Research Article

# Geometric Integrability of Some Generalizations of the Camassa-Holm Equation 

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Received 27 May 2011; Accepted 17 July 2011
Academic Editor: V. A. Yurko
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We study the Camassa-Holm (CH) equation and recently introduced $\mu \mathrm{CH}$ equation from the geometric point of view. We show that Kupershmidt deformations of these equations describe pseudospherical surfaces and hence are geometrically integrable.

## 1. Introduction

The modern theory of integrable nonlinear partial differential equations arose as a result of the inverse scattering method (ISM) discovered by Gardner et al. [1] for Korteweg de Vries (KdV) equation. Soon after, it was realized that this method can be applied to several important nonlinear equations like nonlinear Shrödinger equation, sine-Gordon, and so forth.

Sasaki [2] gave a natural geometric interpretation for ISM in terms of pseudospherical surfaces. Motivated by Sasaki, Chern and Tenenblat [3] introduce the notion of a scalar equation of pseudospherical type and study systematically the evolution equations that describe pseudospherical surfaces. It appears that almost all important equations and systems in mathematical physics enjoy this property [3-7]. The advantage of this geometric treatment is that most of the ingredients connected with the integrable equations such as Lax pair, zero curvature representation, conservation laws, and symmetries come naturally.

Let us recall some facts about the equations we study in this paper.
The Camassa-Holm equation (CH)

$$
\begin{equation*}
u_{t}-u_{x x t}+2 \omega u_{x}+3 u u_{x}-2 u_{x} u_{x x}-u u_{x x x}=0 \tag{1.1}
\end{equation*}
$$

where $\omega \in \mathbb{R}$, has appeared in [8] as an equation with bi-Hamiltonian structure. Later, in [9], it was considered as a model, describing the unidirectional propagation of shallow water
waves over a flat bottom. CH is a completely integrable equation; see, for example, [10-12]. Furthermore, CH is geometrically integrable [6]. Here, we consider the case $\omega=0$. Then, (1.1) is

$$
\begin{equation*}
m_{t}+u m_{x}+2 m u_{x}, \quad m=u_{x x}-u \tag{1.2}
\end{equation*}
$$

The bi-Hamiltonian form of (1.2) is [8, 9]

$$
\begin{equation*}
m_{t}=-\mathbb{B}^{1} \frac{\delta H_{2}[m]}{\delta m}=-\mathbb{B}^{2} \frac{\delta H_{1}[m]}{\delta m} \tag{1.3}
\end{equation*}
$$

where $B^{1}=\partial-\partial^{3}, B^{2}=m \partial+\partial m$ are the two compatible Hamiltonian operators ( $\partial$ stands for $\partial / \partial x)$, and the corresponding Hamiltonians are

$$
\begin{equation*}
H_{1}[m]=\frac{1}{2} \int m u d x, \quad H_{2}[m]=\frac{1}{2} \int\left(u^{3}+u u_{x}^{2}\right) d x \tag{1.4}
\end{equation*}
$$

There exists an infinite sequence of conservation laws (multi-Hamiltonian structure) $H_{n}[m], n=0, \pm 1, \pm 2, \ldots$ including (1.4) such that

$$
\begin{equation*}
B^{1} \frac{\delta H_{n}[m]}{\delta m}=B^{2} \frac{\delta H_{n-1}[m]}{\delta m} \tag{1.5}
\end{equation*}
$$

Very recently a "modified" CH equation (mCH) was introduced by analogy of "Miura transform" from the theory of KdV equation in [5].

The $\mu \mathrm{CH}$ equation was derived recently in $[13,14]$ as

$$
\begin{equation*}
-u_{x x t}=-2 \mu(u) u_{x}+2 u_{x} u_{x x}+u u_{x x x} \tag{1.6}
\end{equation*}
$$

where $\mu(u)=\int_{0}^{1} u d x$ and $u(t, x)$ is a spatially periodic real-valued function of time variable $t$ and space variable $x \in S^{1}=[0,1)$. In order to keep a certain symmetry and analogy with CH , one can write the above equation in the form (see also [15])

$$
\begin{equation*}
\mu\left(u_{t}\right)-u_{x x t}=-2 \mu(u) u_{x}+2 u_{x} u_{x x}+u u_{x x x} . \tag{1.7}
\end{equation*}
$$

Note that $\mu\left(u_{t}\right)=0$ in the periodic case. Introducing $m=A u=\mu(u)-u_{x x}$ (1.7) becomes

$$
\begin{equation*}
m_{t}+u m_{x}+2 m u_{x}=0, \quad m=\mu(u)-u_{x x} \tag{1.8}
\end{equation*}
$$

The $\mu \mathrm{CH}$ is an integrable equation and arises as an asymptotic rotator equation in a liquid crystal with a preferred direction if one takes into account the reciprocal action of dipoles on themselves $[13,14]$.

The bi-Hamiltonian form of (1.8) is

$$
\begin{equation*}
m_{t}=-\beta^{1} \frac{\delta H_{2}[m]}{\delta m}=-\bar{B}^{2} \frac{\delta H_{1}[m]}{\delta m} \tag{1.9}
\end{equation*}
$$

where $B^{1}=\partial A=-\partial^{3}, B^{2}=m \partial+\partial m$ are the two compatible Hamiltonian operators, and the corresponding Hamiltonians are

$$
\begin{equation*}
H_{1}[m]=\frac{1}{2} \int m u d x, \quad H_{2}[m]=\int\left(\mu(u) u^{2}+\frac{1}{2} u u_{x}^{2}\right) d x . \tag{1.10}
\end{equation*}
$$

Also the $\mu \mathrm{CH}$ equation is geometrically integrable [15] (see there for other geometric descriptions of this equation).

Recently in [16], a new 6th-order wave equation, named KdV6, was derived. After some rescaling, this equation can be presented as the following system;

$$
\begin{gather*}
u_{t}=6 u u_{x}+u_{x x x}-w_{x}  \tag{1.11}\\
w_{x x x}+4 u w_{x}+2 u_{x} w=0 .
\end{gather*}
$$

This system gives a perturbation to KdV equation $(w=0)$, and since the constrain on $w$ is differential, this is a nonholonomic deformation.

Kupershmidt [17] suggested a general construction applicable to any bi-Hamiltonian systems providing a nonholonomic perturbation on it. This perturbation is conjectured to preserve integrability. In the case of KdV6, the system (1.11) can be converted into

$$
\begin{gather*}
u_{t}=B^{1} \frac{\delta H_{n+1}}{\delta u}-B^{1}(w)=B^{2} \frac{\delta H_{n}}{\delta u}-B^{1}(w),  \tag{1.12}\\
B^{2}(w)=0
\end{gather*}
$$

where $B^{1}=\partial, B^{2}=\partial^{3}+2(m \partial+\partial m)$ are the two standard Hamiltonian operators of the KdV hierarchy and

$$
\begin{equation*}
H_{1}=\int u d x, \quad H_{2}=\frac{1}{2} \int u^{2} d x, \ldots \tag{1.13}
\end{equation*}
$$

In the same paper, Kupershmidt verifies integrability of KdV6, as well as integrability of such nonholonomic deformations for some representative cases: the classical long-wave equation, the Toda lattice (both continuous and discrete), and the Euler top.

In fact, Kersten et al. [18] prove that the Kupershmidt deformation of every biHamiltonian equation is again bi-Hamiltonian system and every hierarchy of conservation laws of the original bi-Hamiltonian system, gives rise to a hierarchy of conservation laws of the Kupershmidt deformation.

The aim of this paper is to show that the CH and the $\mu \mathrm{CH}$ equations have geometrically integrable Kupershmidt deformations. We also show that the KdV6 equation and twocomponent CH system [19] are also geometrically integrable.

As a matter of fact, Yao and Zeng [20] propose a generalized Kupershmidt deformation and verify that this generalized deformation also preserves integrability in few representative cases: KdV equation, Boussinesq equation, Jaulent-Miodek equation, and CH equation.

Kundu et al. [21] consider slightly generalized form of deformation for the KdV equation and extend this approach to $m K d V$ equation and to AKNS system.

Guha [22] uses Kirillov's theory of coadjoint representation of Virasoro algebra to obtain a large class of KdV6-type equations, equivalent to the original one. Also, applying the Adler-Konstant-Symes scheme, he constructs a new nonholonomic deformation of the coupled KdV equation.

The paper is organized as follows. In Section 2, we recall some facts about equations that describe pseudospherical surfaces. The main results are in Section 3. There we show that the Kupershmidt deformations for the CH equation and for the $\mu \mathrm{CH}$ equation are of pseudospherical type and hence are geometrically integrable. We also derive the corresponding quadratic pseudopotentials which turn out to be very useful in obtaining conservation laws and symmetries. At the end, we make some speculations about "modified" $\mu \mathrm{CH}$ equation.

## 2. Equations of Pseudospherical Type and Pseudopotentials

In this section, we recall some definitions and facts. One can consult, for example, [3-7] for more details.

Definition 2.1. A scalar differential equation $\Xi\left(x, t, u, u_{x}, \ldots, u_{x^{n} t^{m}}\right)=0$ in two independent variables $x, t$ is of pseudospherical type (or, it describes pseudospherical surfaces) if there exist one-forms $\omega^{\alpha} \neq 0$

$$
\begin{equation*}
\omega^{\alpha}=f_{\alpha 1}\left(x, t, u, \ldots, u_{x^{r} t p}\right) d x+f_{\alpha 2}\left(x, t, u, \ldots, u_{x^{s} t q}\right) d t, \quad \alpha=1,2,3 \tag{2.1}
\end{equation*}
$$

whose coefficients $f_{\alpha \beta}$ are smooth functions which depend on $x, t$ and finite number of derivatives of $u$, such that the 1 -forms $\bar{\omega}^{\alpha}=\omega^{\alpha}(u(x, t))$ satisfy the structure equations

$$
\begin{equation*}
d \bar{\omega}^{1}=\bar{\omega}^{3} \wedge \bar{\omega}^{2}, \quad d \bar{\omega}^{2}=\bar{\omega}^{1} \wedge \bar{\omega}^{3}, \quad d \bar{\omega}^{3}=\bar{\omega}^{1} \wedge \bar{\omega}^{2} \tag{2.2}
\end{equation*}
$$

whenever $u=u(x, t)$ is a solution of $\Xi=0$.
Equations (2.2) can be interpreted as follows. The graphs of local solutions of equations of pseudospherical type can be equipped with structure of pseudospherical surface (see $[3,6,7]$ ): if $\bar{\omega}^{1} \wedge \bar{\omega}^{2} \neq 0$, the tensor $\bar{\omega}^{1} \otimes \bar{\omega}^{1}+\bar{\omega}^{2} \otimes \bar{\omega}^{2}$ defines a Riemannian metric of constant Gaussian curvature -1 on the graph of solution $u(x, t)$, and $\bar{\omega}^{3}$ is the corresponding metric connection one-form.

An equation of pseudospherical type is the integrability condition for a $\operatorname{sl}(2, \mathbb{R})$-valued problem

$$
\begin{equation*}
d \psi=\Omega \psi, \tag{2.3}
\end{equation*}
$$

where $\Omega$ is the matrix-valued one-form

$$
\Omega=X d x+T d t=\frac{1}{2}\left(\begin{array}{cc}
\omega^{2} & \omega^{1}-\omega^{3}  \tag{2.4}\\
\omega^{1}+\omega^{3} & -\omega^{2}
\end{array}\right)
$$

Definition 2.2. An equation $\Xi=0$ is geometrically integrable if it describes a nontrivial oneparameter family of pseudospherical surfaces.

Hence, if $\Xi=0$ is geometrically integrable, it is the integrability condition of oneparameter family of linear problems $\psi_{x}=X \psi, \psi_{t}=T \psi$. In fact, this is equivalent to the zero curvature equation

$$
\begin{equation*}
X_{t}-T_{x}+[X, T]=0 \tag{2.5}
\end{equation*}
$$

which is an essential ingredient of integrable equations.
Another important property of equations of pseudospherical type is that they admit quadratic pseudopotentials. Pseudopotentials are a generalization of conservation laws.

Proposition 2.3 (see [6]). Let $\Xi=0$ be a differential equation describing pseudospherical surfaces with associated one-forms $\omega^{\alpha}$. The following two Pfaffian systems are completely integrable whenever $u(x, t)$ is a solution of $\Xi=0$ :

$$
\begin{align*}
& -2 d \Gamma=\omega^{3}+\omega^{2}-2 \Gamma \omega^{1}+\Gamma^{2}\left(\omega^{3}-\omega^{2}\right)  \tag{2.6}\\
& 2 d \gamma=\omega^{3}-\omega^{2}-2 \gamma \omega^{1}+\gamma^{2}\left(\omega^{3}+\omega^{2}\right) \tag{2.7}
\end{align*}
$$

Moreover, the one-forms

$$
\begin{align*}
& \Theta=\omega^{1}-\Gamma\left(\omega^{3}-\omega^{2}\right) \\
& \widehat{\Theta}=-\omega^{1}+\gamma\left(\omega^{3}+\omega^{2}\right) \tag{2.8}
\end{align*}
$$

are closed whenever $u(x, t)$ is a solution of $\Xi=0$ and $\Gamma$ (resp. $\gamma$ ) is a solution of (2.6) (resp. (2.7)).
Geometrically, Pfaffian systems (2.6) and (2.7) determine geodesic coordinates on the pseudospherical surfaces associated with the equation $\Xi=0[3,6]$.

## 3. Results

In this section, we consider the nonholonomic deformation of CH equation and $\mu \mathrm{CH}$ equation. We show that they are geometrically integrable and consider their quadratic pseudopotentials. The nonlocal symmetries will be studied elsewhere.

### 3.1. CH Equation

Recall from Introduction the CH equation (1.2) and its bi-Hamiltonian form (1.3) with the corresponding Hamiltonian operators $B^{1}$ and $B^{2}$.

Following Kupershmidt's construction, we introduce the nonholonomic deformation of the CH equation

$$
\begin{gather*}
m_{t}=-u m_{x}-2 m u_{x}-B^{1}(w)  \tag{3.1}\\
B^{2}(w)=0
\end{gather*}
$$

Proposition 3.1. The system (3.1) describes pseudospherical surface and hence is geometrically integrable.

Let us give the corresponding 1 -forms

$$
\begin{align*}
\omega^{1}= & (m+1) d x+\left[-u m+\frac{\eta+1}{\eta}\left(u_{x}-u\right)+\frac{1}{\eta}-w_{x x}+(\eta+1) w_{x}-\eta w(m+1)\right] d t \\
\omega^{2}= & (\eta+1) d x+\left[\frac{\eta+1}{\eta}-(\eta+1) u+u_{x}-\eta(\eta+1) w+\eta w_{x}\right] d t  \tag{3.2}\\
\omega^{3}= & (m+1+\eta) d x \\
& +\left[-u m+\frac{\eta-(\eta+1)^{2}}{\eta} u+\frac{\eta+1}{\eta} u_{x}+\frac{\eta+1}{\eta}-w_{x x}+(\eta+1) w_{x}-\eta w(m+1+\eta)\right] d t
\end{align*}
$$

For the proof of Proposition 3.1, we need only to verify the structure equations (2.2) and to check that the parameter $\eta$ is intrinsic.

For the matrices $X$ and $T$ in, we get

$$
X=\frac{1}{2}\left(\begin{array}{cc}
\eta+1 & -\eta  \tag{3.3}\\
2(m+1)+\eta & -(\eta+1)
\end{array}\right), \quad T=\frac{1}{2}\left(\begin{array}{cc}
T_{11} & T_{12} \\
T_{21} & -T_{11}
\end{array}\right)
$$

where

$$
\begin{align*}
T_{11}= & \frac{\eta+1}{\eta}-(\eta+1) u+u_{x}-\eta(\eta+1) w+\eta w_{x} \\
T_{12}= & \eta u+\eta^{2} w-1 \\
T_{21}= & -2 u m-\frac{(\eta+1)^{2}+1}{\eta} u+\frac{2(\eta+1)}{\eta} u_{x}  \tag{3.4}\\
& +\frac{\eta+2}{\eta}-2 w_{x x}+2(\eta+1) w_{x}-\eta(\eta+2) w-2 \eta w m
\end{align*}
$$

Hence, we have a zero curvature representation $X_{t}-T_{x}+[X, T]=0$ for the system (3.1). From (3.3), it is straightforward to obtain the corresponding scalar linear problem

$$
\begin{gather*}
\psi_{x x}=\left(\frac{1}{4}-\frac{\eta}{2} m\right) \psi  \tag{3.5}\\
\psi_{t}=\left(-u-\eta w+\frac{1}{\eta}\right) \psi_{x}+\frac{u_{x}+\eta w_{x}}{2} \psi .
\end{gather*}
$$

In order to apply Proposition 2.3 to nonholonomic deformation of the CH equation, we consider new 1-forms $\omega_{\text {new }}^{\alpha}$

$$
\begin{equation*}
\omega_{\text {new }}^{1}=\omega^{2}, \quad \omega_{\text {new }}^{2}=-\omega^{1}, \quad \omega_{\text {new }}^{3}=\omega^{3} . \tag{3.6}
\end{equation*}
$$

With these forms the Pfaffian system (2.7) becomes

$$
\begin{align*}
2 \gamma_{x}= & \eta \gamma^{2}-2(\eta+1) \gamma+2(m+1)+\eta \\
2 \gamma_{t}= & r^{2}-2 w_{x x}-2[\eta \gamma-(\eta+1)] w_{x}-2 w \eta r_{x}  \tag{3.7}\\
& -2\left[\gamma-\frac{\eta+1}{\eta}\right] u_{x}-2 u\left(r_{x}+\frac{1}{\eta}\right)+\frac{\eta+2}{\eta}-2 \gamma \frac{\eta+1}{\eta}
\end{align*}
$$

Applying the transform

$$
\begin{equation*}
\gamma \longmapsto r+\frac{\eta+1}{\eta} \tag{3.8}
\end{equation*}
$$

after some algebraic manipulations and setting $\lambda=-1 / \eta$, we obtain the following result.
Proposition 3.2. The nonholonomic deformation of the CH equation (3.1) admits a quadratic pseudopotential $\gamma$, defined by the equations

$$
\begin{align*}
& m=\frac{r^{2}}{2 \lambda}+r_{x}-\frac{\lambda}{2}  \tag{3.9}\\
& r_{t}=\frac{r^{2}}{2}\left[1+\frac{1}{\lambda}\left(u-\frac{w}{\lambda}\right)\right]-r\left(u-\frac{w}{\lambda}\right)_{x}-\left(u-\frac{w}{\lambda}\right)\left(m+\frac{\lambda}{2}\right)+\lambda u-\frac{\lambda^{2}}{2} \tag{3.10}
\end{align*}
$$

where $\lambda \neq 0, m=u_{x x}-u$. Moreover, (3.1) possesses the parameter-dependent conservation law

$$
\begin{equation*}
r_{t}=\lambda\left[(u+w)_{x}-\gamma-\frac{1}{\lambda}\left(u-\frac{w}{\lambda}\right) \gamma\right]_{x} . \tag{3.11}
\end{equation*}
$$

Conservation densities can be obtained by expanding (3.9) and (3.11) in powers of $\lambda$. Note that the left hand side of (3.11) and (3.9) does not depend on $w$ as it should be. The corresponding expansions are performed in [6].

## 3.2. $\mu \mathrm{CH}$ Equation

Consider now the $\mu \mathrm{CH}$ equation (1.8), its bi-Hamiltonian form (1.9) with the corresponding Hamiltonian operators $B^{1}, B^{2}$. Applying Kupershmidt's procedure to (1.8), we obtain the nonholonomic deformation of the $\mu \mathrm{CH}$ equation

$$
\begin{gather*}
m_{t}=-u m_{x}-2 m u_{x}-B^{1}(w) \\
B^{2}(w)=0 \tag{3.12}
\end{gather*}
$$

or

$$
\begin{gather*}
m_{t}=-u m_{x}-2 m u_{x}+w_{x x x} \\
2 m w_{x}+w m_{x}=0, \quad m=\mu(u)-u_{x x} \tag{3.13}
\end{gather*}
$$

Proposition 3.3. The nonholonomic deformation of the $\mu \mathrm{CH}$ equation (3.13) describes pseudospherical surfaces and, hence, is geometrically integrable.

For validation of Proposition 3.3, we give the 1-forms associated with (3.13)

$$
\begin{align*}
\omega^{1}= & \frac{1}{2}\left(\eta m-\frac{\eta^{2}}{2}+2\right) d x \\
& +\frac{1}{2}\left[\frac{\eta^{2}}{2} u-\eta\left(u_{x}+u m+\frac{1}{2}\right)+\mu(u)-2 u+\frac{2}{\eta}+\left(\frac{\eta^{3}}{2}-2 \eta\right) w-\eta^{2} w_{x}+\eta w_{x x}-\eta^{2} m w\right] d t \\
\omega^{2}= & \eta d x+\left(1-\eta u+u_{x}-\eta^{2} w+\eta w_{x}\right) d t \\
\omega^{3}= & \frac{1}{2}\left(\eta m-\frac{\eta^{2}}{2}-2\right) d x \\
& +\frac{1}{2}\left[\frac{\eta^{2}}{2} u-\eta\left(u_{x}+u m+\frac{1}{2}\right)+\mu(u)+2 u-\frac{2}{\eta}+\left(\frac{\eta^{3}}{2}+2 \eta\right) w-\eta^{2} w_{x}+\eta w_{x x}-\eta^{2} m w\right] d t \tag{3.14}
\end{align*}
$$

For the matrices $X$ and $T$ in, we get

$$
X=\frac{1}{2}\left(\begin{array}{cc}
\eta & 2  \tag{3.15}\\
\eta m-\frac{\eta^{2}}{2} & -\eta
\end{array}\right), \quad T=\frac{1}{2}\left(\begin{array}{cc}
T_{11} & T_{12} \\
T_{21} & -T_{11}
\end{array}\right)
$$

where

$$
\begin{align*}
& T_{11}=1-\eta u+u_{x}-\eta^{2} w+\eta w_{x} \\
& T_{12}=2\left(-u+\frac{1}{\eta}-\eta w\right)  \tag{3.16}\\
& T_{21}=\frac{\eta^{2}}{2} u-\eta\left(u_{x}+u m+\frac{1}{2}\right)+\mu(u)-\eta^{2} w_{x}+\eta w_{x x}-\eta^{2} m w+\frac{\eta^{3}}{2} w
\end{align*}
$$

Hence, we have a zero curvature representation $X_{t}-T_{x}+[X, T]=0$ for the system (3.13). From (3.3), it is straightforward to obtain the corresponding scalar linear problem

$$
\begin{gather*}
\psi_{x x}=\left(\frac{\eta}{2} m\right) \psi \\
\psi_{t}=\left(-u-\eta w+\frac{1}{\eta}\right) \psi_{x}+\frac{u_{x}+\eta w_{x}}{2} \psi \tag{3.17}
\end{gather*}
$$

which coincides with those in [13] upon setting $w=0$ and $\lambda=\eta / 2$.
In order to find pseudopotentials for the nonholonomic deformation of the $\mu \mathrm{CH}$ equation, we proceed as before denoting

$$
\begin{equation*}
\omega_{\text {new }}^{1}=\omega^{2}, \quad \omega_{\text {new }}^{2}=-\omega^{1}, \quad \omega_{\text {new }}^{3}=\omega^{3} . \tag{3.18}
\end{equation*}
$$

With these forms, the Pfaffian system (2.7) becomes

$$
\begin{align*}
2 \gamma_{x}= & -2 \gamma^{2}-2 \eta \gamma+\eta m-\frac{\eta^{2}}{2}  \tag{3.19}\\
2 \gamma_{t}= & -\frac{2 \gamma^{2}}{\eta}+2 \gamma^{2}(u+\eta w)-2 \gamma\left(1-\eta u+u_{x}-\eta^{2} w+\eta w_{x}\right) \\
& +\left[\frac{\eta^{2}}{2} u-\eta\left(u_{x}+m+\frac{1}{2}\right)+\mu(u)+\frac{\eta^{3}}{2} w-\eta^{2} w_{x}+\eta w_{x x}-\eta^{2} m w\right] \tag{3.20}
\end{align*}
$$

After some manipulations, the above system obtains the form

$$
\begin{gather*}
2 \gamma_{x}=-2 \gamma^{2}-2 \eta \gamma+\eta m-\frac{\eta^{2}}{2} \\
2 \gamma_{t}=-\frac{2}{\eta} \gamma^{2}+\eta w_{x x}-[(2 \gamma+\eta)(u+\eta w)]_{x}+\mu(u)-2 \gamma-\frac{\eta}{2} \tag{3.21}
\end{gather*}
$$

Applying the transform $\gamma \mapsto \gamma-\eta / 2$, we get

$$
\begin{gather*}
\gamma_{x}=-\gamma^{2}+\frac{\eta}{2} m  \tag{3.22}\\
\gamma_{t}=-\frac{\gamma^{2}}{\eta}+\frac{\eta}{2} w_{x x}-[\gamma(u+\eta w)]_{x}+\frac{\mu(u)}{2} \tag{3.23}
\end{gather*}
$$

Multiplying the first equation (3.22) by $-1 / \eta$ and then adding the result to the second equation (3.23), we get the following result denoting $\lambda=\eta / 2$.

Proposition 3.4. The nonholonomic deformation of the $\mu \mathrm{CH}$ equation (3.13) admits a quadratic pseudopotential $\gamma$, defined by the equations

$$
\begin{gather*}
m=\frac{\gamma^{2}}{\lambda}+\frac{\gamma_{x}}{\lambda}  \tag{3.24}\\
\gamma_{t}=-\frac{2 \gamma^{2}}{\lambda} \lambda w_{x x}-[\gamma(u+2 \lambda w)]_{x}+\frac{\mu(u)}{2} \tag{3.25}
\end{gather*}
$$

where $\lambda \neq 0, m=\mu(u)-u_{x x}$. Moreover, (3.13) possesses the parameter-dependent conservation law

$$
\begin{equation*}
\gamma_{t}=\frac{1}{2 \lambda}\left[\gamma+\lambda(u+2 \lambda w)_{x}-2 \lambda(u+2 \lambda w) \gamma\right]_{x} \tag{3.26}
\end{equation*}
$$

As the conserved densities for the nonholonomic deformation are the same as for the original bi-Hamiltonian system, we make use of the pseudopotentials to obtain them for the $\mu \mathrm{CH}$ equation. One possible expansion of $\gamma$ is

$$
\begin{equation*}
\gamma=\lambda^{1 / 2} \gamma_{1}+\gamma_{0}+\sum_{j=1}^{\infty} \lambda^{-j / 2} \gamma_{-j} \tag{3.27}
\end{equation*}
$$

Substituting this into (3.24) yields

$$
\begin{equation*}
\gamma_{1}=\sqrt{m}, \quad \gamma_{0}=-\frac{m_{x}}{4 m}, \quad \gamma_{-1}=\frac{1}{32} \frac{m_{x}^{2}}{m^{5 / 2}}+\frac{1}{8}\left(\frac{m_{x}}{m^{3 / 2}}\right)_{x}, \ldots \text { and so forth. } \tag{3.28}
\end{equation*}
$$

In this way, we can obtain local functionals; see [13].
We finish this section with the geometric integrability of one of the most popular twocomponent generalization of CH equation and of KdV6 equation.

Another generalization of the Camassa-Holm equation is the following integrable twocomponent CH system [19]:

$$
\begin{gather*}
u_{t}-u_{x x t}=-3 u u_{x}+2 u_{x} u_{x x}+u u_{x x x}+\sigma \rho \rho_{x}  \tag{3.29}\\
\rho_{t}+(u \rho)_{x}=0
\end{gather*}
$$

where $\sigma= \pm 1$. Introducing a new variable $v=\rho^{2} / 2$, the above system becomes

$$
\begin{gather*}
u_{t}-u_{x x t}=-3 u u_{x}+2 u_{x} u_{x x}+u u_{x x x}+\sigma v_{x} \\
v_{t}+2 v u_{x}+u v_{x}=0 \tag{3.30}
\end{gather*}
$$

The system (3.30) is geometrically integrable. The corresponding 1-forms, satisfying the structure equations (2.2), are the following:

$$
\begin{align*}
\omega^{1}= & \left(u_{x x}-u-\sigma \eta v+1\right) d x+\left[u^{2}-u u_{x x}+\frac{\eta+1}{\eta}\left(u_{x}-u\right)+\frac{1}{\eta}-\sigma v+\eta \sigma u v\right] d t, \\
\omega^{2}= & (\eta+1) d x+\left(\frac{\eta+1}{\eta}-(\eta+1) u+u_{x}\right) d t  \tag{3.31}\\
\omega^{3}= & \left(u_{x x}-u-\sigma \eta v+\eta+1\right) d x \\
& +\left[u^{2}-u u_{x x}+\frac{\eta-(\eta+1)^{2}}{\eta} u+\frac{\eta+1}{\eta}\left(u_{x}+1\right)-\sigma v+\eta \sigma u v\right] d t .
\end{align*}
$$

We could easily include two-component Hunter-Saxon system [19] into this picture (see also [7]).

Finally, we note that nonholonomic perturbation of KdV equation, known as KdV6 equation, is also of pseudospherical type, that is, KdV6 equation is geometrically integrable. We just give the corresponding 1-forms

$$
\begin{align*}
& \omega^{1}=(1-u) d x+\left[-u_{x x}+\eta u_{x}-2 u^{2}+\frac{1}{\eta} w_{x}-\frac{1}{\eta^{2}}\left(w_{x x}+2 u w\right)+\left(2-\eta^{2}\right) u+\frac{2}{\eta^{2}} w+\eta^{2}\right] d t \\
& \omega^{2}=\eta d x+\left(\eta^{3}+2 \eta u+\frac{2}{\eta} w-2 u_{x}-\frac{2}{\eta^{2}} w_{x}\right) d t \\
& \omega^{3}=-(1+u) d x+\left[-u_{x x}+\eta u_{x}-2 u^{2}+\frac{1}{\eta} w_{x}-\frac{1}{\eta^{2}}\left(w_{x x}+2 u w\right)-\left(2+\eta^{2}\right) u-\frac{2}{\eta^{2}} w-\eta^{2}\right] d t \tag{3.32}
\end{align*}
$$

which coincide with those for KdV equation [3] when $w \rightarrow 0$.

## 4. Concluding Remarks

In this paper, we study the CH equation and some of its generalizations from the geometric point of view. We show that Kupershmidt deformations for CH and $\mu \mathrm{CH}$ equations preserve integrability and derive some important objects like quadratic pseudopotentials which turn out to be useful for obtaining conservation laws and nonlocal symmetries. It is also shown that the KdV6 equation and two-component CH system are also geometrically integrable.

Having at hand these examples of geometrically integrable Kupershmidt deformations, it is natural to think that maybe there exists a general link in this sense: a Kupershmidt
deformation of geometrically integrable system is again geometrically integrable. We have not succeeded in establishing such a link up to now, but we believe that this is true at least for the systems with local Hamiltonian pair of operators as in the above examples.

Let us return, however, to the $\mu \mathrm{CH}$ equation ( $w=0$ ). It is obvious that pseudopotentials for the $\mu \mathrm{CH}$ equation (1.8) and parameter-dependent conservation law are obtained from

$$
\begin{gather*}
m=\frac{\gamma^{2}}{\lambda}+\frac{\gamma_{x}}{\lambda},  \tag{4.1}\\
\gamma_{t}=\frac{1}{2 \lambda}\left(\gamma+\lambda u_{x}-2 \lambda u \gamma\right)_{x}=\frac{\gamma_{x}}{2 \lambda}+\frac{u_{x x}}{2}-\partial(\gamma u) . \tag{4.2}
\end{gather*}
$$

Equation (4.1) is an analogue of the Miura transformation of KdV theory. We can repeat, purely formally, the procedure for obtaining the "modified" $\mathrm{CH}(\mathrm{mCH})$ equation [5] in this case. However, it is clear that since $\mu \mathrm{CH}$ contains nonlocal term, one can expect that the "modified" equation also will have nonlocal terms.

Denote by $A$ the operator $A=\mu-\partial^{2}, A(u)=m=\mu(u)-u_{x x}$. The operators $A^{-1}$ and $\partial$ commute and $\mu(u)=\mu(A u)$.

We have

$$
\begin{equation*}
u=A^{-1} m, \quad u_{x}=A^{-1} m_{x}, \quad u_{x x}=A^{-1} m_{x x}, \tag{4.3}
\end{equation*}
$$

in which $m$ is determined by (4.1). Then, the second equation (4.2) takes the form

$$
\begin{equation*}
\gamma_{t}=\frac{\gamma_{x}}{2 \lambda}+\frac{A^{-1} m_{x x}}{2}-\gamma A^{-1} m_{x}-\gamma_{x} A^{-1} m, \tag{4.4}
\end{equation*}
$$

or

$$
\begin{equation*}
A \gamma_{t}=\frac{A \gamma_{x}}{2 \lambda}+\frac{m_{x x}}{2}-A \partial\left(r A^{-1} m\right) \tag{4.5}
\end{equation*}
$$

Formally, this equation can be named as a "modified" $\mu \mathrm{CH}$ equation. One can simplify further this equation using (4.1) or even to present it as a system as in [5], it remains nonlocal and, hence, it is of no immediate advantage.

## Acknowledgments

This work is partially supported by Grant no. 169/2010 of Sofia University and by Grant no. DD VU 02/90 with NSF of Bulgaria.

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