# THE SECTIONAL CURVATURE OF A 5-DIMENSIONAL HARMONIC RIEMANNIAN MANIFOLD 

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In 1939, E. T. Copson and H. S. Ruse [4] initiated to study harmonic Riemannian manifolds. And, in 1944, A. Lichnerowicz [13] proved the curvature identities (cf. §2) in a harmonic Riemannian manifold and gave the following

Conjecture. If a Riemannaan manıfold $M$ with positive definite metric is harmonic, then $M$ is locally symmetric.

For Riemannian manifolds of dimension 2 or 3 the conjecture is trivially affirmative, because a harmonic manifold is Einsteinian and therefore of constant curvature. A.G. Walker [22] showed that the conjecture is affirmative for Riemannian manifolds of dimension 4 (cf. [16], [2]). Later, harmonic Riemannian manifolds were studied by T. J. Willmore [28], [29], A. J. Ledger [11], [12], A. Allamigeon [1], S. Tachibana [19], Y. Watanabe [23], [24], [25], [26], [27], A. Besse [2], L. Vanhecke [20], [21], M. Kôzaki [9], K. Sakamoto [18] and others. But it is an open problem to show that the Lichnerowicz's conjecture is affirmative for Riemannian manifolds of dimension $>4$, and no counterexample is known up to now. The main purpose of this paper is to prove the main Theorem 3.3 giving a sufficient condition, by pinching the sectional curvature at a point, for a harmonic Riemannian manifold of dimension 5 to be locally symmetric, i.e. of constant curvature. In § 3, we prove Lemma 3.2, which implies immediately the main theorem, because a locally symmetric harmonic manifold is locally flat or locally isometric to a rank one symmetric space (cf. [12], [6], [3]), i. e., because it is of constant curvature in the case where it is odd-dimensional.

In $\S 1$, we give some preliminaries concerning Riemannian manifolds. In $\S 2$, we give definitions and curvature conditions concerning harmonic Riemannian manifolds. The last section $\S 3$ is devoted to the proof of the main theorem and another.

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## § 1. Preliminaries.

First, we shall recall in Riemannian manifolds some curvature identities which will be useful in the sequel. In the present paper, every Riemannian manifold we consider is assumed to be of class $C^{\omega}$ and connected. Let $M$ be an $n$-dimensional Riemannian manifold with positive definite metric $g$ and $\nabla$ be the Levi-Civita connection. The Riemannian curvature tensor $R$ is defined by

$$
R(X, Y) Z=\nabla_{X} \nabla_{Y} Z-\nabla_{Y} \nabla_{X} Z-\nabla_{[X, Y]} Z
$$

for any vector fields $X, Y$ and $Z$. With local components the curvature tensor can be written as $R=\left(R_{k j i}{ }^{h}\right)$. Let $R_{1}=\left(R_{a j i}{ }^{a}\right)=\left(R_{j i}\right)$ and $S=g^{j i} R_{j i}$ be the Ricci tensor and the scalar curvature, respectively. Denote by $T_{x} M$ the tangent space to $M$ at a point $x$ of $M$. The sectional curvature $\kappa_{x}(X, Y)$ for an orthonormal pair $\{X, Y\}$, where $X$ and $Y$ belonging to $T_{x} M$, is given by

$$
\kappa_{x}(X, Y)=-g(R(X, Y) X, Y)
$$

We recall the well known Bianchi's identities
(a) $R_{k j i}{ }^{h}+R_{j i k}{ }^{h}+R_{\imath k j}{ }^{h}=0$,
(b) $\nabla_{l} R_{k j i}{ }^{h}+\nabla_{k} R_{j l_{2}}{ }^{h}+\nabla_{j} R_{l k_{2}}{ }^{h}=0$,
where $\nabla_{l}$ denotes the covariant differentiation with respect to the Levi-Civita connection. Generally speaking, we put $|T|^{2}=T_{k j i} T^{k j i}$ for any tensor field of any type, say $T=\left(T_{k j i}\right)$. Then from (1.1) we get the following well known formulas (cf. [17])

$$
\begin{align*}
\nabla^{u} R^{a b c d} \nabla^{u} R_{a d c b} & =\nabla_{u} R^{a b c d} \nabla_{u} R_{c b a d}=\nabla^{u} R^{a b c d} \nabla_{c} R_{a b u d}  \tag{1.2}\\
& =\nabla^{u} R^{a b c d} \nabla_{d} R_{a b c u}=\nabla^{u} R^{a b c d} \nabla_{a} R_{u b c d} \\
& =\nabla^{u} R^{a b c d} \nabla_{c} R_{u b a d}=\frac{1}{2}|\nabla R|^{2} .
\end{align*}
$$

On putting (cf. [2])

$$
\hat{R}=R^{a b c d} R_{a b}{ }^{u v} R_{c d u v} \text { and } \hat{R}=R^{a b c a} R_{a}{ }^{u}{ }_{c}^{v} R_{b u d v},
$$

we have the following formulas (cf. T. Sakai [17]):
(a) $R^{a b c d} R_{a b}{ }^{u v} R_{c u d v}=R^{a b c d} R_{a}{ }^{u}{ }_{b}{ }^{v} R_{c d u v}=R^{a b c d} R_{a c}{ }^{u v} R_{b d u v}=\frac{1}{2} \hat{R}$,
(b) $R^{a b c d} R_{a}{ }^{u}{ }_{b}^{v} R_{c u d v}=R^{a b c d} R_{a}{ }^{u}{ }_{b}{ }^{v} R_{c v u d}=R^{a b c d} R_{a}{ }^{u}{ }_{c}{ }^{v} R_{b d u v}$

$$
=R^{a b c d} R_{a c}{ }^{u v} R_{b u d v}=\frac{1}{4} \hat{R},
$$

(c) $\quad R^{a b c d} R_{a}{ }^{u}{ }_{c}{ }^{v} R_{b v d u}=R^{a b c d} R_{a}{ }^{v}{ }_{c}{ }^{u} R_{b u d v}=\hat{R}-\frac{1}{4} \hat{R}$.
A. Lichnerowicz [14] gave the following identity

$$
\begin{equation*}
\frac{1}{2} \Delta|R|^{2}=|\nabla R|^{2}-4 R^{j i n k} \nabla_{j} \nabla_{h} R_{\imath k}+2 R_{\imath \jmath} R^{i \hbar k l} R^{\jmath}{ }_{h k l}+\hat{R}+4 \hat{R}, \tag{1.4}
\end{equation*}
$$

where $\Delta$ is the Laplace-Beltrami operator acting on differentiable functions on $M$ (cf. [30]).

If $M$ is Einsteinian and $|R|^{2}=$ constant, then (1.4) reduces to

$$
\begin{equation*}
|\nabla R|^{2}+\frac{2}{n} S|R|^{2}+\hat{R}+4 \hat{R}=0 . \tag{1.5}
\end{equation*}
$$

## § 2. Harmonic Riemannian manifolds.

Let $M$ be a Riemannian manifold of dimension $n$. Take a normal neighborhood $N$ centered at a point $x_{0}$ of $M$. Denoting by $s(x)$ the geodesic distance measured from $x_{0}$ to a point $x$ of $N$, we define in $N$ a function $s$ by $s: x \rightarrow$ $s(x)(x \in N)$. Given a fixed point $x_{0}$ of $M$, the Riemannian manifold $M$ is said to be harmonic at the point $x_{0}$, if there is a normal neighborhood $U$ centered at $x_{0}$ in such a way that there is in $U-\left\{x_{0}\right\}$ a nontrivial solution $u$, analytically depending only on $\Omega=(1 / 2) s^{2}$, of the Laplace equation $\Delta u=0$. When $M$ is harmonic at every point of $M$, it is called a harmonic Rzemannıan manifold. It is well known (cf. [13], [16]) that in a harmonic Riemannian manifold the local function $\Delta \Omega$ has the form $f(\Omega)$ in each normal neighborhood $U$, where $f(\Omega)$ is a function analytically depending on $\Omega$, and the function $f(\Omega)$ is independent of the choice of the center $x_{0}$. The function $f(\Omega)$ is called the characteristic function of the harmonic Riemannian manifold.

Let $M$ be a harmonic Riemannian manifold of dimension $n$ and $\left\{y^{i}\right\}$ be a normal coordinate system, covering a sufficient small normal neighborhood $U$ centered at a point $x_{0}$ of $M$. Then $\Delta \Omega=f(\Omega)$ and $f(\Omega)$ thus admits the Maclaurin expansion

$$
\begin{align*}
f(\Omega) & =f(0)+\dot{f}(0) \Omega+\frac{1}{2!} \tilde{f}(0) \Omega^{2}+\frac{1}{3!} \ddot{f}(0) \Omega^{3}+\cdots  \tag{2.1}\\
& =f(0)+\frac{1}{2} \dot{f}(0) s^{2}+\frac{1}{2!2^{2}} \dot{f}(0) s^{4}+\frac{1}{3!2^{3}} \ddot{f}(0) s^{6}+\cdots,
\end{align*}
$$

taking account of $\Omega=(1 / 2) s^{2}$, where $(\cdot)$ means the operator taking the derivative with respect to $\Omega$. On the other hand, in any Riemannian manifold the formula

$$
\Delta \Omega=n+\left\{\begin{array}{l}
2  \tag{2.2}\\
2 k
\end{array}\right\} y^{k}
$$

holds with respect to normal coordinates $\left\{y^{k}\right\}$, where $\left\{\begin{array}{c}i \\ j k\end{array}\right\}$ denotes the Christoffel symbols. If $\left\{\begin{array}{c}l \\ l k\end{array}\right\}$ are expanded in Taylor expansion with respect to $y^{h}$, then
using (2.1) and (2.2), the following curvature conditions are obtained (cf. [13], [10], [16]):

$$
\begin{align*}
& R_{j i}=-\frac{3}{2} \dot{f}(0) g_{\jmath i}, \quad S=-\frac{3 n}{2} \dot{f}(0)  \tag{2.3}\\
& \Theta\left(R_{p \imath j}^{q} R_{q k l}{ }^{p}+\frac{45}{8} \dot{f}(0) g_{\imath \jmath} g_{k l}\right)=0 \tag{2.4}
\end{align*}
$$

$$
\begin{equation*}
\varsigma\left(9 \nabla_{k} R_{p \imath \jmath}^{q} \nabla_{l} R_{q m n}^{p}-32 R_{p \imath j}^{q} R_{q k l}^{r} R_{r m n}^{p}-315 \ddot{f}(0) g_{\imath \jmath} g_{k l} g_{m n}\right)=0 \tag{2.5}
\end{equation*}
$$

where $\subseteq$ means the summation taken over all permutation of the free indices appearing incide the parenthesis ( ). By the definition of harmonicity we see that $\dot{f}(0), \vec{f}(0)$ and $\vec{f}(0)$ are absolute constants, i. e. that they are independent of choice of the center $x_{0}$. Then transvecting $g^{2 g} g^{k l}$ with (2.4) and using (2.3), we obtain

$$
\begin{equation*}
|R|^{2}=-\frac{3 n}{2}\left\{\dot{f}(0)^{2}+\frac{5(n+2)}{2} \ddot{f}(0)\right\} \tag{2.6}
\end{equation*}
$$

and see that $|R|^{2}$ is constant.
We now need the following two lemmas for computing $\ddot{f}(0)$ in terms of the scalar functions constructed by the curvature tensors.

Lemma 2.1. For a tensor (field) $T=\left(T_{i j k l m n}\right)$ of type ( 0,6 ), we have

Proof. See [23].
Lemma 2.2. In a harmonic Riemannian manafold, we have

> (a) $g^{\imath \jmath} g^{k l} g^{m n} \subseteq\left(A_{\imath j k l m n}\right)=144|\nabla R|^{2}$,
> (b) $g^{2 \jmath} g^{k l} g^{m n} \subseteq\left(B_{\imath j k l m n}\right)=48\left(\frac{S^{3}}{n^{2}}+\frac{9}{2 n} S|R|^{2}-\frac{7}{2} \hat{R}+\hat{R}\right)$,
> (c) $g^{\imath \jmath} g^{k l} g^{m n} \subseteq\left(g_{\imath j} g_{k l} g_{m n}\right)=48 n(n+2)(n+4)$,
where

$$
A_{i j k l m n}=\nabla_{k} R_{p l j}{ }^{q} \nabla_{l} R_{q m n}{ }^{p} \quad \text { and } \quad B_{\imath j k l m n}=R_{p i j}{ }^{q} R_{q k l}{ }^{r} R_{r m n}{ }^{p} .
$$

Proof. Putting $T_{\imath j k l m n}=g_{\imath \jmath} g_{k l} g_{m n}$ in Lemma 2.1, the formula (c) follows immediately. Next putting $T_{\imath \jmath k l m n}=A_{\imath \jmath k l m n}$ in Lemma 2.1, (1.1), (1.2) and (2.3) imply that

$$
A^{2 j k_{i j k}}=\frac{1}{2}|\nabla R|^{2}, \quad A^{2 j k}{ }_{\imath k_{j}}=\frac{1}{4}|\nabla R|^{2}, \quad A^{2 j k_{j i k}}=|\nabla R|^{2},
$$

$$
\begin{align*}
& +T^{\imath \jmath_{2}{ }_{2}{ }_{k j}+T^{\imath{ }_{j} i^{k}{ }_{k}}+T^{\imath \jmath_{j}{ }_{j}{ }_{2 k}}+T^{\imath \jmath}{ }_{j k}{ }^{k}{ }_{2}+T^{\imath j k}{ }_{\imath \jmath k} .}  \tag{2.7}\\
& \left.+T^{\imath j k}{ }_{\imath k j}+T^{\imath j k_{j i k}}+T^{\imath j k}{ }_{j k i}+T^{\imath j k_{k 2 j}}+T^{\imath j k_{k j 2}}\right) .
\end{align*}
$$

$$
A_{\jmath k \imath}^{\imath j k}=\frac{1}{2}|\nabla R|^{2}, \quad A_{k \imath \jmath}^{\imath j k}=\frac{1}{2}|\nabla R|^{2}, \quad A_{k \jmath \imath}^{\imath j k}=\frac{1}{2}|\nabla R|^{2}
$$

and all the others corresponding to the terms appearing in the right hand side of (2.7) vanish. Thus we obtain the formula (a). Lastly putting $T_{\imath j k l m n}=B_{\imath j k l m n}$ in Lemma 2.1, (1.1), (1.3) and (2.3) imply

$$
\begin{aligned}
& B_{i}^{\imath}{ }_{i}{ }_{j}{ }^{k}{ }_{k}=\frac{1}{n^{2}} S^{3}, \quad B_{i}^{\imath j k}{ }_{j k}=\frac{1}{2 n} S|R|^{2}, \quad B_{i}^{2}{ }_{i}{ }_{k j}=\frac{1}{n} S|R|^{2},
\end{aligned}
$$

$$
\begin{aligned}
& B^{\imath \jmath_{j i}{ }_{k}}=\frac{1}{n} S|R|^{2}, \quad B^{\imath \jmath_{j}{ }_{\imath k}}=-\frac{1}{2} \hat{R}, \quad B^{\imath \jmath}{ }_{\jmath}{ }_{k \imath}=-\hat{R}, \\
& B_{\imath \jmath k}^{\imath j k}=\stackrel{\circ}{R}-\frac{1}{4} \hat{R}, \quad B_{\imath{ }_{\imath j \jmath}}^{\imath j k}=-\frac{1}{4} \hat{R}, \quad B_{\jmath \jmath k}^{\imath j k}=-\frac{1}{4} \hat{R}, \\
& B_{j k \imath}^{\imath j k}=-\frac{1}{2} \hat{R}, \quad B_{k \imath \jmath}^{\imath j k}=-\frac{1}{2 n} S|R|^{2}, \quad B_{k \jmath j}^{\imath \jmath k}=\frac{1}{n} S|R|^{2} .
\end{aligned}
$$

Thus we obtain the formula (b).
Remark. Recently A. Gray and L. Vanhecke [7] has given in a Riemannian manifold many formulas which are useful in obtaining systematically scalar functions such as given in Lemma 2.2.

If we transvect $g^{2 \jmath} g^{k l} g^{m n}$ with (2.5), then taking account of the formulas (2.8), we have the following lemma (cf. Y. Watanabe [23]).

Lemma 2.3. In a harmonic Riemannian manifold of dimension $n$, we have

$$
\begin{equation*}
27|\nabla R|^{2}-32\left(\frac{S^{3}}{n^{2}}+\frac{9}{2 n} S|R|^{2}-\frac{7}{2} \hat{R}+\stackrel{\circ}{R}\right)=315 n(n+2)(n+4) \ddot{f}(0) \tag{2.9}
\end{equation*}
$$

## §3. 5-dimensional harmonic Riemannian manifolds.

We shall now prove a curvature identity (3.4) in a harmonic Riemannian manifold $M$ of dimension 5 . To do so, we introduce in a Riemannian manifold of dimension $n$ a function $G_{(m)}$ by

$$
\begin{equation*}
G_{(m)}=\delta_{\imath_{1} \imath_{2} \cdots \imath_{2 m}}^{\jmath_{1} \lambda_{2} \cdots \jmath_{2}} R^{\imath_{1} \imath_{2}}{ }_{\jmath_{1} \jmath_{2}} \cdots R^{\imath_{2 m-1} \imath_{2 m}}{ }_{\jmath_{2 m-1} \jmath_{2 m}} \tag{3.1}
\end{equation*}
$$

for any natural number $m \geqq 1$, where

$$
\delta_{\imath_{1} 2_{2} \cdots 2 m}^{\jmath_{1} \jmath_{2} \cdots \jmath_{2} m}=\left|\delta_{\imath_{b}}^{\jmath a}\right| \quad(a, b=1,2, \cdots, 2 m)
$$

is the so-called generalized Kronecker delta (cf. [15]). When $M$ is compact and $n=2 m$, the $G_{(m)}$ is, as is well known, the integrand of the Gauss-Bonnet
theorem (see, for example, S. Kobayashi and K. Nomizu [8]). However, the function $G_{(m)}$ vanishes identically for a Riemannian manifold $M$ of dimension $n<2 m$, because in such a case $\delta_{i_{1}}^{L_{1} 11_{2} \cdots i_{2 m} \cdots}$ in equal to zero. Therefore if the dimension of $M$ is 5 , then

$$
\begin{equation*}
G_{(3)}=0, \tag{3.2}
\end{equation*}
$$

which will be used in the sequel.
On the other hand, the function $G_{(3)}$ has in a Riemannian manifold of dimension $n$ the following form (cf. [17], [5]):

$$
\begin{align*}
G_{(3)}= & 8 \hat{R}-4 \hat{R}-24 R^{a b} R^{c d} R_{a b c d}-24 R^{u v} R_{u}^{a b c} R_{v a b c}  \tag{3.3}\\
& +16 R^{a b} R_{a}^{c} R_{b c}+S^{3}-12 S\left|R_{1}\right|^{2}+3 S|R|^{2},
\end{align*}
$$

as a consequence of (1.3). Thus using (3.3), we see that in an Einsteinian manifold of dimension 5 the formula (3.2) reduces to

$$
\begin{equation*}
4 \hat{R}-2 \hat{R}=\frac{S}{10}\left(9|R|^{2}-S^{2}\right) . \tag{3.4}
\end{equation*}
$$

The identity (3.4) is also obtained by transvecting $R^{h i j k}$ with the identity

$$
\begin{aligned}
& R_{h \imath p q} R^{p q}{ }_{j k}+2 R_{h p q k} R_{j}{ }^{p q}{ }_{i}-2 R_{h p q j} R_{k}{ }^{p q}{ }_{i}+\frac{3}{5} S R_{h \imath j k} \\
& \quad=\left(\frac{3}{20}|R|^{2}-\frac{1}{20} S^{3}\right)\left(g_{h j} g_{\imath k}-g_{h k} g_{\imath j}\right),
\end{aligned}
$$

proved by E. M. Patterson [15] in a harmonic Riemannian manifold of dimension 5.

From now on, let $M$ be a 5 -dimensional harmonic Riemannian manifold. We note here that any harmonic Riemannian manifold is necessarily Einsteinian (see $\S 2$ ). Eliminating $R$ from (1.5) and (3.4), we have

$$
\begin{equation*}
|\nabla R|^{2}+\frac{13}{10} S|R|^{2}+3 \hat{R}-\frac{S^{3}}{10}=0 \tag{3.5}
\end{equation*}
$$

Eliminating $\stackrel{R}{ }$ from (2.9) and (3.4), we have

$$
\begin{equation*}
27|\nabla R|^{2}+96 \hat{R}-36 S|R|^{2}-\frac{12}{25} S^{3}=315^{2} \ddot{f}(0) . \tag{3.6}
\end{equation*}
$$

Next eliminating $\hat{R}$ from (3.5) and (3.6), we have

$$
\begin{equation*}
5|\nabla R|^{2}+\frac{388}{5} S|R|^{2}-\frac{68}{25} S^{3}=-315^{2} \ddot{f}(0) \tag{3.7}
\end{equation*}
$$

Since $|R|^{2}$ and $S$ are constant in a harmonic Riemannian manifold because of (2.3) and (2.6), we have from (3.7), (3.5) and (3.4)

Proposition 3.1. In a 5-dimensional harmonic Riemannian manifold, the
scalar functions $|\nabla R|^{2}, \hat{R}$ and $\stackrel{R}{R}$ are all constant.
In a harmonic Riemannian manifold $M$ of dimension 5, we take a fixed point $x_{0}$ and a fixed unit vector $X$ belonging to the tangent space $T_{0} M$ to $M$ at $x_{0}$. Transvecting $X^{k} X^{i} X^{j} X^{l} X^{m} X^{n}$ with (2.5), we have at the point $x_{0}$

$$
9 \nabla_{k} R_{p i j}{ }^{q} \nabla_{l} R_{q m n}{ }^{p} X^{k} X^{i} X^{j} X^{l} X^{m} X^{n}=32 R_{p i j}{ }^{q} R_{q k l}{ }^{r} R_{r m n}{ }^{p} X^{k} X^{i} X^{j} X^{l} X^{m} X^{n}+315 \dddot{f}(0) .
$$

Substituting $\dddot{f}(0)$ given by (3.7) into the equation above, we have at the point $x_{0}$ the following key equation

$$
\begin{align*}
& 9 \nabla_{k} R_{p ı \jmath}{ }^{q} \nabla_{l} R_{q m n}{ }^{p} X^{k} X^{i} X^{j} X^{l} X^{m} X^{n}+\frac{1}{63}|\nabla R|^{2}  \tag{3.8}\\
& \quad=32 R_{p \imath j}^{q} R_{q k l} R_{r m n}^{p} X^{k} X^{i} X^{j} X^{l} X^{m} X^{n}+\frac{1}{315}\left(-\frac{388}{5} S|R|^{2}+\frac{68}{25} S^{3}\right) .
\end{align*}
$$

We now define a linear transformation $\Pi_{X}$ in the tangent space $T_{0} M$ by

$$
\begin{equation*}
\Pi_{X}(Y)=-R(X, Y) X \tag{3.9}
\end{equation*}
$$

for $Y \in T_{0} M$, which implies immediately

$$
\begin{equation*}
\Pi_{X}(X)=0 . \tag{3.10}
\end{equation*}
$$

For simplicity, we put $\Pi_{X}=\Pi$. Then $X$ is obviously an eigen vector of $\Pi$ with eigen value 0 , because of (3.10). Since the linear transformation $\Pi$ is symmetric because of (3.9), there is an orthonormal basis $\left\{X, e_{1}, e_{2}, e_{3}, e_{4}\right\}$ such that

$$
\begin{equation*}
\Pi\left(e_{\alpha}\right)=\lambda_{\alpha} e_{\alpha} \quad(\alpha=1,2,3,4) \tag{3.11}
\end{equation*}
$$

Then by (2.3) and (2.4) we get

$$
\begin{align*}
& \operatorname{Tr}(\Pi)=R_{j i} X^{j} X^{i}=\frac{S}{5}=-\frac{3}{2} \dot{f}(0),  \tag{3.12}\\
& \operatorname{Tr}\left(\Pi^{2}\right)=R_{p ı}{ }^{q} R_{q k l^{p}} X^{i} X^{j} X^{k} X^{l}=-\frac{45}{8} \tilde{f}(0), \\
& \operatorname{Tr}\left(\Pi^{3}\right)=R_{p ı \jmath}^{q} R_{q k l}{ }^{r} R_{r m n}{ }^{p} X^{i} X^{j} X^{k} X^{l} X^{m} X^{n},
\end{align*}
$$

where $\operatorname{Tr}$ means the trace of each of linear transformations $\Pi, \Pi^{2}=\Pi \circ \Pi$ and $\Pi^{3}=\Pi \circ \Pi \circ \Pi$. Thus (3.8) implies the following equation (3.13) at the point $x_{0}$ because of (2.6) and (3.11).

$$
\begin{align*}
& 9 \nabla_{k} R_{p 2 \rho}{ }^{q} \nabla_{l} R_{q m n}{ }^{p} X^{k} X^{i} X^{j} X^{l} X^{m} X^{n}+\frac{1}{63}|\nabla R|^{2}  \tag{3.13}\\
& \quad=32 \operatorname{Tr}\left(\Pi^{3}\right)+\frac{1}{315}\left\{-388 \operatorname{Tr}(\Pi)\left(-\frac{15}{2} \dot{f}(0)^{2}-\frac{525}{4} \dot{f}(0)\right)+340(\operatorname{Tr}(\Pi))^{3}\right\}
\end{align*}
$$

Then arranging (3.13) and using (3.12), we get at the point $x_{0}$

$$
\begin{equation*}
9 \nabla_{k} R_{p \imath \jmath}{ }^{q} \nabla_{l} R_{q m n}^{p} X^{k} X^{i} X^{j} X^{l} X^{m} X^{n}+\frac{1}{63}|\nabla R|^{2}={ }_{27}^{4} F(\lambda), \tag{3.14}
\end{equation*}
$$

where

$$
F(\lambda)=216 \operatorname{Tr}\left(\Pi^{3}\right)+35(\operatorname{Tr}(\Pi))^{3}-194 \operatorname{Tr}(\Pi) \operatorname{Tr}\left(\Pi^{2}\right) .
$$

As a consequence of (3.11), we have

$$
\begin{aligned}
F(\lambda) & =216 \sum_{\alpha} \lambda_{\alpha}^{3}+35\left(\sum_{\alpha} \lambda_{\alpha}\right)^{3}-194\left(\sum_{\alpha} \lambda_{\alpha}\right) \sum_{\beta} \lambda_{\beta}^{2} \\
& =54\left\{4 \sum_{\alpha} \lambda_{\alpha}^{3}-\left(\sum_{\alpha} \lambda_{\alpha}\right) \sum_{\beta} \lambda_{\beta}^{2}\right\}+35 \sum_{\alpha} \lambda_{a}\left\{\left(\sum_{\beta} \lambda_{\beta}\right)^{2}-4 \sum_{\beta} \lambda_{\beta}^{\gamma}\right\} \\
& =54 \sum_{\beta<r}\left(\lambda_{\beta}-\lambda_{\gamma}\right)^{2}\left(\lambda_{\beta}+\lambda_{\gamma}\right)+35\left(\sum_{\alpha} \lambda_{\alpha}\right) \sum_{\beta<r}\left(\lambda_{\beta}-\lambda_{i}\right)^{2} \\
& =\sum_{\beta<r}\left(\lambda_{\beta}-\lambda_{\gamma}\right)^{2}\left\{54\left(\lambda_{\beta}+\lambda_{i}\right)-35 \sum_{\alpha} \lambda_{\alpha}\right\},
\end{aligned}
$$

which implies the following formulas

$$
\begin{align*}
F(\lambda)= & \left(\lambda_{1}-\lambda_{2}\right)^{2}\left\{19\left(\lambda_{1}+\lambda_{2}\right)-35\left(\lambda_{3}+\lambda_{4}\right)\right\}+\left(\lambda_{1}-\lambda_{3}\right)^{2}\left\{19\left(\lambda_{1}+\lambda_{3}\right)-35\left(\lambda_{2}+\lambda_{4}\right)\right\}  \tag{3.15}\\
& +\left(\lambda_{1}-\lambda_{4}\right)^{2}\left\{19\left(\lambda_{1}+\lambda_{4}\right)-35\left(\lambda_{2}+\lambda_{3}\right)\right\}+\left(\lambda_{2}-\lambda_{3}\right)^{2}\left\{19\left(\lambda_{2}+\lambda_{3}\right)-35\left(\lambda_{1}+\lambda_{1}\right)\right\} \\
& +\left(\lambda_{2}-\lambda_{4}\right)^{2}\left\{19\left(\lambda_{2}+\lambda_{4}\right)-35\left(\lambda_{1}+\lambda_{3}\right)\right\}+\left(\lambda_{3}-\lambda_{4}\right)^{2}\left\{19\left(\lambda_{3}+\lambda_{4}\right)-35\left(\lambda_{1}+\lambda_{2}\right)\right\}
\end{align*}
$$

and

$$
\begin{align*}
F(\lambda)= & \left(\lambda_{1}-\lambda_{2}\right)^{2}\left\{19 \sum_{\alpha} \lambda_{\alpha}-54\left(\lambda_{3}+\lambda_{4}\right)\right\}+\left(\lambda_{1}-\lambda_{3}\right)^{2}\left\{19 \sum_{\alpha} \lambda_{\alpha}-54\left(\lambda_{2}+\lambda_{4}\right)\right\}  \tag{3.16}\\
& +\left(\lambda_{1}-\lambda_{4}\right)^{2}\left\{19 \sum_{\alpha} \lambda_{\alpha}-54\left(\lambda_{2}+\lambda_{3}\right)\right\}+\left(\lambda_{2}-\lambda_{3}\right)^{2}\left\{19 \sum_{\alpha} \lambda_{\alpha}-54\left(\lambda_{1}+\lambda_{4}\right)\right\} \\
& +\left(\lambda_{2}-\lambda_{4}\right)^{2}\left\{19 \sum_{\alpha} \lambda_{\alpha}-54\left(\lambda_{1}+\lambda_{3}\right)\right\}+\left(\lambda_{3}-\lambda_{4}\right)^{2}\left\{19 \sum_{\alpha} \lambda_{\alpha}-54\left(\lambda_{1}+\lambda_{2}\right)\right\} .
\end{align*}
$$

As a consequence of (3.14), we have the following inequality

$$
\begin{equation*}
F(\lambda) \geqq 0 . \tag{3.17}
\end{equation*}
$$

We first note that $\lambda_{\alpha}$ is the sectional curvature for the orthonormal pair $\left\{X, e_{\alpha}\right\} \quad(\alpha=1,2,3,4)$, because of the definition (3.9) of $\Pi=\Pi_{X}$. Suppose that all $\lambda_{\alpha}$ satisfy $\delta \geqq \lambda_{\alpha} \geqq(19 / 35) \delta$ for some $\delta \geqq 0$. Then we see from (3.15) that the right hand side of (3.14) is non-positive. Consequently it follows from (3.14) and (3.17) that $|\nabla R|^{2}=0$ at the point $x_{0}$ of $M$. Since $M$ is connected, Proposition 3.1 implies that $\nabla R=0$, i.e. that $M$ is locally symmetric. Thus, we have the following

Lemma 3.2. Let $M$ be a 5 -dimensional harmonic Rremanman mamifold all of whose sectıonal curvatures $\kappa_{x}(X, Y)$ at a pornt $x$ satisfy $\delta \geqq \kappa_{x}(X, Y) \geqq(19 / 35) \delta$ for some $\delta \geqq 0$. Then $M$ is locally symmetric.

Since a locally symmetric harmonic manifold is locally flat or locally isometric to a rank one symmetric space (cf. [12], [6]), it is of constant curvature if it is odd-dimensional. Thus, Lemma 3.2 implies

Theorem 3.3. Let $M$ be a 5-dimensional harmonic Riemannian manafold all of whose sectıonal curvatures $\kappa_{x}(X, Y)$ at a point $x$ satısfy $\delta \geqq \kappa_{x}(X, Y) \geqq(19 / 35) \delta$ for some $\delta \geqq 0$. Then $M$ is of constant curvature.

Similarly using (3.16) and noting $S=5 \sum_{\alpha} \lambda_{\alpha}$, we obtain
Theorem 3.4. Let $M$ be a 5-dimensional harmonic Riemannian mannfold. If all sectional curvatures $\kappa_{x}(X, Y)$ at a point $x$ satzsfy $\kappa_{x}(X, Y) \geqq(19 / 540) S$, where $S$ is the scalar curvature, then $M$ is of constant curvature.

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