leq m \leq 9$. $J = 2J_1$ was evaluated for those values, as well as for the classical approximations. The results are shown in Tables 1 and 2.

During the computations on the system of equations (12), as expected, increasing m necessitates more accurate initial guesses at the solution. For m = 10 and 11, convergence for all initial guesses attempted was to repeated nodes. An alternative, and in fact simpler procedure than the above, is as follows. We want

$$\frac{\partial}{\partial a_{k-1}} \int_{-1}^{0} [K(t)]^2 dt = 2u \int_{a_{k-1}}^{0} K(t) dt = 0$$

where K(t) is given by equation (6), with n = 2. Since $\int_{a_{k-1}}^{0} K(t) dt = 0$ must hold for $k = 1, \dots, m^*$, we have $\int_{a_{k-1}}^{a_k} K(t) dt = 0$ for $k = 1, \dots, m^* - 1$. Simplifying, we find

$$\int_{a_{k-1}}^{a_{k}} K(t)dt = \int_{a_{k-1}}^{a_{k}} \left[\frac{(-1-t)^{2}}{2!} + u \sum_{i=1}^{k} (a_{i-1}-t) \right] dt$$

= $(a_{k} - a_{k-1}) \left[\frac{(1+a_{k})^{2} + (1+a_{k})(1+a_{k-1}) + (1+a_{k-1})^{2}}{6} + u \sum_{i=1}^{k} a_{i-1} - \frac{ku}{2} (a_{k} + a_{k-1}) \right]$
= 0

Because we require that $a_{k-1} < a_k$, the quantity in brackets must be

zero, and this then simplifies to

$$a_{k}^{2} + a_{k}(3 - 3ku + a_{k-1})$$

$$(13) + \left(3 + 3a_{k-1} + a_{k-1}^{2} + 6u \sum_{i=1}^{k} a_{i-1} - 3kua_{k-1}\right) = 0,$$

$$k = 1, 2, \cdots, m^{*} - 1.$$

We note that if the nodes a_0, \dots, a_{k-1} are known, a_k can be computed from (13). Equation (13) is a quadratic in a_k , hence has two solutions. For n = 2, K(t) can be made continuous by properly defining it at the a_k , and is piecewise parabolic, the parabolas all opening upward. We recall that the solutions of (13) correspond to $\int_{a_{k-1}}^{a_k} K(t) dt = 0$. Proceeding inductively on k, and assuming that the two solutions are real and distinct, the smaller solution corresponds to a point where K(t) < 0 and the larger to a point where K(t) > 0. Noting the shape of K(t) and that $dK/dt|_{a_k^-} > dK/dt|_{a_k^+}$, one is convinced that the larger of the two solutions leads to the smaller value of J_1 .

The numerical procedure is as follows. For an initial guess at a_0 , one can compute in turn, $a_1, a_2, \dots, a_{m^*-1}$ from equation (13). Then let

$$F(a_0) = \int_{a_{m^*-1}}^{0} K(t)dt$$

= $\frac{a_{m^*-1}}{6} \left[3 + 3a_{m^*-1} + a_{m^*-1}^2 + 6u \sum_{i=1}^{m^*} a_{i-1} - 3um^*a_{m^*-1} \right].$

From above, we must have $\int_{a_m^*-1}^0 K(t)dt = 0$, thus we want to solve the equation $F(a_0) = 0$. The procedure is quite sensitive, and some choices for a_0 result in complex nodes. However, searching for an appropriate initial guess is considerably simplified over that for equation (13), since only one variable is involved. This procedure was successfully applied for m = 10, 11, as well as for previously solved values, as a check on the two methods. The results for m = 8, 10, and 11 are shown in Tables 3 and 4, along with the approximations of Barnhill, Dennis, and Nielson.

Conclusion. Some observations may be made from the tables. First note that the classical approximation for m = 6 seems to be relatively poorer than the others, and in fact for the cases considered, the corresponding value for J exceeds that for m = 5. Without exception,

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the nodes move toward their classical position from their position for n = 1, when *n* is increased to 2.

The existence of approximations for values of m for which the classical approximations do not exist is noted for n = 1 and 2. It should be mentioned that the criterion used by Barnhill, Dennis, and Nielson to obtain their approximations is different from that used here. The approximations obtained here would not compare favorably against theirs, using their criterion.

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