# BOUNDARY VALUE PROBLEM FOR THE THREE DIMENSIONAL TIME PERIODIC VLASOV-MAXWELL SYSTEM* 

M. BOSTAN ${ }^{\dagger}$


#### Abstract

In this work we study the existence of time periodic weak solution for the three dimensional Vlasov-Maxwell system with boundary conditions. The main idea consists of using the mass, momentum and energy conservation laws which allow us to obtain a priori estimates in the case of a star-shaped bounded spatial domain. We start by constructing time periodic smooth solutions for a regularized system. The existence for the Vlasov-Maxwell system follows by weak stability under uniform estimates. These results apply for both classical and relativistic cases and for systems with several species of particles.


Key words. Equations, weak/mild formulation, regularization.
AMS subject classifications. 35F30, 35L40.

## 1. Introduction

The coupled nonlinear system given by the Vlasov-Maxwell equations is a classical model in the kinetic theory of plasma. The main assumption underlying the model is that collisions are so rare that they may be neglected.

Consider $\Omega$ an open bounded subset of $\mathbb{R}_{x}^{3}$, with boundary $\partial \Omega$ regular. We introduce the notations $\Sigma=\partial \Omega \times \mathbb{R}_{p}^{3}$ and:

$$
\begin{equation*}
\Sigma^{ \pm}=\left\{(x, p) \in \partial \Omega \times \mathbb{R}_{p}^{3} \mid \pm(v(p) \cdot n(x))>0\right\} \tag{1.1}
\end{equation*}
$$

where $n(x)$ is the unit outward normal to $\partial \Omega$ at $x$ and $v(p)$ is the velocity associated to some energy function $\mathcal{E}(p)$ by $v(p)=\nabla_{p} \mathcal{E}(p), p \in \mathbb{R}_{p}^{3}$. The functions to be considered are:

$$
\begin{equation*}
\mathcal{E}(p)=\frac{|p|^{2}}{2 m}, v(p)=\frac{p}{m} \tag{1.2}
\end{equation*}
$$

for the classical case and:

$$
\begin{equation*}
\mathcal{E}(p)=m c_{0}^{2}\left(\left(1+\frac{|p|^{2}}{m^{2} c_{0}^{2}}\right)^{1 / 2}-1\right), v(p)=\frac{p}{m}\left(1+\frac{|p|^{2}}{m^{2} c_{0}^{2}}\right)^{-1 / 2} \tag{1.3}
\end{equation*}
$$

for the relativistic case, where $m$ is the mass of particles, $c_{0}$ is the light speed in the vacuum. We denote by $f(t, x, p)$ the particles distribution depending on the time $t$, the position $x \in \Omega$ and the momentum $p \in \mathbb{R}_{p}^{3}$ and by $(E(t, x), B(t, x))$ the electromagnetic field depending on $t$ and $x$. If we note by $F(t, x, p)=q(E(t, x)+v(p) \wedge B(t, x))$ the electro-magnetic force, the Vlasov problem is given by:

$$
\begin{gather*}
\partial_{t} f+v(p) \cdot \nabla_{x} f+F(t, x, p) \cdot \nabla_{p} f=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3},  \tag{1.4}\\
f(t, x, p)=g(t, x, p),(t, x, p) \in \mathbb{R}_{t} \times \Sigma^{-}, \tag{1.5}
\end{gather*}
$$

[^0]where $q$ is the charge of particles and $g$ represents the distribution of the incoming particles, which is a given $T$ periodic function. Some other boundary conditions can be considered as we will see later on. The problem $(1.4),(1.5)$ is coupled with the Maxwell equations:
$\partial_{t} E-c_{0}^{2} \operatorname{rot} B=-\frac{j(t, x)}{\varepsilon_{0}}, \partial_{t} B+\operatorname{rot} E=0, \operatorname{div} E=\frac{\rho(t, x)}{\varepsilon_{0}}, \operatorname{div} B=0,(t, x) \in \mathbb{R}_{t} \times \Omega$,
with the boundary condition:
\[

$$
\begin{equation*}
n(x) \wedge E(t, x)+c_{0} n(x) \wedge(n(x) \wedge B(t, x))=h(t, x),(t, x) \in \mathbb{R}_{t} \times \partial \Omega \tag{1.7}
\end{equation*}
$$

\]

where $\varepsilon_{0}$ is the permittivity of the vacuum, $\rho(t, x)=q \int_{\mathbb{R}_{p}^{3}} f(t, x, p) d p$ is the charge density, $j(t, x)=q \int_{\mathbb{R}_{p}^{3}} f(t, x, p) v(p) d p$ is the current density and $h$ is a given $T$ periodic function on the boundary $\mathbb{R}_{t} \times \partial \Omega$ such that $\left.(n \cdot h)\right|_{\mathbb{R}_{t} \times \partial \Omega}=0$. We introduce also the permeability of the vacuum, $\mu_{0}$, given by $\varepsilon_{0} \mu_{0} c_{0}^{2}=1$.

When the magnetic field is neglected, the electric field derives from a potential $E=-\nabla_{x} \Phi$ and we obtain the Vlasov-Poisson system:

$$
\begin{array}{r}
\partial_{t} f+v(p) \cdot \nabla_{x} f-q \nabla_{x} \Phi \cdot \nabla_{p} f=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}, \\
f(t, x, p)=g(t, x, p),(t, x, p) \in \mathbb{R}_{t} \times \Sigma^{-}, \\
-\Delta_{x} \Phi=\frac{\rho}{\varepsilon_{0}},(t, x) \in \mathbb{R}_{t} \times \Omega, \Phi(t, x)=\varphi_{0}(t, x),(t, x) \in \mathbb{R}_{t} \times \partial \Omega
\end{array}
$$

where $\varphi_{0}$ is the potential on the boundary $\mathbb{R}_{t} \times \partial \Omega$. This model can be derived from the relativistic Vlasov-Maxwell system by letting $c_{0} \rightarrow+\infty$, see Degond [12] for the case of smooth solutions and [9] for the case of weak solutions.

Various results were obtained for the free space Vlasov-Poisson system. Weak solutions were constructed by Arseneev [1], Horst and Hunze [26]. The existence of classical solutions has been studied by Ukai and Okabe [35], Horst [25], Batt [2], Pfaffelmoser [31]. The existence of global classical solutions for the Vlasov-Poisson equations with small initial data is a result of Bardos and Degond [5], see also Schaeffer [33], [34]. A powerful method has been proposed by Lions and Perthame [30] in order to study the propagation of the moments for the three dimensional Vlasov-Poisson system. The existence of global weak solution for the Vlasov-Maxwell system in three dimensions was obtained by DiPerna and Lions [14], one of the key points being the compactness result of velocity averages (see also [21]). The global existence of a strong solution is still an open problem. Results for the relativistic case were obtained by Glassey and Schaeffer [17], [18], Glassey and Strauss [19], [20], Klainerman and Staffilani [27], Bouchut, Golse and Pallard [10].

However, for applications like vacuum diodes, tube discharges, cold plasma, solar wind, satellite ionization, thrusters, etc. boundary conditions have to be taken into account. Results for the initial-boundary value problem were obtained by Ben Abdallah [6] for the Vlasov-Poisson system in three dimensions and Guo [23] for the Vlasov-Maxwell system. Permanent regimes are particularly important. They are of two types and they are modeled by stationary solutions or time periodic solutions for boundary value problems. Another strong motivation to study such solutions is the great difficulty to compute it numerically. The stationary problem for the VlasovPoisson equations was studied by Greengard and Raviart [22] in one dimension and by Poupaud [32] in three dimensions for the Vlasov-Maxwell system. An asymptotic
analysis of the Vlasov-Poisson system was done by Degond and Raviart [13] in the case of the plane diode. The regularity of the solutions for the Vlasov-Maxwell system has been studied by Guo [24]. Results for the one dimensional time periodic case can be found in [8] for the Vlasov-Poisson system and in [7] for the Vlasov-Maxwell system.

The aim of this paper is to prove the existence of time $T$ periodic solution for the three dimensional Vlasov-Maxwell system (1.4),(1.5), (1.6),(1.7) when the boundary conditions are supposed $T$ periodic, with $T>0$ fixed. The techniques introduced for the analysis of the one dimensional case cannot be applied in the three dimensional case. Indeed, the proof in [7] relies on the fact that in one dimension the solution of the Maxwell equations can be computed explicitly, where in three dimensions such a formula is not available. Our main result is the following theorem.

Theorem 1.1. Assume that $\Omega$ is bounded, with $\partial \Omega$ smooth and strictly star-shaped, $g \in L^{\infty}\left(\mathbb{R}_{t} \times \Sigma^{-}\right), h$ are $T$ periodic such that $g \geq 0,\left.(n \cdot h)\right|_{\mathbb{R}_{t} \times \partial \Omega}=0$ and:

$$
W_{0}:=\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))|(1+\mathcal{E}(p)) g(t, x, p) d t d \sigma d p+\int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma<+\infty
$$

Then there is a $T$ periodic weak solution $(f \geq 0, E, B) \in L^{\infty}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right) \times$ $L_{\text {loc }}^{2}\left(\mathbb{R}_{t} ; L^{2}(\Omega)^{3}\right)^{2}$ for the Vlasov-Maxwell system (classical or relativistic case):

$$
\partial_{t} f+v(p) \cdot \nabla_{x} f+q(E(t, x)+v(p) \wedge B(t, x)) \cdot \nabla_{p} f=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}
$$

$$
\begin{gathered}
\partial_{t} E-c_{0}^{2} \operatorname{rot} B=-\frac{j}{\varepsilon_{0}}, \partial_{t} B+\operatorname{rot} E=0, \text { div } E=\frac{\rho}{\varepsilon_{0}}, \text { div } B=0,(t, x) \in \mathbb{R}_{t} \times \Omega, \\
f(t, x, p)=g(t, x, p),(t, x, p) \in \mathbb{R}_{t} \times \Sigma^{-}, \\
n \wedge E(t, x)+c_{0} n \wedge(n \wedge B(t, x))=h(t, x),(t, x) \in \mathbb{R}_{t} \times \partial \Omega .
\end{gathered}
$$

Moreover the continuity equation is satisfied $\partial_{t} \rho+\operatorname{div} j=0$ in $\mathcal{D}^{\prime}(] 0, T[\times \Omega)$, there are trace functions $\gamma^{+} f \geq 0,\left\|\gamma^{+} f\right\|_{\infty} \leq\|g\|_{\infty}$, normal and tangential traces $(n \cdot E, n$. $B),(n \wedge E, n \wedge B)$ and for some constant $C$ depending on $m, \varepsilon_{0}, \mu_{0}, \Omega$ we have:

$$
\begin{aligned}
& \text { ess } \sup _{s \in \mathbb{R}}\left\{\int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f(s, x, p) d x d p+\int_{\Omega}\left\{\varepsilon_{0}|E(s, x)|^{2}+\frac{1}{\mu_{0}}|B(s, x)|^{2}\right\} d x\right\} \\
& +\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x))(1+\mathcal{E}(p)) \gamma^{+} f(t, x, p) d t d \sigma d p \\
& +\int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}\left[(n \cdot E)^{2}+|n \wedge E|^{2}\right]+\frac{1}{\mu_{0}}\left[(n \cdot B)^{2}+|n \wedge B|^{2}\right]\right\} d t d \sigma \\
& \leq C \cdot W_{0} \text {. }
\end{aligned}
$$

As in [14] we construct a $T$ periodic weak solution as weak limit of solutions for a regularized system. One of the crucial points consists of finding uniform a priori bounds for the total (kinetic and electro-magnetic) energy by using the physical conservation laws. As usual we prove the existence of a solution for the regularized problem by the fixed point method. By the conservation laws of the mass and the energy we obtain:

$$
\begin{align*}
& \frac{d}{d t} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p)(1+\mathcal{E}(p)) d x d p+\frac{\varepsilon_{0}}{2} \frac{d}{d t} \int_{\Omega}\left(|E(t, x)|^{2}+c_{0}^{2}|B(t, x)|^{2}\right) d x \\
& +\int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f(t, x, p)(1+\mathcal{E}(p)) d \sigma d p \\
& \quad+\frac{\varepsilon_{0} c_{0}}{2} \int_{\partial \Omega}\left(|n \wedge E|^{2}+c_{0}^{2}|n \wedge B|^{2}\right) d \sigma \\
& =\int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(t, x, p)(1+\mathcal{E}(p)) d \sigma d p+\frac{\varepsilon_{0} c_{0}}{2} \int_{\partial \Omega}|h(t, x)|^{2} d \sigma, t \in \mathbb{R}_{t} \tag{1.8}
\end{align*}
$$

If for the initial-boundary value problems the equation (1.8) provides immediately bounds for the total energy after integration on $[0, t], t>0$, the situation is different for the time periodic case, since in this case initial data are not available. Nevertheless, after integration of (1.8) over a period we obtain:

$$
\begin{gather*}
\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f(t, x, p)(1+\mathcal{E}(p)) d t d \sigma d p \\
+\frac{\varepsilon_{0} c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}\left(|n \wedge E|^{2}+c_{0}^{2}|n \wedge B|^{2}\right) d t d \sigma \\
=\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(t, x, p)(1+\mathcal{E}(p)) d t d \sigma d p+\frac{\varepsilon_{0} c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma . \tag{1.9}
\end{gather*}
$$

In order to estimate the total energy we use also the momentum conservation law. We suppose that $\partial \Omega$ is strictly star-shaped with respect to some point $x_{0} \in \Omega$ i.e., $\exists r>0$ such that $\left(n(x) \cdot\left(x-x_{0}\right)\right) \geq r, \forall x \in \partial \Omega$, and we multiply the Vlasov equation by the test function $\left(p \cdot\left(x-x_{0}\right)\right)$ :

$$
\begin{array}{rl}
\frac{d}{d t} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} & f(t, x, p)\left(p \cdot\left(x-x_{0}\right)\right) d x d p+\int_{\Sigma}(v(p) \cdot n(x)) \gamma f(t, x, p)\left(p \cdot\left(x-x_{0}\right)\right) d \sigma d p \\
= & \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p)(v(p) \cdot p) d x d p+\int_{\Omega} \int_{\mathbb{R}_{p}^{3}} q f(t, x, p)(E+v(p) \wedge B) \cdot\left(x-x_{0}\right) d x d p \tag{1.10}
\end{array}
$$

In the previous equality $\gamma f$ denotes the trace of $f$ on $\mathbb{R}_{t} \times \Sigma$ (in fact $\gamma f=\gamma^{+} f$ on $\mathbb{R}_{t} \times \Sigma^{+}$and $\gamma f=g$ on $\mathbb{R}_{t} \times \Sigma^{-}$). By using the Maxwell equations the last integral in the above equation can be written:

$$
\begin{array}{rl}
\int_{\Omega} \int_{\mathbb{R}_{p}^{3}} & q f(t, x, p)(E+v(p) \wedge B) \cdot\left(x-x_{0}\right) d x d p=\int_{\Omega}(\rho E+j \wedge B) \cdot\left(x-x_{0}\right) d x \\
\quad= & \varepsilon_{0} \int_{\Omega}\left[(E \operatorname{div} E-E \wedge \operatorname{rot} E)+c_{0}^{2}(B \operatorname{div} B-B \wedge \operatorname{rot} B)\right] \cdot\left(x-x_{0}\right) d x \\
& -\varepsilon_{0} \int_{\Omega} \partial_{t}(E \wedge B) \cdot\left(x-x_{0}\right) d x \tag{1.11}
\end{array}
$$

We use also the identity ( $u$ div $u-u \wedge \operatorname{rot} u)_{i}=\sum_{j=1}^{3} \frac{\partial}{\partial x_{j}}\left(u_{i} u_{j}\right)-\frac{1}{2} \frac{\partial}{\partial x_{i}}|u|^{2}, 1 \leq i \leq 3$ and the inequality $(v(p) \cdot p) \geq \mathcal{E}(p), \forall p \in \mathbb{R}_{p}^{3}$. After integration by parts and direct
computations the equations (1.10), (1.11) yield:

$$
\begin{align*}
& \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f \mathcal{E}(p) d t d x d p+\frac{\varepsilon_{0}}{2} \int_{0}^{T} \int_{\Omega}\left(|E|^{2}+c_{0}^{2}|B|^{2}\right) d t d x \\
&+\frac{\varepsilon_{0} r}{2} \int_{0}^{T} \int_{\partial \Omega}\left[(n \cdot E)^{2}+c_{0}^{2}\left(n \cdot B_{0}\right)^{2}\right] d t d \sigma \\
& \leq R \int_{0}^{T} \int_{\Sigma}|(v(p) \cdot n(x))| \cdot|p| \gamma f d t d \sigma d p+\frac{\varepsilon_{0} R}{2} \int_{0}^{T} \int_{\partial \Omega}\left[|n \wedge E|^{2}+c_{0}^{2}|n \wedge B|^{2}\right] d t d \sigma \\
&+\varepsilon_{0} R \int_{0}^{T} \int_{\partial \Omega}\left[|(n \cdot E)| \cdot|n \wedge E|+c_{0}^{2}|(n \cdot B)| \cdot|n \wedge B|\right] d t d \sigma \tag{1.12}
\end{align*}
$$

where $R=\sup _{x \in \partial \Omega}\left|x-x_{0}\right|$. The estimate (1.12) together with (1.9) clearly give the desired bounds. There is another important point to be clarified: in the above computations we used the divergence equations $\operatorname{div} E=\frac{\rho}{\varepsilon_{0}}$, $\operatorname{div} B=0$. In the case of the initial value problem these equations hold true as soon as they are verified by the initial data. In the time periodic case the idea is to regularize the Vlasov-Maxwell equations and to use the time periodicity. Indeed we consider $T$ periodic solutions for the perturbed equations:

$$
\begin{aligned}
& \alpha f+\partial_{t} f+v(p) \cdot \nabla_{x} f+q(E+v(p) \wedge B) \cdot \nabla_{p} f=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3} \\
& \alpha E+\partial_{t} E-c_{0}^{2} \operatorname{rot} B=-\frac{1}{\varepsilon_{0}} j(t, x), \alpha B+\partial_{t} B+\operatorname{rot} E=0,(t, x) \in \mathbb{R}_{t} \times \Omega
\end{aligned}
$$

where $\alpha>0$ is a small parameter. Note that the continuity equation is in this case $\alpha \rho+\partial_{t} \rho+\operatorname{div} j=0$. As usual, by taking the divergence in the perturbed Maxwell equations we find $\left(\alpha+\partial_{t}\right)\left(\operatorname{div} E-\frac{\rho}{\varepsilon_{0}}\right)=0,\left(\alpha+\partial_{t}\right) \operatorname{div} B=0$ and by time periodicity we deduce that $\operatorname{div} E=\frac{\rho}{\varepsilon_{0}}$, $\operatorname{div} B=0$. Once we have obtained uniform estimates for the total energy of the solutions for the regularized Vlasov-Maxwell system, the existence of the $T$ periodic weak solution for the non perturbed Vlasov-Maxwell system follows easily by weak stability results (cf. [14]).

The content of this paper is organized as follows: first we recall some basic definitions and results concerning the Vlasov problem. In the next section we prove general existence and uniqueness results of the time periodic solution for evolution equations with time periodic source terms. In particular, existence and uniqueness results for the regularized Maxwell equations are obtained in Section 4. In section 5 we prove the existence for the regularized Vlasov-Maxwell system by using a fixed point technique. In the next section we obtain the a priori estimates by using the conservation laws of the mass, momentum and energy. In section 7 we justify the passing to the limit for the sequence of regularized solutions. We end with some remarks concerning the system with specular a boundary condition.

## 2. The Vlasov equation

The Vlasov equation describes the evolution of a population of charged particles under the action of the electro-magnetic force. In this section we suppose that the electro-magnetic field is a given $T$ periodic function $(E, B)$. The time periodic Vlasov problem is given by:

$$
\begin{gather*}
\partial_{t} f+v(p) \cdot \nabla_{x} f+F(t, x, p) \cdot \nabla_{p} f=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}  \tag{2.1}\\
f(t, x, p)=g(t, x, p),(t, x, p) \in \mathbb{R}_{t} \times \Sigma^{-} \tag{2.2}
\end{gather*}
$$

By taking into account that $\nabla_{(x, p)} \cdot(v(p), F(t, x, p))=0$, the equation (2.1) can be written also:

$$
\partial_{t} f+\nabla_{x} \cdot(v(p) f)+\nabla_{p} \cdot(F(t, x, p) f)=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}
$$

Since there is no uniqueness for the Vlasov problem (2.1), (2.2) (because the distribution function can take arbitrary constant values on the characteristics which remain in the domain), it is convenient to consider also the perturbed problem:

$$
\begin{equation*}
\alpha f+\partial_{t} f+v(p) \cdot \nabla_{x} f+F(t, x, p) \cdot \nabla_{p} f=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3} \tag{2.3}
\end{equation*}
$$

with the boundary condition (2.2), where $\alpha>0$ is fixed. We introduce the definitions of weak/mild solution for the perturbed Vlasov problem:
Definition 2.1. Assume that $E, B \in L^{\infty}\left(\mathbb{R}_{t} \times \Omega\right)^{3}$ and $(v(p) \cdot n(x)) g \in L_{l o c}^{1}\left(\mathbb{R}_{t} \times \Sigma^{-}\right)$ are $T$ periodic. We say that $f \in L_{\text {loc }}^{1}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right)$ is a $T$ periodic weak solution for the perturbed Vlasov problem (2.3),(2.2) iff $f$ is $T$ periodic and:

$$
\begin{align*}
\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p)\left(\alpha \varphi-\partial_{t} \varphi\right. & \left.-v(p) \cdot \nabla_{x} \varphi-F(t, x, p) \cdot \nabla_{p} \varphi\right) d t d x d p \\
& =-\int_{0}^{T} \int_{\Sigma^{-}}(v(p) \cdot n(x)) g(t, x, p) \varphi(t, x, p) d t d \sigma d p \tag{2.4}
\end{align*}
$$

for all test functions which belong to:

$$
\mathcal{T}_{w}=\left\{\varphi \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega} \times \mathbb{R}_{p}^{3}\right)\left|\exists R>0: \varphi=\varphi \cdot \mathbf{1}_{\{|p| \leq R\}}, \varphi\right|_{\mathbb{R}_{t} \times \Sigma^{+}}=0, \varphi(\cdot+T)=\varphi\right\}
$$

REmARK 2.2. In the above definition we can assume that $E, B$ are only in $L^{s}(] 0, T[\times \Omega)^{3}$ by requiring more regularity on $f$, namely $f \in L_{l o c}^{r}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right)$, where $s$ is the conjugate exponent of $r$.

Suppose now that $E, B$ are $T$ periodic and belong to $L^{\infty}\left(\mathbb{R}_{t} ; W^{1, \infty}(\Omega)\right)^{3}$. In this case we can define the notion of solution by characteristics or mild solution. First of all let us introduce the characteristics: for $(t, x, p) \in \mathbb{R}_{t} \times \bar{\Omega} \times \mathbb{R}_{p}^{3}$ we denote by $(X(s), P(s))=(X(s ; t, x, p), P(s ; t, x, p))$ the unique solution of the system:

$$
\begin{equation*}
\frac{d X}{d s}=v(P(s ; t, x, p)), \frac{d P}{d s}=F(s, X(s ; t, x, p), P(s ; t, x, p)), s_{\text {in }}(t, x, p) \leq s \leq s_{\text {out }}(t, x, p) \tag{2.5}
\end{equation*}
$$

with the conditions $X(s=t ; t, x, p)=x, P(s=t ; t, x, p)=p$. Here $s_{\text {in }}, s_{\text {out }}$ represent the incoming, respectively outgoing time given by:

$$
\begin{aligned}
& s_{\text {in }}(t, x, p)=\sup \{s \leq t \mid X(s ; t, x, p) \in \partial \Omega\}, \\
& s_{\text {out }}(t, x, p)=\inf \{s \geq t \mid X(s ; t, x, p) \in \partial \Omega\} .
\end{aligned}
$$

By using the time periodicity of the electro-magnetic field we check easily that

$$
\begin{equation*}
s_{\text {in }}(t+T, x, p)=s_{\text {in }}(t, x, p)+T, s_{\text {out }}(t+T, x, p)=s_{\text {out }}(t, x, p)+T \tag{2.6}
\end{equation*}
$$

and:

$$
\begin{equation*}
X(s+T ; t+T, x, p)=X(s ; t, x, p), P(s+T ; t+T, x, p)=P(s ; t, x, p) \tag{2.7}
\end{equation*}
$$

for all $(t, x, p) \in \mathbb{R}_{t} \times \bar{\Omega} \times \mathbb{R}_{p}^{3}, s_{\text {in }}(t, x, p) \leq s \leq s_{\text {out }}(t, x, p)$. The mild formulation follows formally by solving:

$$
\alpha \varphi-\partial_{t} \varphi-v(p) \cdot \nabla_{x} \varphi-F(t, x, p) \cdot \nabla_{p} \varphi=\psi(t, x, p),(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}
$$

with the boundary condition $\left.\varphi\right|_{\mathbb{R}_{t} \times \Sigma^{+}}=0$. By integration along the characteristic curves we obtain:

$$
\varphi_{\psi}^{\alpha}(t, x, p)=\int_{t}^{s_{o u t}(t, x, p)} e^{-\alpha(s-t)} \psi(s, X(s ; t, x, p), P(s ; t, x, p)) d s
$$

and we define the mild solution by:
Definition 2.3. Assume that $E, B \in L^{\infty}\left(\mathbb{R}_{t} ; W^{1, \infty}(\Omega)\right)^{3}$ and $(v(p) \cdot n(x)) g \in$ $L_{\text {loc }}^{1}\left(\mathbb{R}_{t} \times \Sigma^{-}\right)$are $T$ periodic, $\alpha>0$. We say that $f \in L_{\text {loc }}^{1}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right)$ is a $T$ periodic mild solution for the perturbed Vlasov problem (2.3),(2.2) iff $f$ is $T$ periodic and:

$$
\begin{equation*}
\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p) \psi(t, x, p) d t d x d p=-\int_{0}^{T} \int_{\Sigma^{-}}(v(p) \cdot n(x)) g(t, x, p) \varphi_{\psi}^{\alpha}(t, x, p) d t d \sigma d p \tag{2.8}
\end{equation*}
$$

for all test functions which belong to:

$$
\mathcal{T}_{m}=\left\{\psi \in C^{0}\left(\mathbb{R}_{t} \times \bar{\Omega} \times \mathbb{R}_{p}^{3}\right) \mid \exists R>0: \psi=\psi \cdot \mathbf{1}_{\{|p| \leq R\}}, \psi(\cdot+T)=\psi\right\}
$$

For $\alpha=0$ one gets the definitions of the weak/mild solution for the Vlasov problem $(2.1),(2.2)$. The existence of the $T$ periodic mild solution is a standard result and follows by change of variables along characteristics (see also the Remark 2.6).
Proposition 2.4. Assume that $E, B \in L^{\infty}\left(\mathbb{R}_{t} ; W^{1, \infty}(\Omega)\right)^{3}$ and $g \in L^{\infty}\left(\mathbb{R}_{t} \times \Sigma^{-}\right)$are $T$ periodic, $\alpha>0$. Then the perturbed Vlasov problem (2.3),(2.2) has a unique $T$ periodic mild solution $f \in L^{\infty}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right)$, verifying $\|f\|_{\infty} \leq\|g\|_{\infty}$. Moreover, if $g \geq$ 0 then $f \geq 0$.
REmark 2.5. It is easy to check that all $T$ periodic mild solutions are also $T$ periodic weak solutions.
REMARK 2.6. It is well known that the mild solution of problem (2.3), (2.2) is given by $f(t, x, p)=e^{-\alpha\left(t-s_{i n}(t, x, p)\right)} g\left(s_{i n}, X\left(s_{i n} ; t, x, p\right), P\left(s_{i n} ; t, x, p\right)\right)$ if $s_{i n}(t, x, p)>-\infty$ and $f(t, x, p)=0$ otherwise. Observe that when the electro-magnetic field and the boundary data $g$ are $T$ periodic, the equalities (2.6), (2.7) imply immediately that $f$ is $T$ periodic.
Remark 2.7. Under the same hypothesis as in Proposition 2.4, the $T$ periodic mild solution $f$ has a trace $\gamma^{+} f \in L^{\infty}\left(\mathbb{R}_{t} \times \Sigma^{+}\right)$verifying the following Green formula:

$$
\begin{align*}
& \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p)\left(\alpha \varphi-\partial_{t} \varphi-v(p) \cdot \nabla_{x} \varphi-F(t, x, p) \cdot \nabla_{p} \varphi\right) d t d x d p \\
= & -\int_{0}^{T} \int_{\Sigma^{-}}(v(p) \cdot n(x)) g(t, x, p) \varphi(t, x, p) d t d \sigma d p \\
& -\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f(t, x, p) \varphi(t, x, p) d t d \sigma d p \tag{2.9}
\end{align*}
$$

for all $\varphi \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega} \times \mathbb{R}_{p}^{3}\right)$ with compact support in momentum and $T$ periodic in time. The trace $\gamma^{+} f$ is given by the same formula as in the Remark 2.6 and we have $\left\|\gamma^{+} f\right\|_{\infty} \leq\|g\|_{\infty}$. Moreover, if $g \geq 0$ then $\gamma^{+} f \geq 0$.
Proposition 2.8. Under the same hypothesis as in Proposition 2.4, a T periodic bounded weak solution of the perturbed Vlasov problem (2.3),(2.2) is unique and therefore coincides with the $T$ periodic mild solution.

Proof. Assume that $f \in L^{\infty}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right)$ is a $T$ periodic weak solution with boundary data $g=0$. We have $\partial_{t} f+v(p) \cdot \nabla_{x} f+F(t, x, p) \cdot \nabla_{p} f=-\alpha f \in L^{\infty}\left(\mathbb{R}_{t} \times \Omega \times\right.$ $\mathbb{R}_{p}^{3}$ ) and therefore (cf. [4], [15]) we obtain:

$$
\frac{1}{2}\left(\partial_{t} f^{2}+v(p) \cdot \nabla_{x} f^{2}+F \cdot \nabla_{p} f^{2}\right)=-\alpha f^{2}
$$

After integration on $] 0, T\left[\times \Omega \times \mathbb{R}_{p}^{3}\right.$ we deduce that:

$$
\alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f^{2}(t, x, p) d t d x d p+\frac{1}{2} \int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x))\left(\gamma^{+} f\right)^{2}(t, x, p) d t d \sigma d p=0
$$

or $f=0, \gamma^{+} f=0$.
Proposition 2.9. Under the same hypothesis as in Proposition 2.4, with $g \geq 0$, $\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(t, x, p) d t d \sigma d p<+\infty$, the $T$ periodic mild/weak solution belongs to $L^{1}(] 0, T\left[\times \Omega \times \mathbb{R}_{p}^{3}\right)$ and

$$
\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p) d t d x d p \leq \frac{1}{\alpha} \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(t, x, p) d t d \sigma d p
$$

Proof. By applying the mild formulation with the test function $\psi(t, x, p)=\chi_{R}(|p|)$ where $\chi_{R}(u)=\chi\left(\frac{u}{R}\right), \chi \in C_{c}(\mathbb{R}), \chi(u)=1$ if $|u| \leq 1, \chi(u)=0$ if $|u| \geq 2$ and $0 \leq \chi \leq 1$, we deduce that $0 \leq \varphi_{\psi}^{\alpha} \leq \frac{1}{\alpha}$ and therefore:

$$
\begin{aligned}
& \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f \mathbf{1}_{\{|p| \leq R\}} d t d x d p \leq \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f \psi d t d x d p \\
& \quad \leq \frac{1}{\alpha} \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g d t d \sigma d p, \forall R>0
\end{aligned}
$$

The conclusion follows by letting $R \rightarrow+\infty$ and by using the monotone convergence theorem.

## 3. Time periodic evolution equations

We intend to solve the time periodic Maxwell equations by using standard theory of maximal monotone operators. We present here some easy results of existence and uniqueness for linear evolution equations with time periodic source terms. We need the following lemma.

Lemma 3.1. Assume that $g \in L_{l o c}^{1}\left(\mathbb{R}_{t} ; \mathbb{R}\right)$ is a $T$ periodic function and $\alpha>0$ is fixed. Then:

$$
\left|\int_{0}^{t} e^{-\alpha(t-s)} g(s) d s\right| \leq\left(\frac{1}{\alpha T}+4\right)\|g\|_{L^{1}(] 0, T[)}, \forall t \geq 0
$$

Proof. Consider $G:\left[0,+\infty\left[\rightarrow \mathbb{R}, G(t)=\int_{0}^{t}\{g(s)-\langle g\rangle\} d s\right.\right.$, where $\langle g\rangle:=\frac{1}{T} \int_{0}^{T} g(t)$ $d t$. Obviously $G$ is $T$ periodic and bounded $|G(t)| \leq 2 \cdot\|g\|_{L^{1}(] 0, T[)}, \forall t \geq 0$. We have:

$$
\begin{aligned}
\int_{0}^{t} e^{-\alpha(t-s)} g(s) d s & =\int_{0}^{t} e^{-\alpha(t-s)}(g(s)-\langle g\rangle) d s+\langle g\rangle \cdot \int_{0}^{t} e^{-\alpha(t-s)} d s \\
& =\int_{0}^{t} e^{-\alpha(t-s)} G^{\prime}(s) d s+\alpha^{-1}\left(1-e^{-\alpha t}\right)\langle g\rangle \\
& =G(t)+\alpha^{-1}\left(1-e^{-\alpha t}\right)\langle g\rangle-\alpha \int_{0}^{t} e^{-\alpha(t-s)} G(s) d s
\end{aligned}
$$

Finally we deduce that:

$$
\begin{align*}
\left|\int_{0}^{t} e^{-\alpha(t-s)} g(s) d s\right| & \leq|G(t)|+\alpha^{-1}|\langle g\rangle|+\|G\|_{L^{\infty}} \cdot \int_{0}^{t} \alpha e^{-\alpha(t-s)} d s \\
& \leq\left(\frac{1}{\alpha T}+4\right)\|g\|_{L^{1}(] 0, T[)} \tag{3.1}
\end{align*}
$$

Proposition 3.2. Assume that $A: D(A) \subset H \rightarrow H$ is a linear maximal monotone operator on a Hilbert space $H, f \in C^{1}\left(\mathbb{R}_{t} ; H\right)$ is $T$ periodic, $\alpha>0$ fixed. Then there is a unique $T$ periodic solution $\xi \in C\left(\mathbb{R}_{t} ; D(A)\right) \cap C^{1}\left(\mathbb{R}_{t} ; H\right)$ for the perturbed evolution equation:

$$
\begin{equation*}
\alpha x(t)+x^{\prime}(t)+A x(t)=f(t), t \in \mathbb{R}_{t} \tag{3.2}
\end{equation*}
$$

Moreover, we have the following estimates:

$$
\|\xi\|_{L^{\infty}\left(\mathbb{R}_{t} ; H\right)} \leq\left(\frac{1}{\alpha T}+4\right)\|f\|_{L^{1}(] 0, T[; H)},\left\|\xi^{\prime}\right\|_{L^{\infty}\left(\mathbb{R}_{t} ; H\right)} \leq\left(\frac{1}{\alpha T}+4\right)\left\|f^{\prime}\right\|_{L^{1}(] 0, T[; H)}
$$

Proof. Consider an arbitrary $x_{0} \in D(A)$ and denote by $x\left(\cdot ; 0, x_{0}\right) \in$ $C\left(\left[0,+\infty[; D(A)) \cap C^{1}([0,+\infty[; H)\right.\right.$ the unique solution of (3.2) with the initial condition $x_{0}$. We have:

$$
\alpha(x(t+T)-x(t))+x^{\prime}(t+T)-x^{\prime}(t)+A x(t+T)-A x(t)=0, t \geq 0
$$

After multiplication by $x(t+T)-x(t)$, by using the monotonicity of $A$ we obtain:

$$
\frac{d}{d t}\left\{e^{2 \alpha t}\|x(t+T)-x(t)\|^{2}\right\} \leq 0, t \geq 0
$$

which implies that $\quad\|x(t+T)-x(t)\| \leq e^{-\alpha t}\|x(T)-x(0)\| \leq e^{-\alpha t}\left(2\left\|x_{0}\right\|+\right.$ $\left.\|f\|_{L^{1}(\mathrm{j}, T[; H)}\right)$. If we denote by $\left(x_{n}\right)_{n}$ the functions $x_{n}(t)=x(n T+t), 0 \leq t \leq T, n \geq 0$ we deduce that:

$$
\begin{aligned}
&\left\|x_{n+1}(t)-x_{n}(t)\right\| \leq e^{-\alpha(n T+t)}\left(2\left\|x_{0}\right\|+\|f\|_{L^{1}(] 0, T[; H)}\right) \\
& \leq e^{-\alpha n T}\left(2\left\|x_{0}\right\|+\|f\|_{L^{1}(] 0, T[; H)}\right)
\end{aligned}
$$

and therefore $x_{n} \rightarrow \xi$ in $C([0, T] ; H)$. We have

$$
\xi(T)=\lim _{n \rightarrow+\infty} x_{n}(T)=\lim _{n \rightarrow+\infty} x(n T+T)=\lim _{n \rightarrow+\infty} x_{n+1}(0)=\xi(0)
$$

With the notation $y_{h}(t)=x(t+h)-x(t)$, we have:

$$
\alpha\left(y_{h}(t+T)-y_{h}(t)\right)+y_{h}^{\prime}(t+T)-y_{h}^{\prime}(t)+A y_{h}(t+T)-A y_{h}(t)=0, t \geq 0 .
$$

After multiplication by $y_{h}(t+T)-y_{h}(t)$ we deduce as before that $\frac{1}{h} \| y_{h}(t+T)-$ $y_{h}(t)\left\|\leq e^{-\alpha t} \frac{1}{h}\right\| y_{h}(T)-y_{h}(0) \|, h>0, t \geq 0$ and by passing $h \searrow 0$ we obtain:

$$
\left\|x^{\prime}(t+T)-x^{\prime}(t)\right\| \leq e^{-\alpha t}\left\|x^{\prime}(T)-x^{\prime}(0)\right\| \leq e^{-\alpha t}\left(2\left\|x^{\prime}(0)\right\|+\left\|f^{\prime}\right\|_{L^{1}(] 0, T[; H)}\right)
$$

In particular $\left\|x_{n+1}^{\prime}(t)-x_{n}^{\prime}(t)\right\| \leq e^{-\alpha n T}\left(2\left\|x^{\prime}(0)\right\|+\left\|f^{\prime}\right\|_{L^{1}(] 0, T[; H)}\right)$ and therefore $x_{n}^{\prime} \rightarrow$ $\eta$ in $C([0, T] ; H)$. Now, by taking into account that $A$ is closed and $\left[x_{n}(t), f(t)-\right.$ $\left.\alpha x_{n}(t)-x_{n}^{\prime}(t)\right] \in \operatorname{Graph}(A), 0 \leq t \leq T, n \geq 0$ we find by passing to the limit for $n \rightarrow+\infty$ that $[\xi(t), f(t)-\alpha \xi(t)-\eta(t)] \in \operatorname{Graph}(A), 0 \leq t \leq T$ or $\xi(t) \in D(A)$ and $\alpha \xi(t)+\eta(t)+$ $A \xi(t)=f(t), 0 \leq t \leq T$. It is easy to check that $\eta=\xi^{\prime}$ and thus $\xi \in C([0, T] ; D(A)) \cap$ $C^{1}([0, T] ; H)$ is a $T$ periodic solution for (3.2). In order to estimate $\xi$ observe that:

$$
\alpha\|x(t)\|^{2}+\frac{1}{2} \frac{d}{d t}\|x(t)\|^{2} \leq\|f(t)\| \cdot\|x(t)\|, t \geq 0
$$

which implies by using the Bellman lemma that:

$$
\|x(t)\| \leq e^{-\alpha t}\|x(0)\|+\int_{0}^{t} e^{-\alpha(t-s)}\|f(s)\| d s
$$

From the Lemma 3.1 we deduce that $\left\|x_{n}(t)\right\| \leq e^{-\alpha(n T+t)}\left\|x_{0}\right\|+$ $\left(\frac{1}{\alpha T}+4\right)\|f\|_{L^{1}(] 0, T[; H)}$, or $\|\xi\|_{L^{\infty}\left(\mathbb{R}_{t} ; H\right)} \leq\left(\frac{1}{\alpha T}+4\right)\|f\|_{L^{1}(j 0, T[; H)}$. In order to estimate $\xi^{\prime}$ we write $\alpha y_{h}(t)+y_{h}^{\prime}(t)+A y_{h}(t)=f(t+h)-f(t)$ and as before we deduce that for $h>0, t \geq 0$ :

$$
\frac{1}{h}\left\|y_{h}(t)\right\| \leq e^{-\alpha t} \frac{1}{h}\left\|y_{h}(0)\right\|+\int_{0}^{t} e^{-\alpha(t-s)} \frac{1}{h}\|f(s+h)-f(s)\| d s
$$

By passing to the limit for $h \searrow 0$ one gets:

$$
\left\|x^{\prime}(t)\right\| \leq e^{-\alpha t}\left\|x^{\prime}(0)\right\|+\int_{0}^{t} e^{-\alpha(t-s)}\left\|f^{\prime}(s)\right\| d s
$$

and thus by using the Lemma 3.1 finally we find that $\left\|\xi^{\prime}\right\|_{L^{\infty}\left(\mathbb{R}_{t} ; H\right)} \leq\left(\frac{1}{\alpha T}+\right.$ 4) $\left\|f^{\prime}\right\|_{L^{1}(j 0, T[; H)}$. The uniqueness of the periodic solution follows easily by standard arguments: consider $\xi_{1}, \xi_{2}$ two periodic solutions. We have as before that:

$$
\alpha\left\|\xi_{1}(t)-\xi_{2}(t)\right\|^{2}+\frac{1}{2} \frac{d}{d t}\left\|\xi_{1}(t)-\xi_{2}(t)\right\|^{2}+\left(A \xi_{1}(t)-A \xi_{2}(t), \xi_{1}(t)-\xi_{2}(t)\right)=0
$$

After integration on $[0, T]$ one gets that $\alpha \int_{0}^{T}\left\|\xi_{1}(t)-\xi_{2}(t)\right\|^{2} d t \leq 0$, or $\xi_{1}=\xi_{2}$.
Remark 3.3. The previous result holds for $f \in W_{l o c}^{1,1}\left(\mathbb{R}_{t} ; H\right), T$ periodic (see also [3], p. 138).

Indeed, approximate $f$ by $f_{\varepsilon}$ in $W_{\text {loc }}^{1,1}\left(\mathbb{R}_{t} ; H\right)$ with $f_{\varepsilon} \in C^{1}\left(\mathbb{R}_{t} ; H\right), T$ periodic and notice that the corresponding solutions $\left(x_{\varepsilon}\right)_{\varepsilon>0}$ converge in $C^{1}\left(\mathbb{R}_{t} ; H\right)$.
REmark 3.4. The equation $\alpha x(t)-x^{\prime}(t)+A x(t)=f(t), t \in \mathbb{R}_{t}$ has also a unique $T$ periodic solution verifying the same estimates (take $x(t)=y(T-t), 0 \leq t \leq T$, where $y$ solves $\alpha y+y^{\prime}+A y=\tilde{f}$, with $\left.\tilde{f}(t)=f(T-t), 0 \leq t \leq T\right)$.

By transposition we can define the notion of $T$ periodic weak solution as follows:
Definition 3.5. Assume that $A: D(A) \subset H \rightarrow H$ is a linear operator densely defined on a Hilbert space, $f \in L_{\text {loc }}^{1}\left(\mathbb{R}_{t} ; H\right)$, $T$ periodic. We say that $x \in C\left(\mathbb{R}_{t} ; H\right)$ is a $T$ periodic weak solution of (3.2) iff $x$ is $T$ periodic and:

$$
\int_{0}^{T}\left\langle x(t), \alpha \varphi(t)-\varphi^{\prime}(t)+A^{\star} \varphi(t)\right\rangle d t=\int_{0}^{T}\langle f(t), \varphi(t)\rangle d t
$$

for all $\varphi \in C\left(\mathbb{R}_{t} ; D\left(A^{\star}\right)\right) \cap C^{1}\left(\mathbb{R}_{t} ; H\right)$, $T$ periodic.
Proposition 3.6. If $A: D(A) \subset H \rightarrow H$ is a linear maximal monotone operator on a Hilbert space, $f \in L_{l o c}^{1}\left(\mathbb{R}_{t} ; H\right)$ is $T$ periodic, $\alpha>0$ is fixed, then there is a unique $T$ periodic weak solution $\xi \in C\left(\mathbb{R}_{t} ; H\right)$ of $(3.2)$ verifying $\|\xi\|_{L^{\infty}\left(\mathbb{R}_{t} ; H\right)} \leq\left(\frac{1}{\alpha T}+4\right)\|f\|_{L^{1}(] 0, T[; H)}$.

Proof. Consider $f_{n} \in C^{1}\left(\mathbb{R}_{t} ; H\right), T$ periodic, such that $\lim _{n \rightarrow+\infty} f_{n}=f$ in $L^{1}(] 0, T[; H)$ and denote by $x_{n} \in C\left(\mathbb{R}_{t} ; D(A)\right) \cap C^{1}\left(\mathbb{R}_{t} ; H\right)$ the corresponding strong solution. By the Proposition 3.2 we have that $\left\|x_{n}-x_{m}\right\|_{L^{\infty}\left(\mathbb{R}_{t} ; H\right)} \leq\left(\frac{1}{\alpha T}+4\right) \| f_{n}-$ $f_{m} \|_{L^{1}(] 0, T[; H)}$ and thus $\left(x_{n}\right)_{n}$ converges to some $T$ periodic function $\xi \in C\left(\mathbb{R}_{t} ; H\right)$ such that $\|\xi\|_{L^{\infty}\left(\mathbb{R}_{t} ; H\right)} \leq\left(\frac{1}{\alpha T}+4\right)\|f\|_{L^{1}(] 0, T[; H)}$. If $\varphi \in C\left(\mathbb{R}_{t} ; D\left(A^{\star}\right)\right) \cap C^{1}\left(\mathbb{R}_{t} ; H\right)$ is $T$ periodic, we have for all $n$ :

$$
\int_{0}^{T}\left\langle x_{n}(t), \alpha \varphi(t)-\varphi^{\prime}(t)+A^{\star} \varphi(t)\right\rangle d t=\int_{0}^{T}\left\langle f_{n}(t), \varphi(t)\right\rangle d t
$$

By passing to the limit in respect to $n$ we deduce that $\xi$ is a $T$ periodic weak solution. In order to prove the uniqueness, consider $x_{1}, x_{2}$ two $T$ periodic weak solutions and therefore we have:

$$
\int_{0}^{T}\left\langle x_{1}(t)-x_{2}(t), \alpha \varphi(t)-\varphi^{\prime}(t)+A^{\star} \varphi(t)\right\rangle d t=0, \forall \varphi \in C\left(\mathbb{R}_{t} ; D\left(A^{\star}\right)\right) \cap C^{1}\left(\mathbb{R}_{t} ; H\right)
$$

$$
T \text { periodic. }
$$

In particular we deduce that $\int_{0}^{T}\left\langle x_{1}(t)-x_{2}(t), g(t)\right\rangle d t=0, \forall g \in C^{1}\left(\mathbb{R}_{t} ; H\right), T$ periodic (take $\varphi$ the strong $T$ periodic solution for $\alpha \varphi(t)-\varphi^{\prime}+A^{\star} \varphi(t)=g(t), t \in \mathbb{R}_{t}$ ). It follows by density that $\int_{0}^{T}\left\langle x_{1}(t)-x_{2}(t), g(t)\right\rangle d t=0, \forall g \in L_{l o c}^{1}\left(\mathbb{R}_{t} ; H\right), T$ periodic, or $x_{1}=x_{2}$. $\square$

## 4. The Maxwell equations

In this section we assume that the charge and current densities $\rho=q \int_{\mathbb{R}_{p}^{3}} f d p, j=$ $q \int_{\mathbb{R}_{p}^{3}} v(p) f d p$ are given $T$ periodic functions and we study the time periodic Maxwell equations:

$$
\begin{equation*}
\partial_{t} E-c_{0}^{2} \operatorname{rot} B=-\frac{1}{\varepsilon_{0}} j(t, x), \partial_{t} B+\operatorname{rot} E=0,(t, x) \in \mathbb{R}_{t} \times \Omega \tag{4.1}
\end{equation*}
$$

On the boundary we impose the Silver-Müller condition:

$$
\begin{equation*}
n(x) \wedge E(t, x)+c_{0} n(x) \wedge(n(x) \wedge B(t, x))=h(t, x),(t, x) \in \mathbb{R}_{t} \times \partial \Omega \tag{4.2}
\end{equation*}
$$

In order to be consistent to the perturbed Vlasov problem we consider also the perturbed Maxwell equations:

$$
\begin{equation*}
\alpha E(t, x)+\partial_{t} E-c_{0}^{2} \operatorname{rot} B=-\frac{1}{\varepsilon_{0}} j(t, x), \alpha B(t, x)+\partial_{t} B+\operatorname{rot} E=0,(t, x) \in \mathbb{R}_{t} \times \Omega \tag{4.3}
\end{equation*}
$$

where $\alpha>0$ is fixed. Let us introduce the standard Hilbert spaces $H(\operatorname{rot} ; \Omega), H(\operatorname{div} ; \Omega)$ defined by:

$$
H(\operatorname{rot} ; \Omega)=\left\{u \in L^{2}(\Omega)^{3} \mid \operatorname{rot} u \in L^{2}(\Omega)^{3}\right\}, H(\operatorname{div} ; \Omega)=\left\{u \in L^{2}(\Omega)^{3} \mid \operatorname{div} u \in L^{2}(\Omega)\right\}
$$

endowed with the norms:

$$
\left(\|u\|_{L^{2}(\Omega)^{3}}^{2}+\|\operatorname{rot} u\|_{L^{2}(\Omega)^{3}}^{2}\right)^{1 / 2}, \text { respectively }\left(\|u\|_{L^{2}(\Omega)^{3}}^{2}+\|\operatorname{div} u\|_{L^{2}(\Omega)}^{2}\right)^{1 / 2}
$$

It is well known that $C^{1}(\bar{\Omega})^{3}$ is dense in $H(\operatorname{rot} ; \Omega)$ and $H(\operatorname{div} ; \Omega)$ (see [16]). The application $\varphi \in C^{1}(\bar{\Omega})^{3} \rightarrow n \wedge \varphi \in C^{1}(\partial \Omega)^{3}$ extends by continuity to a continuous linear $\operatorname{map} n \wedge: H(\operatorname{rot} ; \Omega) \rightarrow\left(H^{1 / 2}(\partial \Omega)^{3}\right)^{\prime}=H^{-1 / 2}(\partial \Omega)^{3}$ such that:

$$
\int_{\Omega} \operatorname{rot} u \cdot \Phi d x-\int_{\Omega} u \cdot \operatorname{rot} \Phi d x=\langle n \wedge u, \varphi\rangle_{H^{-1 / 2}(\partial \Omega)^{3}, H^{1 / 2}(\partial \Omega)^{3}}
$$

for all functions $u \in H(\operatorname{rot} ; \Omega), \Phi \in H^{1}(\Omega)^{3}, \varphi=\left.\Phi\right|_{\partial \Omega} \in H^{1 / 2}(\partial \Omega)^{3}$. On the other hand, the application $\varphi \in C^{1}(\bar{\Omega})^{3} \rightarrow n \cdot \varphi \in C^{1}(\partial \Omega)$ extends by continuity to a continuous linear map $n \cdot: H(\operatorname{div} ; \Omega) \rightarrow\left(H^{1 / 2}(\partial \Omega)\right)^{\prime}=H^{-1 / 2}(\partial \Omega)$ such that:

$$
\int_{\Omega} \operatorname{div} u \Phi d x+\int_{\Omega} u \cdot \operatorname{grad} \Phi d x=\langle n \cdot u, \varphi\rangle_{H^{-1 / 2}(\partial \Omega), H^{1 / 2}(\partial \Omega)}
$$

for all functions $u \in H(\operatorname{div} ; \Omega), \Phi \in H^{1}(\Omega), \varphi=\left.\Phi\right|_{\partial \Omega} \in H^{1 / 2}(\partial \Omega)$. We note $\mathcal{H}=L^{2}(\Omega)^{6}$ endowed with the norm $\left(\|E\|_{L^{2}(\Omega)^{3}}^{2}+c_{0}^{2} \cdot\|B\|_{L^{2}(\Omega)^{3}}^{2}\right)^{1 / 2}$ and define $\mathcal{A}: D(\mathcal{A}) \subset \mathcal{H} \rightarrow \mathcal{H}$ given by:

$$
\begin{array}{r}
D(\mathcal{A})=\left\{(E, B) \in \mathcal{H} \mid \text { rot } E, \text { rot } B \in L^{2}(\Omega)^{3}, n \wedge E, n \wedge B \in L^{2}(\partial \Omega)^{3}\right. \\
\left.n \wedge E+\left.c_{0} n \wedge(n \wedge B)\right|_{\partial \Omega}=0\right\}
\end{array}
$$

and:

$$
\mathcal{A}(E, B)=\left(-c_{0}^{2} \operatorname{rot} B, \operatorname{rot} E\right), \forall(E, B) \in D(\mathcal{A})
$$

We check by direct computations that for $(E, B) \in D(\mathcal{A})$ we have:

$$
\begin{array}{r}
\langle\mathcal{A}(E, B),(E, B)\rangle=-c_{0}^{2} \int_{\Omega}\{\operatorname{rot} B \cdot E-\operatorname{rot} E \cdot B\} d x \\
=c_{0} \int_{\partial \Omega}|n \wedge E|^{2} d \sigma=c_{0}^{3} \int_{\partial \Omega}|n \wedge B|^{2} d \sigma \geq 0
\end{array}
$$

We have the following result (see [29]):
Proposition 4.1. The operator $\mathcal{A}$ is maximal monotone.
Proposition 4.2. The adjoint of the unbounded operator $\mathcal{A}$ is given by $\mathcal{A}^{\star}: D\left(\mathcal{A}^{\star}\right) \subset$ $\mathcal{H} \rightarrow \mathcal{H}$ where:

$$
\begin{array}{r}
D\left(\mathcal{A}^{\star}\right)=\left\{(E, B) \in \mathcal{H} \mid \text { rot } E, \text { rot } B \in L^{2}(\Omega)^{3}, n \wedge E\right. \\
\left.n \wedge B \in L^{2}(\partial \Omega)^{3}, n \wedge E-\left.c_{0} n \wedge(n \wedge B)\right|_{\partial \Omega}=0\right\}
\end{array}
$$

and:

$$
\mathcal{A}^{\star}(E, B)=\left(c_{0}^{2} \operatorname{rot} B,-\operatorname{rot} E\right), \forall(E, B) \in D\left(\mathcal{A}^{\star}\right) .
$$

Remark 4.3. By direct computation we check that for $(E, B) \in D\left(\mathcal{A}^{\star}\right)$ we have:

$$
\left\langle\mathcal{A}^{\star}(E, B),(E, B)\right\rangle=c_{0} \int_{\partial \Omega}|n \wedge E|^{2} d \sigma \geq 0
$$

and as before we can prove that $\mathcal{A}^{\star}$ is maximal monotone (see also [11], p.113).
We consider also the equation:

$$
\begin{equation*}
\alpha(\varphi(t), \psi(t))-\frac{d}{d t}(\varphi, \psi)+\mathcal{A}^{\star}(\varphi(t), \psi(t))=(f(t), g(t)), t \in \mathbb{R}_{t} \tag{4.4}
\end{equation*}
$$

Proposition 4.4. Assume that $(f, g) \in C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$ is $T$ periodic, $\alpha>0$ fixed. Then there is a unique $T$ periodic solution $(\varphi, \psi) \in C\left(\mathbb{R}_{t} ; D\left(\mathcal{A}^{\star}\right)\right) \cap C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$ for the equation (4.4) which verifies the estimates:

$$
\|(\varphi, \psi)\|_{L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)} \leq\left(\frac{1}{\alpha T}+4\right)\|(f, g)\|_{L^{1}(] 0, T[; \mathcal{H})}
$$

and:

$$
\begin{array}{r}
c_{0}^{1 / 2}\|n \wedge \varphi\|_{L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)}=c_{0}^{3 / 2}\|n \wedge \psi\|_{L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)} \\
\leq\left(\frac{1}{\alpha T}+4\right)^{1 / 2}\|(f, g)\|_{L^{1}(] 0, T[; \mathcal{H})} .
\end{array}
$$

Proof. The existence and uniqueness of the $T$ periodic solution as well as the first estimate follow by the Proposition 3.2, the Remarks 4.3,3.4. By the other hand observe that:

$$
\left\langle(\varphi, \psi), \mathcal{A}^{\star}(\varphi, \psi)\right\rangle=c_{0} \int_{\partial \Omega}|n \wedge \varphi(t)|^{2} d \sigma=c_{0}^{3} \int_{\partial \Omega}|n \wedge \psi(t)|^{2} d \sigma
$$

After integration on $] 0, T$ [, equation (4.4) gives:

$$
\begin{aligned}
\alpha\|(\varphi, \psi)\|_{L^{2}(] 0, T[; \mathcal{H})}^{2} & +c_{0}\|n \wedge \varphi\|_{L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)}^{2} \leq \int_{0}^{T}\|(f(t), g(t))\|_{\mathcal{H}} \cdot\|(\varphi(t), \psi(t))\|_{\mathcal{H}} d t \\
& \leq\|(f, g)\|_{L^{1}(] 0, T[; \mathcal{H})} \cdot\|(\varphi, \psi)\|_{L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)} \leq\left(\frac{1}{\alpha T}+4\right)\|(f, g)\|_{L^{1}(] 0, T[; \mathcal{H})}^{2}
\end{aligned}
$$

and thus the second estimate follows.
Remark 4.5. If $(f, g) \in L_{l o c}^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$ is $T$ periodic, then the $T$ periodic weak solution of the equation (4.4) verifies the same estimates as in the Proposition 4.4.

We can prove now the existence and uniqueness of the $T$ periodic solution for the perturbed Maxwell equations. Let us start with a result for strong solutions.
Proposition 4.6. Assume that $j \in C^{1}\left(\mathbb{R}_{t} ; L^{2}(\Omega)^{3}\right)$ is $T$ periodic and that there is $\tilde{h} \in C^{1}\left(\mathbb{R}_{t} ; H^{1}(\Omega)^{3}\right) \cap C^{2}\left(\mathbb{R}_{t} ; L^{2}(\Omega)^{3}\right) T$ periodic such that $\left.n \wedge \tilde{h}\right|_{\mathbb{R}_{t} \times \partial \Omega}=h$. Then for
$\alpha>0$ fixed there is a unique $T$ periodic solution $(E, B) \in C\left(\mathbb{R}_{t} ; H(\operatorname{rot} ; \Omega)^{2}\right) \cap C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$ for the perturbed Maxwell problem (4.3),(4.2). Moreover we have:

$$
\left.\begin{array}{rl}
\alpha \int_{0}^{T} \int_{\Omega}\left\{|E(t, x)|^{2}\right. & \left.+c_{0}^{2}|B(t, x)|^{2}\right\} d t d x+\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}\left\{|n \wedge E(t, x)|^{2}+c_{0}^{2}|n \wedge B(t, x)|^{2}\right\} d t d \sigma \\
& =-\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega} j(t, x) \cdot E(t, x) d t d x+\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma
\end{array}\right\}
$$

and:
$\int_{0}^{T}\langle(E(t), B(t)),(f(t), g(t))\rangle_{\mathcal{H}} d t=c_{0} \int_{0}^{T} \int_{\partial \Omega}(n \wedge \varphi) h d t d \sigma-\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega} j(t, x) \cdot \varphi(t, x) d t d x$,
for all $(f, g) \in L_{\text {loc }}^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$, T periodic, where $(\varphi, \psi)$ is the $T$ periodic weak solution for (4.4).

Proof. In order to prove the uniqueness consider $\left(E_{1}, B_{1}\right),\left(E_{2}, B_{2}\right)$ two periodic solutions and observe that, since $n \wedge\left(E_{1}-E_{2}\right)+\left.c_{0} n \wedge\left(n \wedge\left(B_{1}-B_{2}\right)\right)\right|_{\mathbb{R}_{t} \times \partial \Omega}=h-h=0$, then $(E, B)=\left(E_{1}-E_{2}, B_{1}-B_{2}\right) \in C\left(\mathbb{R}_{t} ; D(\mathcal{A})\right) \cap C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$ is also a $T$ periodic solution for the perturbed evolution equation $\alpha(E, B)+\frac{d}{d t}(E, B)+\mathcal{A}(E, B)=0$. After multiplication by $(E, B)$ and integration on $] 0, T[$ we deduce that $(E, B)=(0,0)$. In order to prove the existence let us take $\left(E_{1}, B_{1}\right) \in C\left(\mathbb{R}_{t} ; D(\mathcal{A})\right) \cap C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$ the unique $T$ periodic solution for:

$$
\alpha\left(E_{1}, B_{1}\right)+\frac{d}{d t}\left(E_{1}, B_{1}\right)+\mathcal{A}\left(E_{1}, B_{1}\right)=\left(-\frac{j}{\varepsilon_{0}}-\alpha \tilde{h}-\frac{d}{d t} \tilde{h},-\operatorname{rot} \tilde{h}\right) \in C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right)
$$

We verify that $(E, B)=\left(E_{1}+\tilde{h}, B_{1}\right)$ is a $T$ periodic solution for (4.3), (4.2) with $(n \wedge$ $E(t), n \wedge B(t)) \in L^{2}(\partial \Omega)^{6}$. By multiplying (4.3) by ( $E, B$ ) and by using (4.2) we obtain: $\alpha\|(E, B)\|^{2}+\frac{1}{2} \frac{d}{d t}\|(E, B)\|^{2}+c_{0} \int_{\partial \Omega}(n \wedge E) \cdot(n \wedge E-h) d \sigma=-\frac{1}{\varepsilon_{0}} \int_{\Omega} j(t, x) \cdot E(t, x) d x$.
Using again (4.2) we find that $(n \wedge E) \cdot(n \wedge E-h)=\frac{1}{2}\left(|n \wedge E|^{2}+c_{0}^{2}|n \wedge B|^{2}-|h|^{2}\right)$ and therefore:

$$
\begin{align*}
\alpha \int_{\Omega}\left\{|E(t, x)|^{2}\right. & \left.+c_{0}^{2}|B(t, x)|^{2}\right\} d x+\frac{1}{2} \frac{d}{d t} \int_{\Omega}\left\{|E(t, x)|^{2}+c_{0}^{2}|B(t, x)|^{2}\right\} d x \\
& +\frac{c_{0}}{2} \int_{\partial \Omega}\left\{|n \wedge E(t, x)|^{2}+c_{0}^{2}|n \wedge B(t, x)|^{2}\right\} d \sigma \\
& =-\frac{1}{\varepsilon_{0}} \int_{\Omega} j(t, x) \cdot E(t, x) d x+\frac{c_{0}}{2} \int_{\partial \Omega}|h(t, x)|^{2} d \sigma \tag{4.8}
\end{align*}
$$

Finally, after integration on $] 0, T$ [ we deduce:
$\alpha \int_{0}^{T} \int_{\Omega}\left\{|E(t, x)|^{2}+c_{0}^{2}|B(t, x)|^{2}\right\} d t d x+\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}\left\{|n \wedge E(t, x)|^{2}+c_{0}^{2}|n \wedge B(t, x)|^{2}\right\} d t d \sigma$ $=-\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega} j(t, x) \cdot E(t, x) d t d x+\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma$.

For $(f, g) \in C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right), T$ periodic, consider $(\varphi, \psi)$ the unique $T$ periodic solution of (4.4). After multiplication by $(\varphi, \psi)$ of the perturbed Maxwell equations and integration on $] 0, T$ [ we find that:

$$
\begin{aligned}
\int_{0}^{T}\langle(E(t), B(t)),(f(t), g(t))\rangle_{\mathcal{H}} d t= & \int_{0}^{T}\langle(E(t), B(t)), \alpha(\varphi(t), \psi(t)) \\
& \left.-\frac{d}{d t}(\varphi, \psi)+\mathcal{A}^{\star}(\varphi(t), \psi(t))\right\rangle_{\mathcal{H}} d t \\
= & c_{0} \cdot \int_{0}^{T} \int_{\partial \Omega}(n \wedge \varphi) \cdot h(t, x) d t d \sigma \\
& -\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega} j(t, x) \cdot \varphi(t, x) d t d x
\end{aligned}
$$

By using Proposition 3.6 and Remark 4.5 we verify easily that the previous equality holds for $(f, g) \in L_{l o c}^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right), T$ periodic, with $(\varphi, \psi)$ the associated $T$ periodic weak solution. By the Proposition 4.4 we deduce that for all $(f, g) \in L_{l o c}^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right), T$ periodic, we have:

$$
\begin{aligned}
& \quad\left|\int_{0}^{T}\langle(E(t), B(t)),(f(t), g(t))\rangle_{\mathcal{H}} d t\right| \\
& \quad \leq c_{0}^{1 / 2}\left(\frac{1}{\alpha T}+4\right)^{1 / 2}\|h\|_{L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)}\|(f, g)\|_{L^{1}(] 0, T[; \mathcal{H})} \\
& +\frac{1}{\varepsilon_{0}}\left(\frac{1}{\alpha T}+4\right)\|j\|_{L^{1}(] 0, T\left[; L^{2}(\Omega)^{3}\right)}\|(f, g)\|_{L^{1}(] 0, T[; \mathcal{H})}
\end{aligned}
$$

and thus the estimate (4.6) of our proposition follows.
REmARK 4.7. In particular the solution constructed above verifies:

$$
\begin{aligned}
& \int_{0}^{T} \int_{\Omega}\left\{E(t, x) \cdot\left(\alpha \varphi-\partial_{t} \varphi\right)-c_{0}^{2} B(t, x) \cdot \operatorname{rot} \varphi\right\} d t d x-c_{0}^{2} \int_{0}^{T} \int_{\partial \Omega}(n \wedge B) \cdot \varphi d t d \sigma \\
&=-\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega} j \cdot \varphi d t d x
\end{aligned}
$$

and:

$$
\int_{0}^{T} \int_{\Omega}\left\{B(t, x) \cdot\left(\alpha \psi-\partial_{t} \psi\right)+E(t, x) \cdot \operatorname{rot} \psi\right\} d t d x+\int_{0}^{T} \int_{\partial \Omega}(n \wedge E) \cdot \psi d t d \sigma=0
$$

for all $\varphi, \psi \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right)^{3}, T$ periodic.
Definition 4.8. Assume that $j \in L_{\text {loc }}^{1}\left(\mathbb{R}_{t} ; L^{2}(\Omega)^{3}\right)$ and $h \in L_{\text {loc }}^{2}\left(\mathbb{R}_{t} ; L^{2}(\partial \Omega)^{3}\right)$ are $T$ periodic, $\left.(n \cdot h)\right|_{\mathbb{R}_{t} \times \partial \Omega}=0, \alpha>0$ fixed. We say that $(E, B) \in C\left(\mathbb{R}_{t} ; \mathcal{H}\right)$ is a $T$ periodic weak solution for the perturbed Maxwell problem (4.3), (4.2) iff $(E, B)$ is $T$ periodic and:

$$
\begin{aligned}
& \int_{0}^{T}\left\langle(E(t), B(t)), \alpha(\varphi(t), \psi(t))-\frac{d}{d t}(\varphi, \psi)+\mathcal{A}^{\star}(\varphi(t), \psi(t))\right\rangle_{\mathcal{H}} d t \\
& =c_{0} \int_{0}^{T} \int_{\partial \Omega}(n \wedge \varphi) \cdot h(t, x) d t d \sigma \\
& -\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega} j(t, x) \cdot \varphi(t, x) d t d x
\end{aligned}
$$

for all $(\varphi, \psi) \in C\left(\mathbb{R}_{t} ; D\left(\mathcal{A}^{\star}\right)\right) \cap C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$, $T$ periodic.
Proposition 4.9. Assume that $j \in L_{\text {loc }}^{1}\left(\mathbb{R}_{t} ; L^{2}(\Omega)^{3}\right)$ and $h \in L_{l o c}^{2}\left(\mathbb{R}_{t} ; L^{2}(\partial \Omega)^{3}\right)$ are $T$ periodic, $\left.(n \cdot h)\right|_{\mathbb{R}_{t} \times \partial \Omega}=0$. Then, for $\alpha>0$ fixed there is a unique $T$ periodic weak solution for the perturbed Maxwell problem (4.3),(4.2) verifying the equalities (4.5),(4.7) and the estimate (4.6).

Proof. Consider $j_{k} \in C^{1}\left(\mathbb{R}_{t} ; L^{2}(\Omega)^{3}\right)$ and $\tilde{h}_{k} \in C^{1}\left(\mathbb{R}_{t} ; H^{1}(\Omega)^{3}\right) \cap C^{2}\left(\mathbb{R}_{t} ; L^{2}(\Omega)^{3}\right) T$ periodic such that $j_{k} \rightarrow j$ in $L^{1}(] 0, T\left[; L^{2}(\Omega)^{3}\right)$ and $n \wedge \tilde{h}_{k} \rightarrow h$ in $L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)$. Denote by $\left(E_{k}, B_{k}\right)$ the corresponding $T$ periodic strong solutions. By (4.6), (4.5) we deduce that $\left(E_{k}, B_{k}\right) \rightarrow(E, B)$ in $C\left(\mathbb{R}_{t} ; \mathcal{H}\right)$ and $\left(n \wedge E_{k}, n \wedge B_{k}\right)$ converges in $L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)$ to some function denoted $(n \wedge E, n \wedge B)$. Finally, by passing to the limit for $k \rightarrow+\infty$ we deduce that $(E, B)$ verifies (4.5),(4.7), (4.6). Observe also that by the Remark 4.7 we have:

$$
\begin{aligned}
& \int_{0}^{T} \int_{\Omega}\left\{E_{k}(t, x) \cdot\left(\alpha \varphi-\partial_{t} \varphi\right)-c_{0}^{2} B_{k}(t, x) \cdot \operatorname{rot} \varphi\right\} d t d x-c_{0}^{2} \int_{0}^{T} \int_{\partial \Omega}\left(n \wedge B_{k}\right) \cdot \varphi d t d \sigma \\
&=-\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega} j_{k} \cdot \varphi d t d x
\end{aligned}
$$

for all $\varphi \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right)^{3}, T$ periodic, and by passing to the limit for $k \rightarrow+\infty$ we deduce that:

$$
\begin{aligned}
& \int_{0}^{T} \int_{\Omega}\left\{E(t, x) \cdot\left(\alpha \varphi-\partial_{t} \varphi\right)-c_{0}^{2} B(t, x) \cdot \operatorname{rot} \varphi\right\} d t d x-c_{0}^{2} \int_{0}^{T} \int_{\partial \Omega}(n \wedge B) \cdot \varphi d t d \sigma \\
&=-\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega} j \cdot \varphi d t d x
\end{aligned}
$$

Similarly we obtain that:

$$
\int_{0}^{T} \int_{\Omega}\left\{B(t, x) \cdot\left(\alpha \psi-\partial_{t} \psi\right)+E(t, x) \cdot \operatorname{rot} \psi\right\} d t d x+\int_{0}^{T} \int_{\partial \Omega}(n \wedge E) \cdot \psi d t d \sigma=0
$$

for all $\psi \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right)^{3}$. In order to prove the uniqueness, consider $\left(E_{1}, B_{1}\right),\left(E_{2}, B_{2}\right)$ two $T$ periodic weak solutions and observe that $\left(E_{1}-E_{2}, B_{1}-B_{2}\right)$ is a $T$ periodic weak solution corresponding to $j=0, h=0$. Thus, by Proposition 3.6 we obtain that $E_{1}-E_{2}=0, B_{1}-B_{2}=0$.

## 5. The perturbed Vlasov-Maxwell system

We study now the full perturbed Vlasov-Maxwell system (2.3), (2.2), (4.3), (4.2). We prove the existence of a $T$ periodic solution by using the Schauder fixed point theorem. Let us start by the relativistic case.
5.1. The relativistic case. In this case we have $|j(t, x)| \leq|q|$. $\int_{\mathbb{R}_{p}^{3}}|v(p)| f(t, x, p) d p \leq c_{0}|\rho(t, x)|,(t, x) \in \mathbb{R}_{t} \times \Omega$. We consider the set:

$$
\mathcal{X}=\left\{(E, B) \in L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right) \mid(E, B)(\cdot+T)=(E, B)(\cdot)\right\}
$$

and define the application $\mathcal{F}: \mathcal{X} \rightarrow \mathcal{X}, \mathcal{F}(E, B)=(\tilde{E}, \tilde{B})$ by:

$$
(E, B) \rightarrow\left(E_{\varepsilon}, B_{\varepsilon}\right) \rightarrow f \rightarrow j=q \int_{\mathbb{R}_{p}^{3}} v(p) f(t, x, p) d p \rightarrow j_{\varepsilon} \rightarrow(\tilde{E}, \tilde{B})
$$

where:
(i) $\left(E_{\varepsilon}, B_{\varepsilon}\right)$ is the convolution of $(\bar{E}, \bar{B})$ (the extension of $(E, B)$ by 0 outside $\Omega$ ) by $\zeta_{\varepsilon}(t, x):$

$$
\begin{aligned}
&\left(E_{\varepsilon}, B_{\varepsilon}\right)(t, x)=\left((\bar{E}, \bar{B}) \star \zeta_{\varepsilon}\right)(t, x) \\
& \quad=\int_{0}^{T} \int_{\Omega}(E(s, y), B(s, y)) \zeta_{\varepsilon}(t-s, x-y) d s d y,(t, x) \in \mathbb{R}_{t} \times \Omega
\end{aligned}
$$

with
$\zeta_{\varepsilon}(t, x)=\zeta_{\varepsilon_{1}, \varepsilon_{2}}(t, x)=\zeta_{1, \varepsilon_{1}}(t) \cdot \zeta_{2, \varepsilon_{2}}(x)=\left[\frac{1}{\varepsilon_{1}} \sum_{k \in Z} \zeta_{3}\left(\frac{t-k T}{\varepsilon_{1}}\right)\right] \frac{1}{\varepsilon_{2}^{3}} \zeta_{2}\left(\frac{x}{\varepsilon_{2}}\right)$,
$\zeta_{2} \in C_{c}^{\infty}\left(\mathbb{R}^{3}\right), \zeta_{3} \in C_{c}^{\infty}(\mathbb{R}), \zeta_{2}, \zeta_{3} \geq 0, \operatorname{supp} \zeta_{2} \subset B(0,1), \operatorname{supp} \zeta_{3} \subset[-1,1], \int_{\mathbb{R}^{3}} \zeta_{2}(u) d u=$ $\int_{\mathbb{R}} \zeta_{3}(u) d u=1$ (note that $\left(E_{\varepsilon}, B_{\varepsilon}\right)$ is also $T$ periodic);
(ii) $f$ is the $T$ periodic mild solution of (2.3), (2.2) corresponding to the regular force $F_{\varepsilon}(t, x, p)=q\left(E_{\varepsilon}(t, x)+v(p) \wedge B_{\varepsilon}(t, x)\right)$ (cf. Proposition 2.4);
(iii) $j_{\varepsilon}$ is the convolution of $\bar{j}$ (the extension of $j$ by 0 outside $\Omega$ ) by $\zeta_{\varepsilon}^{\sqrt{ }}(t, x)=$ $\zeta_{\varepsilon}(-t,-x)$ :

$$
j_{\varepsilon}(t, x)=\left(\bar{j} \star \zeta_{\varepsilon}^{\sqrt{ }}\right)(t, x)=\int_{0}^{T} \int_{\Omega} j(s, y) \zeta_{\varepsilon}(s-t, y-x) d s d y,(t, x) \in \mathbb{R}_{t} \times \Omega
$$

(iv) $(\tilde{E}, \tilde{B})$ is the $T$ periodic weak solution of (4.3),(4.2) (see Proposition 4.9) associated to the current density $j_{\varepsilon}$.

Let us consider $\quad M_{\alpha, \varepsilon}:=c_{0}^{1 / 2}\left(\frac{1}{\alpha T}+4\right)^{1 / 2}\|h\|_{L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)}+\left(\frac{1}{\alpha T}+4\right) \frac{\left\|\zeta_{2}\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}}{\varepsilon_{0} \cdot \varepsilon_{2}^{3 / 2}}$. $\frac{c_{0}|q|}{\alpha} \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g d t d \sigma d p \quad$ and $\quad \mathcal{C}=\left\{(E, B) \in \mathcal{X} \mid\|(E, B)\|_{L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)} \leq M_{\alpha, \varepsilon}\right\}$ which is a convex compact subset in respect to the weak $\star$ topology of $L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$. In order to apply the Schauder theorem, we need the following two propositions. We postpone the details of proofs to the end of this section.
Proposition 5.1. Assume that $g, h$ are $T$ periodic such that $g \geq 0,\left.(n \cdot h)\right|_{\mathbb{R}_{t} \times \partial \Omega}=0$ and

$$
\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(t, x, p) d t d \sigma d p+\int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma<+\infty
$$

Then $\mathcal{F}(\mathcal{X}) \subset \mathcal{C}$.
Proposition 5.2. Assume that $g \in L^{\infty}\left(\mathbb{R}_{t} \times \Sigma^{-}\right)$, h are $T$ periodic, with $g \geq 0$, $(n$. $h)\left.\right|_{\mathbb{R}_{t} \times \partial \Omega}=0$ and

$$
\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| \mathcal{E}(p) g(t, x, p) d t d \sigma d p+\int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma<+\infty
$$

Then the application $\mathcal{F}$ is continuous in respect to the weak $\star$ topology of $L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$.

The main result of this section is given in the following theorem.
THEOREM 5.3. Assume that $g \in L^{\infty}\left(\mathbb{R}_{t} \times \Sigma^{-}\right)$,h are $T$ periodic, $g \geq 0$, $\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| \mathcal{E}(p) g d t d \sigma d p+\int_{0}^{T} \int_{\partial \Omega}|h|^{2} d t d \sigma<+\infty \quad$ and $\quad$ there $\quad$ is $\quad \tilde{h} \in$ $C^{1}\left(\mathbb{R}_{t} ; H^{1}(\Omega)^{3}\right) \cap C^{2}\left(\mathbb{R}_{t} ; L^{2}(\Omega)^{3}\right) \quad T$ periodic such that $\left.n \wedge \tilde{h}\right|_{\mathbb{R}_{t} \times \partial \Omega}=h$. Then,
for every $\alpha, \varepsilon_{1}, \varepsilon_{2}>0$ there is a $T$ periodic solution for the perturbed relativistic Vlasov-Maxwell system:

$$
\begin{array}{r}
\alpha f+\partial_{t} f+v(p) \cdot \nabla_{x} f+q\left(\left(\bar{E} \star \zeta_{\varepsilon}\right)+v(p) \wedge\left(\bar{B} \star \zeta_{\varepsilon}\right)\right) \cdot \nabla_{p} f \\
=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3} \\
\alpha E+\partial_{t} E-c_{0}^{2} \operatorname{rot} B=-\frac{\bar{j} \star \zeta_{\varepsilon}^{\vee}}{\varepsilon_{0}}, \alpha B+\partial_{t} B+\operatorname{rot} E=0,(t, x) \in \mathbb{R}_{t} \times \Omega  \tag{5.1}\\
f(t, x, p)=g(t, x, p),(t, x, p) \in \mathbb{R}_{t} \times \Sigma^{-}, n \wedge E(t, x)+c_{0} n \wedge(n \wedge B(t, x)) \\
=h(t, x),(t, x) \in \mathbb{R}_{t} \times \partial \Omega
\end{array}
$$

with $j=q \int_{\mathbb{R}_{p}^{3}} f(t, x, p) v(p) d p$.
Moreover $\|(E, B)\|_{L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)} \leq M_{\alpha, \varepsilon},(E, B) \in C\left(\mathbb{R}_{t} ; H(\text { rot } ; \Omega)^{2}\right) \cap C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$ and:
$\alpha \int_{0}^{T} \int_{\Omega}\left\{|E(t, x)|^{2}+c_{0}^{2}|B(t, x)|^{2}\right\} d t d x+\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}\left\{|n \wedge E(t, x)|^{2}+c_{0}^{2}|n \wedge B(t, x)|^{2}\right\} d t d \sigma$

$$
\begin{equation*}
=-\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega}\left(\bar{j} \star \zeta_{\varepsilon}^{\vee}\right)(t, x) \cdot E(t, x) d t d x+\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma \tag{5.2}
\end{equation*}
$$

Proof. By the Propositions 5.1,5.2, the fixed point theorem of Schauder applies and thus we deduce the existence of a $T$ periodic solution for the perturbed VlasovMaxwell system. The other statements follow by the Proposition 4.6.

The following proposition establishes an a priori estimate for boundary terms.
Proposition 5.4. Under the hypothesis of Theorem 5.3, consider $(f, E, B)$ the $T$ periodic solution constructed above. Then we have:

$$
\begin{align*}
& \alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p)(1+\mathcal{E}(p)) d t d x d p+\alpha \int_{0}^{T} \int_{\Omega}\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d t d x \\
& +\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f(1+\mathcal{E}(p)) d t d \sigma d p+\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}|n \wedge E|^{2}+\frac{1}{\mu_{0}}|n \wedge B|^{2}\right\} d t d \sigma \\
& =\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(1+\mathcal{E}(p)) d t d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma \tag{5.3}
\end{align*}
$$

Proof. Consider the test function $\varphi(t, x, p)=\mathcal{E}(p) \chi_{R}(|p|)$. By using the Green formula (2.9) we have:

$$
\begin{aligned}
& \int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f \mathcal{E}(p) \chi_{R}(|p|) d t d \sigma d p+\int_{0}^{T} \int_{\Sigma^{-}}(v(p) \cdot n(x)) g \mathcal{E}(p) \chi_{R}(|p|) d t d \sigma d p \\
& +\alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p) \mathcal{E}(p) \chi_{R}(|p|) d t d x d p \\
& =\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} q f\left(\left(\bar{E} \star \zeta_{\varepsilon}\right)+v(p) \wedge\left(\bar{B} \star \zeta_{\varepsilon}\right)\right) \cdot\left(v(p) \chi_{R}(|p|)+\mathcal{E}(p) \chi^{\prime}\left(\frac{|p|}{R}\right) \cdot \frac{1}{R} \cdot \frac{p}{|p|}\right) d t d x d p
\end{aligned}
$$

Observe that we can pass to the limit for $R \rightarrow+\infty$ in the last integral since $\bar{E} \star \zeta_{\varepsilon} \in$ $L^{\infty}\left(\mathbb{R}_{t} \times \Omega\right)$ and

$$
\begin{aligned}
f \cdot\left|v(p) \chi_{R}(|p|)+\mathcal{E}(p) \chi^{\prime}\left(\frac{|p|}{R}\right) \cdot \frac{1}{R} \cdot \frac{p}{|p|}\right| & \leq f \cdot\left(|v(p)|+2 \frac{\mathcal{E}(p)}{|p|} \cdot\left\|\chi^{\prime}\right\|_{\infty}\right) \\
& \leq f \cdot|v(p)| \cdot\left(1+2 \cdot\left\|\chi^{\prime}\right\|_{\infty}\right) \\
& \leq f \cdot c_{0} \cdot\left(1+2 \cdot\left\|\chi^{\prime}\right\|_{\infty}\right) \in L^{1}(] 0, T\left[\times \Omega \times \mathbb{R}_{p}^{3}\right)
\end{aligned}
$$

By passing to the limit for $R \rightarrow+\infty$ and by using the monotone convergence theorem we deduce that:

$$
\begin{aligned}
\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f \mathcal{E}(p) d t d \sigma d p+\int_{0}^{T} \int_{\Sigma^{-}} & (v(p) \cdot n(x)) g \mathcal{E}(p) d t d \sigma d p \\
& +\alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f \mathcal{E}(p) d t d x d p \\
= & \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} q f\left(\bar{E} \star \zeta_{\varepsilon}\right) v(p) d t d x d p \\
= & \int_{0}^{T} \int_{\Omega} E(t, x) \cdot\left(\bar{j} \star \zeta_{\varepsilon}^{\sqrt{ }}\right)(t, x) d t d x
\end{aligned}
$$

Similarly, by using the test function $\varphi(t, x, p)=\chi_{R}(|p|)$ and by letting $R \rightarrow+\infty$ we obtain that:
$\alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f d t d x d p+\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f d t d \sigma d p=\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g d t d \sigma d p$.
By combining with (5.2), finally one gets (5.3).
REMARK 5.5. In particular the above computations imply

$$
\alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f d t d x d p \leq \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g d t d \sigma d p,
$$

and

$$
\alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f d t d x d p \leq \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| \mathcal{E}(p) g d t d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h|^{2} d t d \sigma
$$

Proposition 5.6. Under the hypothesis of Theorem 5.3 we have for a.e. $t \in \mathbb{R}_{t}$ :

$$
\begin{aligned}
& \alpha \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}(1+\mathcal{E}(p)) f(t, x, p) d x d p+\alpha \int_{\Omega}\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d x \\
& +\frac{d}{d t} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p)(1+\mathcal{E}(p)) d x d p+\frac{1}{2} \frac{d}{d t} \int_{\Omega}\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d x \\
& \quad+\int_{\Sigma^{+}}(v(p) \cdot n(x))(1+\mathcal{E}(p)) \gamma^{+} f(t, x, p) d \sigma d p \\
& \quad+\frac{c_{0}}{2} \int_{\partial \Omega}\left\{\varepsilon_{0}|n \wedge E(t, x)|^{2}+\frac{1}{\mu_{0}}|n \wedge B(t, x)|^{2}\right\} d \sigma \\
& =\int_{\Sigma^{-}}|(v(p) \cdot n(x))|(1+\mathcal{E}(p)) g(t, x, p) d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{\partial \Omega}|h(t, x)|^{2} d \sigma .
\end{aligned}
$$

Proof. By using the Green formula (2.9) with the test function $\varphi(t, x, p)=\theta(t)$. $\mathcal{E}(p) \cdot \chi_{R}(|p|), \theta \in C_{c}^{1}(] 0, T[)$ and by letting $R \rightarrow+\infty$ we obtain:

$$
\begin{aligned}
\int_{0}^{T} \theta(t) & \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \alpha f \mathcal{E}(p) d x d p d t+\int_{0}^{T} \theta(t) \int_{\Sigma}(v(p) \cdot n(x)) \mathcal{E}(p) \gamma f d \sigma d p d t \\
= & \int_{0}^{T} \theta^{\prime}(t) \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f \mathcal{E}(p) d x d p d t+\int_{0}^{T} \theta(t) \int_{\Omega} E(t, x) \cdot\left(\bar{j} \star \zeta_{\varepsilon}^{\sqrt{ }}\right)(t, x) d x d t,(5.4)
\end{aligned}
$$

or:

$$
\begin{array}{rl}
\alpha \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} & f(t, x, p) \mathcal{E}(p) d x d p+\frac{d}{d t} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p) \mathcal{E}(p) d x d p \\
& +\int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f(t, x, p) \mathcal{E}(p) d \sigma d p \\
= & -\int_{\Sigma^{-}}(v(p) \cdot n(x)) g(t, x, p) \mathcal{E}(p) d \sigma d p+\int_{\Omega} E(t, x) \cdot\left(\bar{j} \star \zeta_{\mathcal{E}}^{\sqrt{2}}\right)(t, x) d x, \text { a.e. } t \in \mathbb{R}_{t} . \tag{5.5}
\end{array}
$$

By combining (4.8) (written for the current density $\bar{j} \star \zeta_{\varepsilon}^{\vee}$ ) with (5.5) one gets

$$
\begin{array}{rl}
\alpha \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} & \mathcal{E}(p) f(t, x, p) d x d p+\alpha \int_{\Omega}\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d x \\
& +\frac{d}{d t} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p) \mathcal{E}(p) d x d p+\frac{1}{2} \frac{d}{d t} \int_{\Omega}\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d x \\
& +\int_{\Sigma^{+}}(v(p) \cdot n(x)) \mathcal{E}(p) \gamma^{+} f(t, x, p) d \sigma d p+\frac{c_{0}}{2} \int_{\partial \Omega}\left\{\varepsilon_{0}|n \wedge E(t, x)|^{2}\right. \\
& \left.+\frac{1}{\mu_{0}}|n \wedge B(t, x)|^{2}\right\} d \sigma \\
= & \int_{\Sigma^{-}}|(v(p) \cdot n(x))| \mathcal{E}(p) g(t, x, p) d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{\partial \Omega}|h(t, x)|^{2} d \sigma \tag{5.6}
\end{array}
$$

Similarly we obtain for a.e. $t \in \mathbb{R}_{t}$ :

$$
\begin{equation*}
\alpha \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p) d x d p+\frac{d}{d t} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p) d x d p+\int_{\Sigma}(v(p) \cdot n(x)) \gamma f(t, x, p) d \sigma d p=0 . \tag{5.7}
\end{equation*}
$$

The conclusion follows by using the mass and energy balances (5.7), (5.6).
Proposition 5.7. Under the hypothesis of the Theorem 5.3 we have $\alpha \rho+\partial_{t} \rho+\operatorname{div} j=$ 0 , in $\mathcal{D}^{\prime}(] 0, T[\times \Omega)$, div $E=\left(\bar{\rho} \star \zeta_{\varepsilon}^{\sqrt{ }}\right) / \varepsilon_{0}$, div $B=0$.

Proof. Since $\alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f d t d x d p+\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f d t d \sigma d p=\int_{0}^{T} \int_{\Sigma^{-}} \mid(v(p)$. $n(x)) \mid g d t d \sigma d p$, we can consider as test function in the weak formulation all function $\varphi \in C_{c}^{1}(] 0, T[\times \Omega)$. We have:

$$
\left\langle\alpha \rho+\partial_{t} \rho+\operatorname{div} j, \varphi\right\rangle=\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f\left(\alpha \varphi(t, x)-\partial_{t} \varphi-v(p) \cdot \nabla_{x} \varphi\right) d t d x d p=0
$$

or $\alpha \rho+\partial_{t} \rho+\operatorname{div} j=0$ in $\mathcal{D}^{\prime}(] 0, T[\times \Omega)$ (in fact the above equality holds for all function $\varphi \in C_{c}^{1}([0, T] \times \Omega), T$ periodic $)$. We verify easily that we have also $\alpha\left(\bar{\rho} \star \zeta_{\varepsilon}^{\vee}\right)+\partial_{t}(\bar{\rho} \star$ $\left.\zeta_{\varepsilon}^{\sqrt{ }}\right)+\operatorname{div}\left(\bar{j} \star \zeta_{\varepsilon}^{\sqrt{ }}\right)=0$ in $\mathcal{D}^{\prime}(] 0, T[\times \Omega)$. Now, by taking the divergence of the perturbed Maxwell equations we have:

$$
\alpha \operatorname{div} E+\partial_{t} \operatorname{div} E=-\frac{1}{\varepsilon_{0}} \operatorname{div}\left(\bar{j} \star \zeta_{\varepsilon}^{\sqrt{ }}\right)=\frac{\alpha}{\varepsilon_{0}}\left(\bar{\rho} \star \zeta_{\varepsilon}^{\sqrt{\prime}}\right)+\frac{1}{\varepsilon_{0}} \partial_{t}\left(\bar{\rho} \star \zeta_{\varepsilon}^{\sqrt{ }}\right),
$$

and thus:

$$
\alpha\left(\operatorname{div} E-\frac{1}{\varepsilon_{0}}\left(\bar{\rho} \star \zeta_{\varepsilon}^{\vee}\right)\right)+\partial_{t}\left(\operatorname{div} E-\frac{1}{\varepsilon_{0}}\left(\bar{\rho} \star \zeta_{\varepsilon}^{\vee}\right)\right)=0
$$

and we deduce by periodicity that $\operatorname{div} E=\frac{1}{\varepsilon_{0}}\left(\bar{\rho} \star \zeta_{\varepsilon}^{\vee}\right)$. In the same manner we have $\alpha \operatorname{div} B+\partial_{t} \operatorname{div} B=0$ which implies that $\operatorname{div} B=0$.

We detail here the proofs of Propositions 5.1, 5.2.
Proof. (Proposition 5.1) By the estimate (4.6) we have:

$$
\begin{aligned}
&\|(\tilde{E}, \tilde{B})\|_{L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)} \leq c_{0}^{1 / 2}\left(\frac{1}{\alpha T}+4\right)^{1 / 2}\|h\|_{L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)} \\
&+\left(\frac{1}{\alpha T}+4\right) \frac{1}{\varepsilon_{0}}\left\|j_{\varepsilon}\right\|_{L^{1}(] 0, T\left[; L^{2}(\Omega)^{3}\right)}
\end{aligned}
$$

On the other hand we have:

$$
\left\|j_{\varepsilon}\right\|_{L^{1}(] 0, T\left[; L^{2}(\Omega)^{3}\right)} \leq \int_{0}^{T}\|j(t)\|_{L^{1}} \cdot\left\|\zeta_{2, \varepsilon}^{\sqrt{2}}\right\|_{L^{2}} d t=\frac{1}{\varepsilon_{2}^{3 / 2}}\left\|\zeta_{2}\right\|_{L^{2}}\|j\|_{L^{1}(] 0, T\left[; L^{1}(\Omega)^{3}\right)}
$$

But from the Proposition 2.9 we deduce that:

$$
\begin{array}{r}
\|j\|_{L^{1}\left(j 0, T\left[; L^{1}(\Omega)^{3}\right)\right.} \leq|q| \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}|v(p)| f(t, x, p) d t d x d p \\
\leq \frac{c_{0}|q|}{\alpha} \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(t, x, p) d t d \sigma d p
\end{array}
$$

and therefore $\|(\tilde{E}, \tilde{B})\|_{L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)} \leq M_{\alpha, \varepsilon}$.
Proof. (Proposition 5.2) Consider $\left(E_{k}, B_{k}\right) \rightharpoonup(E, B)$ weakly $\star$ in $L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$. Denote by $f_{k}, f$ the $T$ periodic mild solutions of (2.3) corresponding to the regularized forces $F_{k, \varepsilon}=q\left(\bar{E}_{k} \star \zeta_{\varepsilon}+v(p) \wedge\left(\bar{B}_{k} \star \zeta_{\varepsilon}\right)\right)$ :
$\alpha f_{k}+\partial_{t} f_{k}+v(p) \cdot \nabla_{x} f_{k}+q\left(\bar{E}_{k} \star \zeta_{\varepsilon}+v(p) \wedge\left(\bar{B}_{k} \star \zeta_{\varepsilon}\right)\right) \cdot \nabla_{p} f_{k}=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}$,
respectively $F_{\varepsilon}=q\left(\bar{E} \star \zeta_{\varepsilon}+v(p) \wedge\left(\bar{B} \star \zeta_{\varepsilon}\right)\right)$ :

$$
\alpha f+\partial_{t} f+v(p) \cdot \nabla_{x} f+q\left(\bar{E} \star \zeta_{\varepsilon}+v(p) \wedge\left(\bar{B} \star \zeta_{\varepsilon}\right)\right) \cdot \nabla_{p} f=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}
$$

with the boundary conditions (2.2). Since $\left(E_{k}, B_{k}\right)$ converges weakly $\star$ in $L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$, we have the pointwise convergence $\left(\bar{E}_{k} \star \zeta_{\varepsilon}, \bar{B}_{k} \star \zeta_{\varepsilon}\right)(t, x) \rightarrow\left(\bar{E} \star \zeta_{\varepsilon}, \bar{B} \star \zeta_{\varepsilon}\right)(t, x),(t, x) \in$ $\mathbb{R}_{t} \times \Omega$, as $k \rightarrow+\infty$. Moreover, $\left(\bar{E}_{k} \star \zeta_{\varepsilon}, \bar{B}_{k} \star \zeta_{\varepsilon}\right)$ is bounded in $L^{\infty}\left(\mathbb{R}_{t} \times \Omega\right)^{6}$ and by the dominated convergence theorem we deduce that $\left(\bar{E}_{k} \star \zeta_{\varepsilon}, \bar{B}_{k} \star \zeta_{\varepsilon}\right) \rightarrow\left(\bar{E} \star \zeta_{\varepsilon}, \bar{B} \star \zeta_{\varepsilon}\right)$ in
$L^{2}(] 0, T[\times \Omega)$ as $k$ goes to $+\infty$. Since $\left(f_{k}\right)_{k}$ is bounded in $L^{\infty}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right),\left\|f_{k}\right\|_{\infty} \leq$ $\|g\|_{\infty}$, we can suppose (after extraction of a subsequence) that $f_{k} \rightharpoonup \tilde{f}$ weakly $\star$ in $L^{\infty}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right)$. In order to prove that $\tilde{f}$ is the $T$ periodic weak solution of (2.3),(2.2) corresponding to the field $\left(\bar{E} \star \zeta_{\varepsilon}, \bar{B} \star \zeta_{\varepsilon}\right)$, take $\varphi \in \mathcal{T}_{w}$ and observe that:

$$
\begin{gathered}
\int_{\mathbb{R}_{p}^{3}} f_{k}(t, x, p) \nabla_{p} \varphi d p \rightharpoonup \int_{\mathbb{R}_{p}^{3}} \tilde{f}(t, x, p) \nabla_{p} \varphi d p, \text { weakly in } L^{2}(] 0, T[\times \Omega), \\
\int_{\mathbb{R}_{p}^{3}} f_{k}(t, x, p)\left(\nabla_{p} \varphi \wedge v(p)\right) d p \rightharpoonup \int_{\mathbb{R}_{p}^{3}} \tilde{f}(t, x, p)\left(\nabla_{p} \varphi \wedge v(p)\right) d p, \text { weakly in } L^{2}(] 0, T[\times \Omega) .
\end{gathered}
$$

By combining with the strong convergence of $\left(\bar{E}_{k} \star \zeta_{\varepsilon}, \bar{B}_{k} \star \zeta_{\varepsilon}\right)$ it is easy to prove that $\tilde{f}$ is the $T$ periodic weak solution for (2.3), (2.2) associated to the field ( $\bar{E} \star \zeta_{\varepsilon}, \bar{B} \star \zeta_{\varepsilon}$ ). In fact, by the uniqueness of the $T$ periodic weak solution (see Proposition 2.8) we deduce that $\tilde{f}$ is the $T$ periodic mild solution $\tilde{f}=f$. Moreover we can prove that all the sequence $\left(f_{k}\right)_{k}$ converges to $f$ weakly $\star$ in $L^{\infty}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right)$. Now we want to pass to the limit in the perturbed Maxwell equations. We need to establish the convergence for the regularized current densities $\left(\bar{j}_{k} \star \zeta_{\varepsilon}^{\sqrt{\prime}}\right)_{k}$. By using the Green formula (2.9) with the test function $\varphi(t, x, p)=|p| \cdot \chi_{R}(|p|)$ we find that:

$$
\begin{aligned}
& \alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f_{k}(t, x, p)|p| \chi_{R}(|p|) d t d x d p \leq \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(t, x, p)|p| \chi_{R}(|p|) d t d \sigma d p \\
& +\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} q f_{k}(t, x, p)\left\{\left(\bar{E}_{k} \star \zeta_{\varepsilon}\right)+v(p) \wedge\left(\bar{B}_{k} \star \zeta_{\varepsilon}\right)\right\} \cdot \frac{p}{|p|}\left(\chi_{R}(|p|)+\frac{|p|}{R} \chi^{\prime}\left(\frac{|p|}{R}\right)\right) d t d x d p .
\end{aligned}
$$

By passing to the limit for $R \rightarrow+\infty$ we deduce that:

$$
\begin{align*}
& \alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f_{k}(t, x, p) \cdot|p| d t d x d p \leq \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(t, x, p) \cdot|p| d t d \sigma d p \\
& \quad+\frac{|q|}{\alpha} \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g d t d \sigma d p \cdot \frac{1}{\varepsilon_{2}^{3 / 2}}\left\|\zeta_{2}\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}\left\|E_{k}\right\|_{L^{\infty}\left(\mathbb{R}_{t} ; L^{2}(\Omega)^{3}\right)} \tag{5.8}
\end{align*}
$$

and therefore, since $\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))|(1+\mathcal{E}(p)) g(t, x, p) d t d \sigma d p<+\infty$, we deduce that $\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f_{k} \cdot|p| d t d x d p$ is bounded by some constant $C$, uniformly in $k$. Similarly we obtain that $\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f \cdot|p| d t d x d p \leq C$. We can prove that $j_{k} \rightharpoonup j$ weakly in $L^{1}(] 0, T[\times \Omega)$. Indeed, for $\varphi \in L^{\infty}(] 0, T[\times \Omega)$ we have:

$$
\begin{aligned}
& \left|\int_{0}^{T} \int_{\Omega} j_{k} \varphi d t d x-\int_{0}^{T} \int_{\Omega} j \varphi d t d x\right| \\
\leq & \left|\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} q\left(f_{k}-f\right) v(p) \varphi \cdot \mathbf{1}_{\{|p| \leq R\}} d t d x d p\right| \\
& +\frac{|q|}{R} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}\left(f_{k}+f\right)|v(p)| \cdot|\varphi| \cdot|p| d t d x d p \\
\leq & \left|\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} q\left(f_{k}-f\right) v(p) \varphi \cdot \mathbf{1}_{\{|p| \leq R\}} d t d x d p\right|+\frac{2 C \cdot c_{0} \cdot|q| \cdot\|\varphi\|_{\infty}}{R} .
\end{aligned}
$$

We take $R_{\delta}$ large enough such that $2 C \cdot c_{0} \cdot|q| \cdot\|\varphi\|_{\infty} / R_{\delta}<\delta / 2$ and since $f_{k} \rightharpoonup f$ weakly $\star$ in $L^{\infty}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right)$ we have:

$$
\left|\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} q\left(f_{k}-f\right) v(p) \varphi \cdot \mathbf{1}_{\left\{|p| \leq R_{\delta}\right\}} d t d x d p\right|<\frac{\delta}{2}, \forall k \geq k_{\delta}
$$

which implies that $\left|\int_{0}^{T} \int_{\Omega}\left(j_{k}-j\right) \varphi d t d x\right| \leq \delta, \forall k \geq k_{\delta}$. We deduce the pointwise convergence $\left(\bar{j}_{k} \star \zeta_{\varepsilon}^{\sqrt{ }}\right)(t, x) \rightarrow\left(\bar{j} \star \zeta_{\varepsilon}^{\sqrt{\prime}}\right)(t, x), \forall(t, x) \in \mathbb{R}_{t} \times \Omega$. Moreover, since $\left(\bar{j}_{k} \star \zeta_{\varepsilon}^{\sqrt{\prime}}\right)_{k}$ is bounded we have $\bar{j}_{k} \star \zeta_{\varepsilon}^{\sqrt{\prime}} \rightarrow \bar{j} \star \zeta_{\varepsilon}^{\vee}$ in $L^{2}(] 0, T[\times \Omega)^{3}$. Consider now $(\varphi, \psi) \in$ $C\left(\mathbb{R}_{t}, D\left(\mathcal{A}^{\star}\right)\right) \cap C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right) T$ periodic and note $(f, g)=\alpha(\varphi, \psi)-\frac{d}{d t}(\varphi, \psi)+\mathcal{A}^{\star}(\varphi, \psi)$. We have:

$$
\begin{align*}
\int_{0}^{T}\left\langle\left(\tilde{E}_{k}(t), \tilde{B}_{k}(t)\right),(f(t), g(t))\right\rangle_{\mathcal{H}} d t & =c_{0} \int_{0}^{T} \int_{\partial \Omega}(n \wedge \varphi) \cdot h(t, x) d t d \sigma \\
& -\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega}\left(\bar{j}_{k} \star \zeta_{\varepsilon}^{\sqrt{ }}\right)(t, x) \cdot \varphi(t, x) d t d x \tag{5.9}
\end{align*}
$$

Since $\left\|\left(\tilde{E}_{k}, \tilde{B}_{k}\right)\right\|_{L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)} \leq M_{\alpha, \varepsilon}, \forall k$ we can suppose that, at least for a subsequence, we have $\left(\tilde{E}_{k}, \tilde{B}_{k}\right) \rightharpoonup(e, b)$ weakly $\star$ in $L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$. By passing to the limit in (5.9) for $k \rightarrow+\infty$ one gets:

$$
\begin{aligned}
& \int_{0}^{T}\langle(e(t), b(t)),(f(t), g(t))\rangle_{\mathcal{H}} d t=c_{0} \int_{0}^{T} \int_{\partial \Omega}(n \wedge \varphi) \cdot h(t, x) d t d \sigma \\
&-\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega}\left(\bar{j} \star \zeta_{\varepsilon}^{\sqrt{ }}\right)(t, x) \cdot \varphi(t, x) d t d x
\end{aligned}
$$

and thus $(e, b)=(\tilde{E}, \tilde{B})=\mathcal{F}(E, B)$ (the unique $T$ periodic weak solution of (4.3),(4.2) associated to $\bar{j} \star \zeta_{\varepsilon}^{\sqrt{ }}$ ). By the uniqueness of the $T$ periodic weak solution of (4.3), (4.2) we deduce also that all the sequence $\mathcal{F}\left(E_{k}, B_{k}\right)=\left(\tilde{E}_{k}, \tilde{B}_{k}\right)$ converges to $(\tilde{E}, \tilde{B})=$ $\mathcal{F}(E, B)$ weakly $\star$ in $L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$, or the application $\mathcal{F}$ is continuous in respect to the weak $\star$ topology of $L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$.
5.2. The classical case. Let us analyze now the classical case, with $\mathcal{E}(p)=$ $\frac{|p|^{2}}{2 m}, v(p)=\frac{p}{m}, \forall p \in \mathbb{R}_{p}^{3}$. We introduce also the energy and the velocity functions:

$$
\mathcal{E}_{c}(p)=m c^{2}\left(\left(1+\frac{|p|^{2}}{m^{2} c^{2}}\right)^{1 / 2}-1\right), v_{c}(p)=\frac{p}{m}\left(1+\frac{|p|^{2}}{m^{2} c^{2}}\right)^{-1 / 2}
$$

with $c>0$. Observe that $\lim _{c \rightarrow+\infty} \mathcal{E}_{c}(p)=\mathcal{E}(p), \lim _{c \rightarrow+\infty} v_{c}(p)=v(p)$ uniformly on bounded subsets of $\mathbb{R}_{p}^{3}$. The idea is to get the existence for the perturbed classical Vlasov-Maxwell system by letting $c \rightarrow+\infty$ in the perturbed relativistic Vlasov-Maxwell system (but keeping $c_{0}$ fixed in the perturbed Maxwell equations).

TheOrem 5.8. Assume that $g \in L^{\infty}\left(\mathbb{R}_{t} \times \Sigma^{-}\right)$, $h$ are $T$ periodic, with $g \geq 0$ and

$$
\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| \mathcal{E}(p) g(t, x, p) d t d \sigma d p+\int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma<+\infty
$$

and there is $\tilde{h} \in C^{1}\left(\mathbb{R}_{t} ; H^{1}(\Omega)^{3}\right) \cap C^{2}\left(\mathbb{R}_{t} ; L^{2}(\Omega)^{3}\right) T$ periodic such that $\left.n \wedge \tilde{h}\right|_{\mathbb{R}_{t} \times \partial \Omega}=h$. Then for every $\alpha, \varepsilon_{1}, \varepsilon_{2}>0$ there is a $T$ periodic solution for the perturbed classical Vlasov-Maxwell system

$$
\begin{gather*}
\alpha f+\partial_{t} f+v(p) \cdot \nabla_{x} f+q\left(\left(\bar{E} \star \zeta_{\varepsilon}\right)+v(p) \wedge\left(\bar{B} \star \zeta_{\varepsilon}\right)\right) \cdot \nabla_{p} f=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}, \\
\alpha E+\partial_{t} E-c_{0}^{2} \operatorname{rot} B=-\frac{\bar{j} \star \zeta_{\varepsilon}^{\sqrt{\prime}}}{\varepsilon_{0}}, \alpha B+\partial_{t} B+\operatorname{rot} E=0,(t, x) \in \mathbb{R}_{t} \times \Omega,  \tag{5.10}\\
f(t, x, p)=g(t, x, p),(t, x, p) \in \mathbb{R}_{t} \times \Sigma^{-}, n \wedge E(t, x)+c_{0} n \wedge(n \wedge B(t, x)) \\
=h(t, x),(t, x) \in \mathbb{R}_{t} \times \partial \Omega,
\end{gather*}
$$

with $j=q \int_{\mathbb{R}_{p}^{3}} f(t, x, p) v(p) d p$. Moreover $(E, B) \in C\left(\mathbb{R}_{t} ; H(\operatorname{rot} ; \Omega)^{2}\right) \cap C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$ and the solution verifies:

$$
\begin{array}{rl}
\alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} & f(t, x, p)(1+\mathcal{E}(p)) d t d x d p+\alpha \int_{0}^{T} \int_{\Omega}\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d t d x \\
& +\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f(1+\mathcal{E}(p)) d t d \sigma d p \\
& +\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}|n \wedge E|^{2}+\frac{1}{\mu_{0}}|n \wedge B|^{2}\right\} d t d \sigma \\
= & \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(1+\mathcal{E}(p)) d t d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma \tag{5.11}
\end{array}
$$

$$
\alpha \rho+\partial_{t} \rho+\operatorname{div} j=0, \operatorname{div} E=\frac{1}{\varepsilon_{0}}\left(\bar{\rho} \star \zeta_{\varepsilon}^{\sqrt{ }}\right), \text { div } B=0 \text { in } \mathcal{D}^{\prime}(] 0, T[\times \Omega) .
$$

Proof. Since $\int_{0}^{T} \int_{\Sigma^{-}}\left|\left(v_{c}(p) \cdot n(x)\right)\right| \mathcal{E}_{c}(p) g d t d \sigma d p \leq \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| \mathcal{E}(p) g d t d \sigma d p$ $<+\infty$, by the Theorem 5.3 we deduce that there is a $T$ periodic solution $\left(f_{c}, E_{c}, B_{c}\right)$ for the system:

$$
\begin{array}{r}
\alpha f_{c}+\partial_{t} f_{c}+v_{c}(p) \cdot \nabla_{x} f_{c}+q\left(\left(\bar{E}_{c} \star \zeta_{\varepsilon}\right)+v_{c}(p) \wedge\left(\bar{B}_{c} \star \zeta_{\varepsilon}\right)\right) \cdot \nabla_{p} f_{c} \\
=0,(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}, \\
\alpha E_{c}+\partial_{t} E_{c}-c_{0}^{2} \operatorname{rot} B_{c}=-\frac{\bar{j}_{c} \star \zeta_{\varepsilon}^{\vee}}{\varepsilon_{0}}, \alpha B_{c}+\partial_{t} B_{c}+\operatorname{rot} E_{c}=0,(t, x) \in \mathbb{R}_{t} \times \Omega, \\
\begin{aligned}
f_{c}(t, x, p)=g(t, x, p),(t, x, p) \in \mathbb{R}_{t} \times \Sigma^{-}, n \wedge E_{c}(t, x) & +c_{0} n \wedge\left(n \wedge B_{c}(t, x)\right) \\
& =h,(t, x) \in \mathbb{R}_{t} \times \partial \Omega
\end{aligned}
\end{array}
$$

with $j_{c}=q \int_{\mathbb{R}_{p}^{3}} f_{c}(t, x, p) v_{c}(p) d p$. Indeed the Propositions 5.1, 5.2 hold true by defining the application $\mathcal{F}_{c}$ as before and by taking $\mathcal{C}_{c}=\left\{(E, B) \in \mathcal{X}\| \|(E, B) \|_{L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)} \leq\right.$ $\left.M_{\alpha, \varepsilon, c}\right\}$, where

$$
\begin{aligned}
M_{\alpha, \varepsilon, c}= & c_{0}^{1 / 2}\left(\frac{1}{\alpha T}+4\right)^{1 / 2}\|h\|_{L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)} \\
& +\left(\frac{1}{\alpha T}+4\right) \frac{c|q| \cdot\left\|\zeta_{2}\right\|_{L^{2}}}{\alpha \cdot \varepsilon_{0} \cdot \varepsilon_{2}^{3 / 2}} \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g d t d \sigma d p
\end{aligned}
$$

Moreover, the Proposition 5.4 applies as well and thus we have:

$$
\begin{aligned}
& \alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f_{c}(t, x, p)\left(1+\mathcal{E}_{c}(p)\right) d t d x d p+\alpha \int_{0}^{T} \int_{\Omega}\left\{\varepsilon_{0}\left|E_{c}(t, x)\right|^{2}+\frac{1}{\mu_{0}}\left|B_{c}(t, x)\right|^{2}\right\} d t d x \\
& \quad+\int_{0}^{T} \int_{\Sigma^{+}}\left(v_{c}(p) \cdot n(x)\right) \gamma^{+} f_{c}\left(1+\mathcal{E}_{c}(p)\right) d t d \sigma d p \\
& \quad+\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}\left|n \wedge E_{c}\right|^{2}+\frac{1}{\mu_{0}}\left|n \wedge B_{c}\right|^{2}\right\} d t d \sigma \\
& =\int_{0}^{T} \int_{\Sigma^{-}}\left|\left(v_{c}(p) \cdot n(x)\right)\right| g\left(1+\mathcal{E}_{c}(p)\right) d t d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma \\
& \leq \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(1+\mathcal{E}(p)) d t d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma .
\end{aligned}
$$

We deduce that $E_{c}, B_{c}$ and $n \wedge E_{c}, n \wedge B_{c}$ are uniformly bounded in $L^{2}(] 0, T\left[; L^{2}(\Omega)^{3}\right)$, respectively in $L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)$ in respect to $c>0$. Observe also that $\left(E_{c}, B_{c}\right)_{c}$ is bounded in $L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)$. Indeed, by observing that $\left(v_{c}(p) \cdot p\right)=\left|v_{c}(p)\right| \cdot|p| \in$ $\left[\mathcal{E}_{c}(p), 2 \mathcal{E}_{c}(p)\right]$, we have:

$$
\begin{aligned}
& |q|^{-1}\left\|j_{c}\right\|_{L^{1}\left(10, T\left[; L^{1}(\Omega)^{3}\right)\right.} \\
& \leq \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}\left|v_{c}(p)\right| \cdot f_{c}(t, x, p) d t d x d p \\
= & \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}\left|v_{c}(p)\right| \cdot f_{c} \cdot \mathbf{1}_{\{|p| \leq 1\}} d t d x d p+\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}\left|v_{c}(p)\right| \cdot f_{c} \cdot \mathbf{1}_{\{|p|>1\}} d t d x d p \\
\leq & \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \frac{1}{m} f_{c}(t, x, p) d t d x d p+\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} 2 \mathcal{E}_{c}(p) f_{c}(t, x, p) d t d x d p \\
\leq & \frac{1}{\alpha}\left(\int_{0}^{T} \int_{\Sigma^{-}}\left(m^{-1}+2 \mathcal{E}(p)\right)|(v(p) \cdot n(x))| g d t d \sigma d p+c_{0} \varepsilon_{0} \int_{0}^{T} \int_{\partial \Omega}|h|^{2} d t d \sigma\right) \\
= & C_{1} .
\end{aligned}
$$

For the last inequality we use Remark 5.5. Therefore, as in the proof of Proposition 5.1 we deduce that:

$$
\begin{aligned}
&\left\|\left(E_{c}, B_{c}\right)\right\|_{L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right)} \leq c_{0}^{1 / 2}\left(\frac{1}{\alpha T}+4\right)^{1 / 2}\|h\|_{L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)} \\
&+\left(\frac{1}{\alpha T}+4\right) \frac{\left\|\zeta_{2}\right\|_{L^{2}}}{\varepsilon_{0} \cdot \varepsilon_{2}^{3 / 2}} \cdot|q| \cdot C_{1}=C_{2}
\end{aligned}
$$

Take $\left(c_{k}\right)_{k}$ with $c_{k}>0, \forall k$ and $\lim _{k \rightarrow+\infty} c_{k}=+\infty$. We denote by $\left(f_{k}, E_{k}, B_{k}\right)$ the corresponding solution. After extraction of subsequences we can assume that:

$$
\begin{aligned}
& f_{k} \rightharpoonup f \text { weakly } \star \text { in } L^{\infty}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right), \gamma^{+} f_{k} \rightharpoonup \gamma^{+} f \text { weakly } \star \text { in } L^{\infty}\left(\mathbb{R}_{t} \times \Sigma^{+}\right), \\
& \qquad\left(E_{k}, B_{k}\right) \rightharpoonup(E, B) \text { weakly } \star \text { in } L^{\infty}\left(\mathbb{R}_{t} ; \mathcal{H}\right) \\
& \left(n \wedge E_{k}, n \wedge B_{k}\right) \rightharpoonup(n \wedge E, n \wedge B), \text { weakly in } L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)
\end{aligned}
$$

By standard arguments we deduce that:

$$
\begin{aligned}
& \alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p)(1+\mathcal{E}(p)) d t d x d p+\alpha \int_{0}^{T} \int_{\Omega}\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d t d x \\
& \quad+\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f(1+\mathcal{E}(p)) d t d \sigma d p+\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}|n \wedge E|^{2}+\frac{1}{\mu_{0}}|n \wedge B|^{2}\right\} d t d \sigma \\
& \quad \leq \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(1+\mathcal{E}(p)) d t d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma=C(g, h) .
\end{aligned}
$$

As usual we prove that $\lim _{k \rightarrow+\infty}\left(\bar{E}_{k} \star \zeta_{\varepsilon}, \bar{B}_{k} \star \zeta_{\varepsilon}\right)=\left(\bar{E} \star \zeta_{\varepsilon}, \bar{B} \star \zeta_{\varepsilon}\right)$ in $L^{2}(] 0, T[\times \Omega)$ and by combining with the weak $\star$ convergence of $f_{k}, \gamma^{+} f_{k}$, after passing to the limit for $k \rightarrow+\infty$ in the Green formula (2.9), we deduce that $f, \gamma^{+} f$ verify the weak formulation of the perturbed $T$ periodic Vlasov problem corresponding to the electro-magnetic field $\left(\bar{E} \star \zeta_{\varepsilon}, \bar{B} \star \zeta_{\varepsilon}\right)$ and to the energy and velocity functions $\mathcal{E}(p), v(p)$ (since ( $\bar{E} \star$ $\zeta_{\varepsilon}, \bar{B} \star \zeta_{\varepsilon}$ ) is regular, by the uniqueness of the $T$ periodic weak solution when $\alpha>0$ we deduce that $f$ is also the $T$ periodic mild solution). Now we want to pass to the limit in the perturbed Maxwell equations. By observing that:

$$
\begin{aligned}
\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}\left|v_{k}(p)\right| \cdot f_{k}(t, x, p) \cdot \mathbf{1}_{\{|p|>R\}} d t d x d p & \leq \frac{1}{R} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}\left|v_{k}(p)\right| \cdot|p| \cdot f_{k}(t, x, p) d t d x d p \\
& \leq \frac{2}{R} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}_{k}(p) f_{k}(t, x, p) d t d x d p \\
& \leq \frac{2}{R} \cdot \frac{C(g, h)}{\alpha}
\end{aligned}
$$

and:

$$
\begin{aligned}
\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}|v(p)| \cdot f(t, x, p) \cdot \mathbf{1}_{\{|p|>R\}} d t d x d p & \leq \frac{1}{R} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}|v(p)| \cdot|p| \cdot f(t, x, p) d t d x d p \\
& =\frac{2}{R} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f(t, x, p) d t d x d p \\
& \leq \frac{2}{R} \cdot \frac{C(g, h)}{\alpha}
\end{aligned}
$$

we deduce as in the proof of Proposition 5.2 that $\lim _{k \rightarrow+\infty} j_{k}=j$ weakly in $L^{1}(] 0, T[\times \Omega)^{3}$ and that $\lim _{k \rightarrow+\infty}\left(\bar{j}_{k} \star \zeta_{\varepsilon}^{\sqrt{\prime}}\right)=\bar{j} \star \zeta_{\varepsilon}^{\sqrt{ }}$ strongly in $L^{2}(] 0, T[\times \Omega)^{3}$. Consider $(\varphi, \psi) \in C\left(\mathbb{R}_{t} ; D\left(\mathcal{A}^{\star}\right)\right) \cap C^{1}\left(\mathbb{R}_{t} ; \mathcal{H}\right) T$ periodic and note $(f, g)=\alpha(\varphi, \psi)-\frac{d}{d t}(\varphi, \psi)+$ $\mathcal{A}^{\star}(\varphi, \psi)$. We have:

$$
\begin{aligned}
\int_{0}^{T}\left\langle\left(E_{k}(t), B_{k}(t)\right),(f(t), g(t))\right\rangle_{\mathcal{H}} d t & =c_{0} \int_{0}^{T} \int_{\partial \Omega}(n \wedge \varphi) \cdot h(t, x) d t d \sigma \\
& -\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega}\left(\bar{j}_{k} \star \zeta_{\varepsilon}^{\sqrt{\prime}}\right)(t, x) \cdot \varphi(t, x) d t d x
\end{aligned}
$$

and after passing to the limit for $k \rightarrow+\infty$ we deduce that $(E, B)$ is a $T$ periodic weak solution for the perturbed Maxwell equations, corresponding to the current density $\bar{j} \star \zeta_{\varepsilon}^{\sqrt{ }}$. By the Proposition 4.6 we deduce that $(E, B)$ is a $T$ periodic strong solution. The other statements follow exactly as in the relativistic case.

## 6. A priori estimates

In the following we want to establish a priori estimates for the normal trace of the electro-magnetic field as well as for the total (kinetic and electro-magnetic) energy. For this we suppose that $\partial \Omega$ is strictly star-shaped in respect to some point $x_{0} \in \Omega$ (i.e., $\exists r>0$ such that $\left.\left(x-x_{0}\right) \cdot n(x) \geq r, \forall x \in \partial \Omega\right)$. After translation we can assume that $x_{0}=0 \in \Omega$ and thus $x \cdot n(x) \geq r, \forall x \in \partial \Omega$. This hypothesis was already used in order to estimate the solutions of the Maxwell equations by using the multiplier method (see [28]). These estimates for the normal trace of the electro-magnetic field and the total energy are summarized in the following proposition.
Proposition 6.1. Assume that $\Omega$ is bounded, with $\partial \Omega$ smooth and strictly starshaped (in respect to $0 \in \Omega$ ). Under the hypothesis of the Theorems 5.8, 5.3, consider ( $f, E, B$ ) the $T$ periodic solution of the perturbed Vlasov-Maxwell system (classical case (5.10), or relativistic case (5.1)) with fixed $0<\alpha, \varepsilon_{1}, \varepsilon_{2}<1$. Then the electro-magnetic field has normal trace $(n \cdot E, n \cdot B) \in L^{2}(] 0, T[\times \partial \Omega)^{2}$ :

$$
\begin{array}{r}
\int_{\Omega} \operatorname{div} E(t) \theta(x) d x+\int_{\Omega} E(t, x) \cdot \nabla_{x} \theta d x=\int_{\partial \Omega}(n(x) \cdot E(t, x)) \theta(x) d \sigma, \forall \theta \in C^{1}(\bar{\Omega}) \\
\text { a.e.t } \in \mathbb{R}_{t},
\end{array}
$$

and:

$$
\int_{\Omega} B(t, x) \cdot \nabla_{x} \theta d x=\int_{\partial \Omega}(n(x) \cdot B(t, x)) \theta(x) d \sigma, \forall \theta \in C^{1}(\bar{\Omega}), \text { a.e. } t \in \mathbb{R}_{t} .
$$

Moreover, the solution satisfies the following estimate:

$$
\begin{array}{rl}
\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} & f(t, x, p) \mathcal{E}(p) d t d x d p+\int_{0}^{T} \int_{\Omega}\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d t d x \\
& +\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x))(1+\mathcal{E}(p)) \gamma^{+} f(t, x, p) d t d \sigma d p \\
+ & \int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}\left[(n \cdot E)^{2}+|n \wedge E|^{2}\right]+\frac{1}{\mu_{0}}\left[(n \cdot B)^{2}+|n \wedge B|^{2}\right]\right\} d t d \sigma \\
\leq & C \cdot\left\{\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))|(1+\mathcal{E}(p)) g d t d \sigma d p+\int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma\right\}+\frac{C \cdot \varepsilon_{2}^{s}}{\alpha^{q} \cdot \varepsilon_{1}^{r}}
\end{array}
$$

for some constant $C$ and exponents $q, r, s>0$ and the total energy is uniformly bounded in time:

$$
\begin{aligned}
& \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f(t, x, p) d x d p+\int_{\Omega}\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d x \leq \frac{C \cdot \varepsilon_{2}^{s}}{\alpha^{q} \cdot \varepsilon_{1}^{r}} \\
+ & C \cdot\left\{\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))|(1+\mathcal{E}(p)) g(t, x, p) d t d \sigma d p+\int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma\right\}, \forall t \in \mathbb{R}_{t} .
\end{aligned}
$$

For the proof of Proposition 6.1 we need several lemmas.
Lemma 6.2. Assume that $\partial \Omega$ is regular $\left(C^{1}\right), u \in H(\operatorname{div} ; \Omega) \cap H($ rot $; \Omega)$. Then we have the following equality in $\mathcal{D}^{\prime}(\Omega)$ :

$$
\begin{equation*}
u_{i} \text { div } u-(u \wedge \operatorname{rot} u)_{i}=\sum_{j=1}^{3} \frac{\partial}{\partial x_{j}}\left(u_{i} u_{j}\right)-\frac{1}{2} \frac{\partial}{\partial x_{i}}|u|^{2}, \forall 1 \leq i \leq 3 . \tag{6.1}
\end{equation*}
$$

The proof of the above lemma is immediate and is left to the reader. In order to estimate the normal trace of the electro-magnetic field we need the following lemma. For the sake of presentation we start with the stationary case.

Lemma 6.3. Assume that $\Omega$ is bounded with $\partial \Omega$ regular and strictly star-shaped in respect to $0 \in \Omega$. Consider $u \in H($ div $; \Omega) \cap H(\operatorname{rot} ; \Omega)$ with $n \wedge u \in L^{2}(\partial \Omega)^{3}$ such that

$$
\begin{array}{r}
\int_{\Omega}\left(u_{i} \operatorname{div} u-(u \wedge \operatorname{rot} u)_{i}\right) \theta(x) d x+\int_{\Omega} w_{i} \cdot \nabla_{x} \theta d x \\
=\int_{\Omega} v_{i} \theta(x) d x+\int_{\partial \Omega}\left(n(x) \cdot w_{i}(x)\right) \theta(x) d \sigma \tag{6.2}
\end{array}
$$

for all functions $\theta \in C^{1}(\bar{\Omega})^{3}, 1 \leq i \leq 3$, where $v_{i}, w_{i}=\left(w_{i j}\right)_{1 \leq j \leq 3}, n \cdot w_{i}$ are some given functions verifying $v_{i} \in L^{1}(\Omega), w_{i} \in L^{1}(\Omega)^{3}, w_{i i} \geq 0, n \cdot w_{i} \in L^{1}(\partial \Omega), 1 \leq i \leq 3$. Then $u$ has a normal trace in $L^{2}(\partial \Omega)$ i.e., there is $n \cdot u \in L^{2}(\partial \Omega)$ such that $\int_{\Omega}$ div $u \theta(x) d x+$ $\int_{\Omega} u \cdot \nabla_{x} \theta d x=\int_{\partial \Omega}(n(x) \cdot u(x)) \theta(x) d \sigma, \forall \theta \in C^{1}(\bar{\Omega})$. Moreover we have the estimate:

$$
\begin{aligned}
\|n \cdot u\|_{L^{2}(\partial \Omega)}^{2}+\sum_{i=1}^{3}\left\|w_{i i}\right\|_{L^{1}(\Omega)} & +\|u\|_{L^{2}(\Omega)^{3}}^{2} \leq C\left\{\|n \wedge u\|_{L^{2}(\partial \Omega)^{3}}^{2}\right. \\
& \left.+\|v\|_{L^{1}(\Omega)^{3}}+\sum_{i=1}^{3}\left\|n \cdot w_{i}\right\|_{L^{1}(\partial \Omega)}\right\}
\end{aligned}
$$

Proof. Since $u \in H(\operatorname{div} ; \Omega) \cap H(\operatorname{rot} ; \Omega), n \wedge u \in L^{2}(\partial \Omega)^{3}$, we can approximate $u$ by smooth functions $u^{k} \in C^{1}(\bar{\Omega})^{3}$ such that $u^{k} \rightarrow u, \operatorname{rot} u^{k} \rightarrow \operatorname{rot} u \operatorname{in} L^{2}(\Omega)^{3}$, $\operatorname{div} u^{k} \rightarrow$ $\operatorname{div} u$ in $L^{2}(\Omega), n \wedge u^{k} \rightarrow n \wedge u$ in $L^{2}(\partial \Omega)^{3}$. By using the Lemma 6.2 we have:

$$
\begin{align*}
\sum_{j=1}^{3} \frac{\partial}{\partial x_{j}}\left(u_{i}^{k} u_{j}^{k}\right)-\frac{1}{2} \frac{\partial}{\partial x_{i}}\left|u^{k}\right|^{2} & =u_{i}^{k} \operatorname{div} u^{k}-\left(u^{k} \wedge \operatorname{rot} u^{k}\right)_{i} \\
& =u_{i} \operatorname{div} u-(u \wedge \operatorname{rot} u)_{i}+r_{i}^{k}, \forall 1 \leq i \leq 3 \tag{6.3}
\end{align*}
$$

where $r_{i}^{k} \rightarrow 0$ in $L^{1}(\Omega)$ as $k \rightarrow+\infty$. By using (6.3) and the hypothesis (6.2) with the test function $\theta(x)=x_{i}$ we obtain for $1 \leq i \leq 3$ :

$$
\begin{aligned}
& \int_{\partial \Omega} \sum_{j=1}^{3} u_{i}^{k} u_{j}^{k} x_{i} n_{j} d \sigma-\int_{\Omega} \sum_{j=1}^{3} u_{i}^{k} u_{j}^{k} \delta_{i j} d x-\frac{1}{2} \int_{\partial \Omega}\left|u^{k}\right|^{2} x_{i} n_{i} d \sigma+\frac{1}{2} \int_{\Omega}\left|u^{k}\right|^{2} d x \\
& \quad=\int_{\Omega} x_{i} v_{i}(x) d x+\int_{\partial \Omega}\left(n(x) \cdot w_{i}(x)\right) x_{i} d \sigma-\int_{\Omega} \sum_{j=1}^{3} w_{i j} \delta_{i j} d x+\int_{\Omega} r_{i}^{k} x_{i} d x
\end{aligned}
$$

By taking the sum for $1 \leq i \leq 3$ in (6.4) we obtain:

$$
\begin{align*}
\int_{\Omega} \sum_{i=1}^{3} w_{i i} d x & +\int_{\partial \Omega}\left(u^{k} \cdot x\right)\left(u^{k} \cdot n\right) d \sigma+\frac{1}{2} \int_{\Omega}\left|u^{k}\right|^{2} d x-\frac{1}{2} \int_{\partial \Omega}\left|u^{k}\right|^{2}(n \cdot x) d \sigma \\
& =\int_{\Omega} x \cdot v(x) d x+\int_{\partial \Omega} \sum_{i=1}^{3}\left(n(x) \cdot w_{i}(x)\right) x_{i} d \sigma+\int_{\Omega} r^{k} \cdot x d x \tag{6.5}
\end{align*}
$$

By using the decomposition $u^{k}=\left(n \cdot u^{k}\right) n-n \wedge\left(n \wedge u^{k}\right)$ after easy computations we obtain the identity:

$$
\begin{align*}
& \quad \int_{\Omega} \sum_{i=1}^{3} w_{i i} d x+\frac{1}{2} \int_{\Omega}\left|u^{k}\right|^{2} d x+\frac{1}{2} \int_{\partial \Omega}\left(u^{k} \cdot n\right)^{2}(n(x) \cdot x) d \sigma=\frac{1}{2} \int_{\partial \Omega}\left|u^{k} \wedge n\right|^{2}(n(x) \cdot x) d \sigma \\
& + \\
& \int_{\partial \Omega}\left(u^{k} \cdot n\right)\left[\left(n \wedge\left(n \wedge u^{k}\right)\right) \cdot x\right] d \sigma+\int_{\Omega} x \cdot v(x) d x  \tag{6.6}\\
& +\int_{\partial \Omega} \sum_{i=1}^{3}\left(n(x) \cdot w_{i}(x)\right) x_{i} d \sigma+\int_{\Omega} r^{k} \cdot x d x
\end{align*}
$$

By our hypothesis there is $0<r \leq R$ such that $r \leq(n(x) \cdot x) \leq|x| \leq R, \forall x \in \partial \Omega$ and thus we have:

$$
\begin{array}{r}
\int_{\Omega} \sum_{i=1}^{3} w_{i i} d x+\frac{1}{2}\left\|u^{k}\right\|_{L^{2}(\Omega)^{3}}^{2}+\frac{r}{2}\left\|n \cdot u^{k}\right\|_{L^{2}(\partial \Omega)}^{2} \leq \frac{R}{2}\left\|n \wedge u^{k}\right\|_{L^{2}(\partial \Omega)^{3}}^{2}+R \sum_{i=1}^{3}\left\|n \cdot w_{i}\right\|_{L^{1}(\partial \Omega)} \\
+R \cdot\|v\|_{L^{1}(\Omega)^{3}}+R\left\|r^{k}\right\|_{L^{1}(\Omega)^{3}}+R\left\|n \cdot u^{k}\right\|_{L^{2}(\partial \Omega)}\left\|n \wedge u^{k}\right\|_{L^{2}(\partial \Omega)^{3}} \tag{6.7}
\end{array}
$$

which implies that:

$$
\begin{align*}
\sum_{i=1}^{3}\left\|w_{i i}\right\|_{L^{1}(\Omega)} & +\left\|u^{k}\right\|_{L^{2}(\Omega)^{3}}^{2}+\left\|\left(n \cdot u^{k}\right)\right\|_{L^{2}(\partial \Omega)}^{2} \leq C(r, R)\left(\left\|n \wedge u^{k}\right\|_{L^{2}(\partial \Omega)^{3}}^{2}+\|v\|_{L^{1}(\Omega)^{3}}\right. \\
& \left.+\sum_{i=1}^{3}\left\|n \cdot w_{i}\right\|_{L^{1}(\partial \Omega)}+\left\|r^{k}\right\|_{L^{1}(\Omega)^{3}}\right) \tag{6.8}
\end{align*}
$$

Since $n \wedge u^{k} \rightarrow n \wedge u$ in $L^{2}(\partial \Omega)^{3}$ and $r^{k} \rightarrow 0$ in $L^{1}(\Omega)^{3}$ we deduce that $\left(n \cdot u^{k}\right)_{k}$ is bounded in $L^{2}(\partial \Omega)$. In fact we can prove that $\left(n \cdot u^{k}\right)_{k}$ converges in $L^{2}(\partial \Omega)$. For this let us introduce the bilinear application:

$$
a_{i}(f, g)=f_{i} \operatorname{div} g-(f \wedge \operatorname{rot} g)_{i}, f \in L^{2}(\Omega)^{3}, g \in H(\operatorname{div} ; \Omega) \cap H(\operatorname{rot} ; \Omega)
$$

We have:

$$
\begin{array}{r}
a_{i}(f-g, f-g)=a_{i}(f, f)-a_{i}(g, g)-a_{i}(g, f-g)-a_{i}(f-g, g), \\
\forall f, g \in H(\operatorname{div} ; \Omega) \cap H(\operatorname{rot} ; \Omega) .
\end{array}
$$

Observe that $a_{i}\left(u^{k}, u^{k}\right)=a_{i}(u, u)+r_{i}^{k}, a_{i}\left(u^{l}, u^{l}\right)=a_{i}(u, u)+r_{i}^{l}$. By taking into account that:

$$
\begin{aligned}
\left\|a_{i}\left(u^{l}, u^{k}-u^{l}\right)\right\|_{L^{1}(\Omega)} \leq\left\|u_{i}^{l}\right\|_{L^{2}(\Omega)}\left\|\operatorname{div} u^{k}-\operatorname{div} u^{l}\right\|_{L^{2}(\Omega)} & +\left\|u^{l}\right\|_{L^{2}(\Omega)^{3}} \| \operatorname{rot} u^{k} \\
& -\operatorname{rot} u^{l} \|_{L^{2}(\Omega)^{3}} \rightarrow 0
\end{aligned}
$$

when $k, l \rightarrow+\infty$ and:

$$
\begin{aligned}
&\left\|a_{i}\left(u^{k}-u^{l}, u^{l}\right)\right\|_{L^{1}(\Omega)} \leq\left\|u_{i}^{k}-u_{i}^{l}\right\|_{L^{2}(\Omega)}\left\|\operatorname{div} u^{l}\right\|_{L^{2}(\Omega)} \\
&+\left\|u^{k}-u^{l}\right\|_{L^{2}(\Omega)^{3}}\left\|\operatorname{rot} u^{l}\right\|_{L^{2}(\Omega)^{3}} \rightarrow 0,
\end{aligned}
$$

when $k, l \rightarrow+\infty$ we deduce that for $1 \leq i \leq 3$ we have:

$$
\begin{aligned}
\sum_{j=1}^{3} \frac{\partial}{\partial x_{j}}\left\{\left(u_{i}^{k}-u_{i}^{l}\right)\left(u_{j}^{k}-u_{j}^{l}\right)\right\} & -\frac{1}{2} \frac{\partial}{\partial x_{i}}\left|u^{k}-u^{l}\right|^{2}=a_{i}\left(u^{k}-u^{l}, u^{k}-u^{l}\right) \\
& =a_{i}\left(u^{k}, u^{k}\right)-a_{i}\left(u^{l}, u^{l}\right)-a_{i}\left(u^{l}, u^{k}-u^{l}\right)-a_{i}\left(u^{k}-u^{l}, u^{l}\right) \\
& =r_{i}^{k}-r_{i}^{l}-a_{i}\left(u^{l}, u^{k}-u^{l}\right)-a_{i}\left(u^{k}-u^{l}, u^{l}\right)=r_{i}^{k l}
\end{aligned}
$$

where $r^{k l} \rightarrow 0$ in $L^{1}(\Omega)^{3}$ when $k, l \rightarrow+\infty$. Now, by the previous computations (this time with $v=0, w=0, n \cdot w=0$ ) we deduce that:

$$
\left\|n \cdot u^{k}-n \cdot u^{l}\right\|_{L^{2}(\partial \Omega)} \leq C(r, R)\left\{\left\|n \wedge u^{k}-n \wedge u^{l}\right\|_{L^{2}(\partial \Omega)^{3}}+\left\|r^{k l}\right\|_{L^{1}(\Omega)^{3}}^{1 / 2}\right\}
$$

and thus $\left(n \cdot u^{k}\right)_{k}$ is a Cauchy sequence in $L^{2}(\partial \Omega)$, or $\left(n \cdot u^{k}\right)_{k}$ converges in $L^{2}(\partial \Omega)$. Moreover, the limit doesn't depend on the approximation sequence $\left(u^{k}\right)_{k}$ and we can associate to $u$ the normal trace $n \cdot u:=\lim _{k \rightarrow+\infty} n \cdot u^{k}$ in $L^{2}(\partial \Omega)$. By passing to the limit in (6.8) we find:

$$
\begin{align*}
\sum_{i=1}^{3}\left\|w_{i i}\right\|_{L^{1}(\Omega)} & +\|u\|_{L^{2}(\Omega)^{3}}^{2}+\|(n \cdot u)\|_{L^{2}(\partial \Omega)}^{2} \leq C(r, R)\left(\|n \wedge u\|_{L^{2}(\partial \Omega)^{3}}^{2}+\|v\|_{L^{1}(\Omega)^{3}}\right. \\
& \left.+\sum_{i=1}^{3}\left\|n \cdot w_{i}\right\|_{L^{1}(\partial \Omega)}\right) \tag{6.9}
\end{align*}
$$

Moreover we have $\int_{\Omega} \operatorname{div} u^{k} \theta(x) d x+\int_{\Omega} u^{k} \cdot \nabla_{x} \theta d x=\int_{\partial \Omega}\left(n(x) \cdot u^{k}(x)\right) \theta(x) d \sigma, \forall \theta \in$ $C^{1}(\bar{\Omega})$, and by passing to the limit for $k \rightarrow+\infty$ we find that:

$$
\int_{\Omega} \operatorname{div} u \theta(x) d x+\int_{\Omega} u \cdot \nabla_{x} \theta d x=\int_{\partial \Omega}(n(x) \cdot u(x)) \theta(x) d \sigma, \forall \theta \in C^{1}(\bar{\Omega})
$$

Once we have defined $n \cdot u \in L^{2}(\partial \Omega)$ we can define the trace of $u$ on $\partial \Omega$ by $\gamma u=$ $(n \cdot u) n-n \wedge(n \wedge u) \in L^{2}(\partial \Omega)^{3}$. By construction we have $(n \cdot \gamma u)=(n \cdot u)$ and $n \wedge \gamma u=$ $n \wedge u$ and therefore:

$$
\begin{aligned}
\|\gamma u\|_{L^{2}(\partial \Omega)^{3}}^{2} & =\|n \cdot \gamma u\|_{L^{2}(\partial \Omega)}^{2}+\|n \wedge \gamma u\|_{L^{2}(\partial \Omega)^{3}}^{2} \\
& \leq C\left\{\|n \wedge u\|_{L^{2}(\partial \Omega)^{3}}^{2}+\|v\|_{L^{1}(\Omega)^{3}}+\sum_{i=1}^{3}\left\|n \cdot w_{i}\right\|_{L^{1}(\partial \Omega)}\right\} .
\end{aligned}
$$

Moreover, the equality $\sum_{j=1}^{3} \frac{\partial}{\partial x_{j}}\left(u_{i} u_{j}\right)-\frac{1}{2} \frac{\partial}{\partial x_{i}}|u|^{2}=v_{i}+\operatorname{div} w_{i}$ holds in $\mathcal{D}^{\prime}(\bar{\Omega}), \forall 1 \leq$ $i \leq 3$. Indeed, for $\theta \in C^{1}(\bar{\Omega})$ we can write for $1 \leq i \leq 3$ :

$$
\begin{aligned}
\int_{\Omega}\left(\sum_{j=1}^{3} \frac{\partial}{\partial x_{j}}\left(u_{i}^{k} u_{j}^{k}\right)-\frac{1}{2} \frac{\partial}{\partial x_{i}}\left|u^{k}\right|^{2}\right) \theta(x) d x & =\int_{\Omega}\left(u_{i} \operatorname{div} u-(u \wedge \operatorname{rot} u)_{i}+r_{i}^{k}\right) \theta(x) d x \\
& =\int_{\Omega} v_{i}(x) \theta(x) d x-\int_{\Omega} w_{i}(x) \cdot \nabla_{x} \theta d x \\
& +\int_{\partial \Omega}\left(n(x) \cdot w_{i}(x)\right) \theta(x) d \sigma+\int_{\Omega} r_{i}^{k} \theta(x) d x
\end{aligned}
$$

After integration by parts we deduce that:

$$
\begin{aligned}
\int_{\partial \Omega} u_{i}^{k} \sum_{j=1}^{3} u_{j}^{k} n_{j} \theta(x) d \sigma- & \frac{1}{2} \int_{\partial \Omega}\left|u^{k}\right|^{2} n_{i} \theta(x) d \sigma-\int_{\Omega} u_{i}^{k} \sum_{j=1}^{3} u_{j}^{k} \frac{\partial \theta}{\partial x_{j}} d x+\frac{1}{2} \int_{\Omega}\left|u^{k}\right|^{2} \frac{\partial \theta}{\partial x_{i}} d x \\
= & \int_{\Omega} v_{i}(x) \theta(x) d x-\int_{\Omega} w_{i}(x) \cdot \nabla_{x} \theta d x \\
& +\int_{\partial \Omega}\left(n(x) \cdot w_{i}(x)\right) \theta(x) d \sigma+\int_{\Omega} r_{i}^{k} \theta(x) d x
\end{aligned}
$$

and by passing to the limit for $k \rightarrow+\infty$ we obtain:

$$
\begin{aligned}
& \int_{\partial \Omega}(\gamma u)_{i} \sum_{j=1}^{3}(\gamma u)_{j} n_{j} \theta(x) d \sigma-\frac{1}{2} \int_{\partial \Omega}|\gamma u|^{2} n_{i} \theta(x) d \sigma-\int_{\Omega} u_{i} \sum_{j=1}^{3} u_{j} \frac{\partial \theta}{\partial x_{j}} d x \\
&+\frac{1}{2} \int_{\Omega}|u|^{2} \frac{\partial \theta}{\partial x_{i}} d x \\
&=\int_{\Omega} v_{i}(x) \theta(x) d x-\int_{\Omega} w_{i}(x) \cdot \nabla_{x} \theta d x+\int_{\partial \Omega}\left(n(x) \cdot w_{i}(x)\right) \theta(x) d \sigma
\end{aligned}
$$

We have similar results for the time dependent case. The proof is left to the reader.
LEmma 6.4. Assume that $\Omega$ is a bounded domain with $\partial \Omega$ regular $\left(C^{1}\right)$ and strictly star-shaped in respect to the origin $0 \in \Omega$. Consider $u \in L^{2}(] 0, T[; H($ div $; \Omega)) \cap$ $L^{2}(] 0, T[; H($ rot $; \Omega))$ with $n \wedge u \in L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)$ such that:

$$
\begin{aligned}
\int_{0}^{T} \int_{\Omega}\left\{u_{i} d i v u\right. & \left.-(u \wedge \operatorname{rot} u)_{i}\right\} \theta(t, x) d t d x+\int_{0}^{T} \int_{\Omega} z_{i}(t, x) \partial_{t} \theta d t d x+\int_{0}^{T} \int_{\Omega} w_{i} \cdot \nabla_{x} \theta d t d x \\
= & \int_{0}^{T} \int_{\Omega} v_{i}(t, x) \theta(t, x) d t d x+\int_{0}^{T} \int_{\partial \Omega}\left(n(x) \cdot w_{i}(x)\right) \theta(t, x) d t d \sigma, 1 \leq i \leq 3
\end{aligned}
$$

for all functions $\theta \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right) T$ periodic in time, where $z, v, w_{i}=\left(w_{i j}\right)_{1 \leq j \leq 3}, n \cdot w_{i}$ are some given functions verifying $z, v \in L^{1}(] 0, T\left[; L^{1}(\Omega)^{3}\right), w_{i} \in L^{1}(] 0, T\left[; L^{1}(\Omega)^{3}\right), w_{i i}$ $\geq 0, n \cdot w_{i} \in L^{1}(] 0, T[\times \partial \Omega), 1 \leq i \leq 3$. Then $u$ has a normal trace $n \cdot u \in L^{2}(] 0, T[\times \partial \Omega)$ :

$$
\int_{\Omega} \operatorname{div} u(t) \theta(x) d x+\int_{\Omega} u(t, x) \cdot \nabla_{x} \theta d x=\int_{\partial \Omega}(n \cdot u(t)) \theta(x) d \sigma, \forall \theta \in C^{1}(\bar{\Omega}) \text { a.e.t } \in \mathbb{R}_{t} \text {, }
$$

and we have the estimate:

$$
\begin{aligned}
\|n \cdot u\|_{L^{2}(] 0, T[\times \partial \Omega)}^{2} & +\sum_{i=1}^{3}\left\|w_{i i}\right\|_{L^{1}(] 0, T[\times \Omega)}+\|u\|_{L^{2}(] 0, T\left[; L^{2}(\Omega)^{3}\right)}^{2} \leq C \cdot\left(\|n \wedge u\|_{L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)}^{2}\right. \\
& \left.+\|v\|_{L^{1}(] 0, T\left[; L^{1}(\Omega)^{3}\right)}+\sum_{i=1}^{3}\left\|n \cdot w_{i}\right\|_{\left.L^{1}(] 0, T[\times \partial \Omega)\right)}\right)
\end{aligned}
$$

REMARK 6.5. We can define the trace $\gamma u=(n \cdot u) n-n \wedge(n \wedge u) \in L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)$ and we have $\sum_{j=1}^{3} \frac{\partial}{\partial x_{j}}\left(u_{i} u_{j}\right)-\frac{1}{2} \frac{\partial}{\partial x_{i}}|u|^{2}=v_{i}+\partial_{t} z_{i}+\operatorname{div} w_{i}$ in $\mathcal{D}^{\prime}(] 0, T[\times \bar{\Omega}), 1 \leq i \leq 3$.

Remark 6.6. The previous results adapt easily when replacing $u \operatorname{div} u-u \wedge \operatorname{rot} u$ by $\varepsilon_{0}\{E \operatorname{div} E-E \wedge \operatorname{rot} E\}+\frac{1}{\mu_{0}}\{B \operatorname{div} B-B \wedge \operatorname{rot} B\}$.
We give here the details for the proof of Proposition 6.1.

Proof. (Proposition 6.1) Consider the test function $\varphi_{i}(t, x, p)=p_{i} \theta(t, x), 1 \leq i \leq$ $3, \theta \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right)$, $T$ periodic. By using the Green formula (remark that this is possible since $\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}(1+\mathcal{E}(p)) f d t d x d p<+\infty, \int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x))(1+\mathcal{E}(p)) \gamma^{+} f d t d \sigma d p<+\infty$ and thus in both classical and relativistic cases we have $\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}|p| f d t d x d p<+\infty$, $\left.\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x))|p| \gamma^{+} f d t d \sigma d p<+\infty\right)$ we deduce that:

$$
\begin{aligned}
\alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p) p_{i} \theta(t, x) d t d x d p & +\int_{0}^{T} \int_{\Sigma}(v(p) \cdot n(x)) p_{i} \gamma f(t, x, p) \theta(t, x) d t d \sigma d p \\
= & \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p) p_{i}\left(\partial_{t} \theta+v(p) \cdot \nabla_{x} \theta\right) d t d x d p \\
& +\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} q \theta(t, x) f(t, x, p)\left\{\left(\bar{E} \star \zeta_{\varepsilon}\right)\right. \\
& \left.+v(p) \wedge\left(\bar{B} \star \zeta_{\varepsilon}\right)\right\}_{i} d t d x d p
\end{aligned}
$$

Let us consider $i=1$ and compute the term involving the electro-magnetic field:

$$
\begin{aligned}
\mathcal{I}_{1}= & \int_{0}^{T} \int_{\Omega} \theta(t, x)\left\{\rho(t, x)\left(\bar{E}_{1} \star \zeta_{\varepsilon}\right)+j_{2}(t, x)\left(\bar{B}_{3} \star \zeta_{\varepsilon}\right)-j_{3}(t, x)\left(\bar{B}_{2} \star \zeta_{\varepsilon}\right)\right\} d t d x \\
= & \int_{0}^{T} \int_{\Omega}\left\{E_{1}(t, x)\left((\overline{\theta \rho}) \star \zeta_{\varepsilon}^{\sqrt{\prime}}\right)+B_{3}(t, x)\left(\left(\overline{\theta j_{2}}\right) \star \zeta_{\varepsilon}^{\sqrt{\prime}}\right)-B_{2}(t, x)\left(\left(\overline{\theta j_{3}}\right) \star \zeta_{\varepsilon}^{\sqrt{\prime}}\right)\right\} d t d x \\
= & \int_{0}^{T} \int_{\Omega} \theta\left\{E_{1}(t, x)\left(\bar{\rho} \star \zeta_{\varepsilon}^{\sqrt{ }}\right)+B_{3}(t, x)\left(\bar{j}_{2} \star \zeta_{\varepsilon}^{\sqrt{\prime}}\right)-B_{2}(t, x)\left(\bar{j}_{3} \star \zeta_{\varepsilon}^{\sqrt{ }}\right)\right\} d t d x \\
& +\int_{0}^{T} \int_{\Omega} E_{1}(t, x) \int_{0}^{T} \int_{\Omega}(\theta(s, y)-\theta(t, x)) \rho(s, y) \zeta_{\varepsilon}^{\sqrt{ }}(t-s, x-y) d s d y d t d x \\
& +\int_{0}^{T} \int_{\Omega} B_{3}(t, x) \int_{0}^{T} \int_{\Omega}(\theta(s, y)-\theta(t, x)) j_{2}(s, y) \zeta_{\varepsilon}^{\sqrt{ }}(t-s, x-y) d s d y d t d x \\
& -\int_{0}^{T} \int_{\Omega} B_{2}(t, x) \int_{0}^{T} \int_{\Omega}(\theta(s, y)-\theta(t, x)) j_{3}(s, y) \zeta_{\varepsilon}^{\sqrt{ }}(t-s, x-y) d s d y d t d x \\
= & M_{1}+R_{1}(\theta) .
\end{aligned}
$$

By using the perturbed Maxwell equations as well as the Proposition 5.7 the main term $M_{1}$ can be written as:

$$
\begin{aligned}
M_{1}= & \int_{0}^{T} \int_{\Omega} \theta(t, x)\left\{E_{1} \varepsilon_{0} \operatorname{div} E+\varepsilon_{0} B_{2}\left(\alpha E_{3}+\partial_{t} E_{3}-c_{0}^{2}(\operatorname{rot} B)_{3}\right)\right. \\
& \left.-\varepsilon_{0} B_{3}\left(\alpha E_{2}+\partial_{t} E_{2}-c_{0}^{2}(\operatorname{rot} B)_{2}\right)\right\} d t d x \\
= & \int_{0}^{T} \int_{\Omega} \theta(t, x)\left\{\varepsilon_{0}\left(E_{1} \operatorname{div} E-(E \wedge \operatorname{rot} E)_{1}\right)+\frac{1}{\mu_{0}}\left(B_{1} \operatorname{div} B-(B \wedge \operatorname{rot} B)_{1}\right)\right. \\
& \left.-2 \alpha \varepsilon_{0}(E \wedge B)_{1}\right\} d t d x \\
& +\int_{0}^{T} \int_{\Omega} \varepsilon_{0}(E \wedge B)_{1} \partial_{t} \theta d t d x .
\end{aligned}
$$

Finally we obtain for $1 \leq i \leq 3$ that:
$\int_{0}^{T} \int_{\Omega}\left\{\varepsilon_{0}\left(E_{i} \operatorname{div} E-(E \wedge \operatorname{rot} E)_{i}\right)+\frac{1}{\mu_{0}}\left(B_{i} \operatorname{div} B-(B \wedge \operatorname{rot} B)_{i}\right)\right\} \theta(t, x) d t d x$

$$
+\int_{0}^{T} \int_{\Omega} \varepsilon_{0}(E \wedge B)_{i} \partial_{t} \theta d t d x
$$

$$
+\int_{0}^{T} \int_{\Omega}\left(\int_{\mathbb{R}_{p}^{3}} p_{i} f d p\right) \partial_{t} \theta d t d x+\int_{0}^{T} \int_{\Omega}\left(\int_{\mathbb{R}_{p}^{3}} p_{i} v(p) f d p\right) \cdot \nabla_{x} \theta d t d x
$$

$$
=\alpha \int_{0}^{T} \int_{\Omega}\left(2 \varepsilon_{0}(E \wedge B)_{i}+\int_{\mathbb{R}_{p}^{3}} p_{i} f d p\right) \theta(t, x) d t d x
$$

$$
+\int_{0}^{T} \int_{\partial \Omega} \theta(t, x)\left(\int_{\mathbb{R}_{p}^{3}}(v(p) \cdot n(x)) p_{i} \gamma f d p\right) d t d \sigma-R_{i}(\theta)
$$

for all $\theta \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right), T$ periodic. Let us introduce the notations:

$$
\begin{gathered}
v_{i}=2 \alpha \varepsilon_{0}(E \wedge B)_{i}+\alpha \int_{\mathbb{R}_{p}^{3}} p_{i} f d p \in L^{1}(] 0, T[\times \Omega), 1 \leq i \leq 3, \\
w_{i}=\int_{\mathbb{R}_{p}^{3}} p_{i} v(p) f d p \in L^{1}(] 0, T\left[; L^{1}(\Omega)^{3}\right), 1 \leq i \leq 3 \\
n \cdot w_{i}=\int_{\mathbb{R}_{p}^{3}}(v(p) \cdot n(x)) p_{i} \gamma f d p \in L^{1}(] 0, T[\times \partial \Omega), 1 \leq i \leq 3, \\
z_{i}=\varepsilon_{0}(E \wedge B)_{i}+\int_{\mathbb{R}_{p}^{3}} p_{i} f d p \in L^{1}(] 0, T[\times \Omega), 1 \leq i \leq 3 .
\end{gathered}
$$

Then we have for $1 \leq i \leq 3, \theta \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right), T$ periodic:

$$
\begin{align*}
\int_{0}^{T} \int_{\Omega}\left\{\varepsilon _ { 0 } \left(E_{i} \operatorname{div} E-\right.\right. & \left.\left.(E \wedge \operatorname{rot} E)_{i}\right)+\frac{1}{\mu_{0}}\left(B_{i} \operatorname{div} B-(B \wedge \operatorname{rot} B)_{i}\right)\right\} \theta(t, x) d t d x \\
& +\int_{0}^{T} \int_{\Omega}\left\{z_{i}(t, x) \partial_{t} \theta+w_{i}(t, x) \cdot \nabla_{x} \theta\right\} d t d x \\
= & \int_{0}^{T} \int_{\Omega} v_{i}(t, x) \theta(t, x) d t d x+\int_{0}^{T} \int_{\partial \Omega}\left(n \cdot w_{i}\right) \theta(t, x) d t d \sigma-R_{i}(\theta) \tag{6.10}
\end{align*}
$$

We can not apply directly the Lemma 6.4 since we have the extra term $R_{i}$ in (6.10). By performing the same computations as in the proof of Lemma 6.3 we deduce that there is a constant $C$ such that:

$$
\begin{aligned}
\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} & (p \cdot v(p)) f d t d x d p+\int_{0}^{T} \int_{\Omega}\left\{\varepsilon_{0}|E|^{2}+\frac{1}{\mu_{0}}|B|^{2}\right\} d t d x \\
& +\int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}(n \cdot E)^{2}+\frac{1}{\mu_{0}}(n \cdot B)^{2}\right\} d t d \sigma \\
\leq & C\left\{\int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}|n \wedge E|^{2}+\frac{1}{\mu_{0}}|n \wedge B|^{2}\right\} d t d \sigma+\alpha \varepsilon_{0} \int_{0}^{T} \int_{\Omega}|E \wedge B| d t d x\right.
\end{aligned}
$$

$$
\begin{equation*}
\left.+\alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}|p| f d t d x d p+\int_{0}^{T} \int_{\Sigma}|(v(p) \cdot n(x))| \cdot|p| \cdot \gamma f d t d \sigma d p+\sum_{i=1}^{3}\left|R_{i}\right|\right\} \tag{6.11}
\end{equation*}
$$

where $R_{i}$ is the corresponding term to the test function $\theta(x)=x_{i}$, for example:

$$
\begin{aligned}
R_{1}= & \int_{0}^{T} \int_{\Omega} E_{1}(t, x) \int_{0}^{T} \int_{\Omega}\left(y_{1}-x_{1}\right) \rho(s, y) \zeta_{\varepsilon}^{\sqrt{\prime}}(t-s, x-y) d s d y d t d x \\
& +\int_{0}^{T} \int_{\Omega} B_{3}(t, x) \int_{0}^{T} \int_{\Omega}\left(y_{1}-x_{1}\right) j_{2}(s, y) \zeta_{\varepsilon}^{\sqrt{\prime}}(t-s, x-y) d s d y d t d x \\
& -\int_{0}^{T} \int_{\Omega} B_{2}(t, x) \int_{0}^{T} \int_{\Omega}\left(y_{1}-x_{1}\right) j_{3}(s, y) \zeta_{\varepsilon}^{\sqrt{\prime}}(t-s, x-y) d s d y d t d x
\end{aligned}
$$

Note that from the Proposition 5.4 and the Theorem 5.8 we already have an uniform estimate for the tangential traces $(n \wedge E, n \wedge B)$ in $L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)$. On the other hand, by using one more time (5.3), (5.11) one gets:

$$
\begin{align*}
2 \alpha \varepsilon_{0} c_{0} \int_{0}^{T} \int_{\Omega}|E \wedge B| d t d x & \leq 2 \alpha \sqrt{\frac{\varepsilon_{0}}{\mu_{0}}} \int_{0}^{T} \int_{\Omega}|E(t, x)| \cdot|B(t, x)| d t d x \\
& \leq \alpha \int_{0}^{T} \int_{\Omega}\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d t d x \\
\leq & \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| \mathcal{E}(p) g(t, x, p) d t d \sigma d p \\
& +\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma . \tag{6.12}
\end{align*}
$$

We have also:

$$
\begin{align*}
& \alpha \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}|p| f(t, x, p) d t d x d p \leq \alpha C \cdot \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}(1+\mathcal{E}(p)) f(t, x, p) d t d x d p \\
& \leq C \cdot\left(\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))|(1+\mathcal{E}(p)) g(t, x, p) d t d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma\right), \tag{6.13}
\end{align*}
$$

and:

$$
\begin{align*}
& \int_{0}^{T} \int_{\Sigma} \quad|(v(p) \cdot n(x))| \cdot|p| \gamma f(t, x, p) d t d \sigma d p \\
& \leq C \cdot \int_{0}^{T} \int_{\Sigma}|(v(p) \cdot n(x))|(1+\mathcal{E}(p)) \gamma f(t, x, p) d t d \sigma d p \\
& \leq C \cdot\left(\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))|(1+\mathcal{E}(p)) g(t, x, p) d t d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma\right) \tag{6.14}
\end{align*}
$$

Now we need to estimate the term $R_{i}, 1 \leq i \leq 3$. We have:

$$
\left|R_{1}\right| \leq \varepsilon_{2} \int_{0}^{T} \int_{\Omega}\left\{\left(\left|\bar{E}_{1}\right| \star \zeta_{\varepsilon}\right) \cdot|\rho|+\left(\left|\bar{B}_{3}\right| \star \zeta_{\varepsilon}\right) \cdot\left|j_{2}\right|+\left(\left|\bar{B}_{2}\right| \star \zeta_{\varepsilon}\right) \cdot\left|j_{3}\right|\right\} d t d x .
$$

Note that for both classical or relativistic case we have:

$$
\begin{aligned}
|\rho(t, x)| & =|q| \int_{\mathbb{R}_{p}^{3}} f \cdot \mathbf{1}_{\{|p| \leq R\}} d p+|q| \int_{\mathbb{R}_{p}^{3}} f \cdot \mathbf{1}_{\{|p|>R\}} d p \leq C \cdot R^{3}\|f\|_{\infty}+\frac{C}{R} \int_{\mathbb{R}_{p}^{3}}(1+\mathcal{E}(p)) f d p \\
& \leq C \cdot\|f\|_{\infty}^{1 / 4} \cdot\left(\int_{\mathbb{R}_{p}^{3}}(1+\mathcal{E}(p)) f d p\right)^{3 / 4}
\end{aligned}
$$

and thus $\|\rho\|_{L^{\frac{4}{3}}(j 0, T[\times \Omega)} \leq C \cdot\|f\|_{\infty}^{1 / 4} \cdot\left(\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}(1+\mathcal{E}(p)) f d t d x d p\right)^{3 / 4} \leq \frac{C}{\alpha^{3 / 4}}$, where for the last inequality we used (5.3), (5.11). Thus, by using the Hölder and Young inequalities and (5.3), (5.11) we deduce that:

$$
\begin{aligned}
\int_{0}^{T} \int_{\Omega}\left(\left|\bar{E}_{1}\right| \star \zeta_{\varepsilon}\right) \cdot|\rho(t, x)| d t d x & \leq\left\|\left(\left|\bar{E}_{1}\right| \star \zeta_{\varepsilon}\right)\right\|_{L^{4}} \cdot\|\rho\|_{L^{\frac{4}{3}}} \leq\left\|\bar{E}_{1}\right\|_{L^{2}} \cdot\left\|\zeta_{\varepsilon}\right\|_{L^{\frac{4}{3}}} \cdot\|\rho\|_{L^{\frac{4}{3}}(00, T[\times \Omega)} \\
& \leq \frac{C}{\alpha^{1 / 2} \cdot \varepsilon_{1}^{1 / 4} \cdot \varepsilon_{2}^{3 / 4}}\left\|\zeta_{3}\right\|_{L^{\frac{4}{3}}} \cdot\left\|\zeta_{2}\right\|_{L^{\frac{4}{3}}} \cdot \frac{C}{\alpha^{3 / 4}} \\
& \leq \frac{C}{\alpha^{5 / 4} \cdot \varepsilon_{1}^{1 / 4} \cdot \varepsilon_{2}^{3 / 4}} .
\end{aligned}
$$

In the relativistic case we have also $\|j\|_{L^{\frac{4}{3}(] 0, T[\times \Omega)}} \leq C \cdot \alpha^{-3 / 4}$ and thus we obtain similar bounds for the terms $\int_{0}^{T} \int_{\Omega}\left(\left|\bar{B}_{3}\right| \star \zeta_{\varepsilon}\right) \cdot\left|j_{2}\right| d t d x, \int_{0}^{T} \int_{\Omega}\left(\left|\bar{B}_{2}\right| \star \zeta_{\varepsilon}\right) \cdot\left|j_{3}\right| d t d x$, which implies that $\left|R_{1}\right| \leq C \cdot \alpha^{-5 / 4} \cdot \varepsilon_{2}^{1 / 4} \cdot \varepsilon_{1}^{-1 / 4}$. In the classical case we can estimate $j$ by interpolation:

$$
\begin{aligned}
|j(t, x)| \leq & \leq q \left\lvert\, \int_{\mathbb{R}_{p}^{3}} \frac{|p|}{m} f d p \leq C \cdot\|f\|_{\infty} \cdot R^{4}+\frac{2|q|}{R} \int_{\mathbb{R}_{p}^{3}} \frac{|p|^{2}}{2 m} f(t, x, p) d p\right. \\
& \leq C \cdot\|f\|_{\infty}^{1 / 5} \cdot\left(\int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f d p\right)^{4 / 5}
\end{aligned}
$$

and thus $\|j\|_{L^{\frac{5}{4}}(j 0, T[\times \Omega)} \leq C \cdot\|f\|_{\infty}^{1 / 5} \cdot\left(\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f d t d x d p\right)^{4 / 5} \leq C \cdot \alpha^{-4 / 5}$. Now, by using the Hölder and Young inequalities we deduce that:

$$
\begin{aligned}
\int_{0}^{T} \int_{\Omega}\left(\left|\bar{B}_{3}\right| \star \zeta_{\varepsilon}\right) \cdot\left|j_{2}(t, x)\right| d t d x & \leq\left\|\left(\left|\bar{B}_{3}\right| \star \zeta_{\varepsilon}\right)\right\|_{L^{5}} \cdot\|j\|_{L^{\frac{5}{4}}} \leq\left\|\bar{B}_{3}\right\|_{L^{2}} \cdot\left\|\zeta_{\varepsilon}\right\|_{L^{\frac{10}{7}}} \cdot\|j\|_{L^{\frac{5}{4}}(] 0, T[\times \Omega)} \\
& \leq \frac{C}{\alpha^{1 / 2} \cdot \varepsilon_{1}^{3 / 10} \cdot \varepsilon_{2} 9 / 10}\left\|\zeta_{3}\right\|_{L^{\frac{10}{7}}} \cdot\left\|\zeta_{2}\right\|_{L^{\frac{10}{7}}} \cdot \frac{C}{\alpha^{4 / 5}} \\
& \leq \frac{C}{\alpha^{13 / 10} \cdot \varepsilon_{1}^{3 / 10} \cdot \varepsilon_{2}^{9 / 10}}
\end{aligned}
$$

Therefore, in the classical case we have $\left|R_{1}\right| \leq C \cdot \alpha^{-5 / 4} \cdot \varepsilon_{2}^{1 / 4} \cdot \varepsilon_{1}^{-1 / 4}+C \cdot \alpha^{-13 / 10}$. $\varepsilon_{2}^{1 / 10} \cdot \varepsilon_{1}^{-3 / 10}$. The conclusion follows easily by observing that $(p \cdot v(p)) \geq \mathcal{E}(p), \forall p \in \mathbb{R}_{p}^{3}$ and by combining (5.3), (5.11), (6.11), (6.12), (6.13), (6.14) we can take $q=\frac{13}{10}, r=$
$\frac{3}{10}, s=\frac{1}{10}$. We deduce that there is $\left.t_{0} \in\right] 0, T[$ such that:

$$
\begin{aligned}
& \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f\left(t_{0}, x, p\right) d x d p+\int_{\Omega}\left\{\varepsilon_{0}\left|E\left(t_{0}, x\right)\right|^{2}+\frac{1}{\mu_{0}}\left|B\left(t_{0}, x\right)\right|^{2}\right\} d x \\
\leq & \frac{C}{T}\left\{\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))|(1+\mathcal{E}(p)) g(t, x, p) d t d \sigma d p+\int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma\right\} \\
& +\frac{C}{T} \frac{\varepsilon_{2}^{s}}{\alpha^{q} \cdot \varepsilon_{1}^{r}} .
\end{aligned}
$$

By using the balance of the total energy (see the proofs of Proposition 5.6, especially formula (5.6) and Theorem 5.8) we deduce that for $t \in] t_{0}, t_{0}+T[$ we have:

$$
\begin{array}{rl}
\int_{\Omega} \int_{\mathbb{R}_{p}^{3}} & \mathcal{E}(p) f(t, x, p) d x d p+\frac{1}{2} \int_{\Omega}\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d x \\
& \leq \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f\left(t_{0}, x, p\right) d x d p+\frac{1}{2} \int_{\Omega}\left\{\varepsilon_{0}\left|E\left(t_{0}, x\right)\right|^{2}+\frac{1}{\mu_{0}}\left|B\left(t_{0}, x\right)\right|^{2}\right\} d x \\
& +\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| \mathcal{E}(p) g(s, x, p) d s d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(s, x)|^{2} d s d \sigma
\end{array}
$$

and the last conclusion follows by time periodicity.

## 7. The Vlasov-Maxwell system

We are ready to prove Theorem 1.1 (existence of $T$ periodic weak solution for the three dimensional Vlasov-Maxwell boundary-value problem).

Proof. (Theorem 1.1) We take $\eta>0$ a small parameter and we consider $0<$ $\alpha, \varepsilon_{1}, \varepsilon_{2}<1$ such that $\alpha^{q}=\eta^{1 / 4}, \varepsilon_{1}^{r}=\eta^{1 / 4}$ and $\varepsilon_{2}^{s}=\eta$ where $q, r, s>0$ are given in Proposition 6.1. We regularize also the boundary data $h$ by taking $h_{\eta} \rightarrow h$ in $L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)$ such that there is $\tilde{h}_{\eta} \in C^{1}\left(\mathbb{R}_{t} ; H^{1}(\Omega)^{3}\right) \cap C^{2}\left(\mathbb{R}_{t} ; L^{2}(\Omega)^{3}\right) T$ periodic with $\left.n \wedge \tilde{h}_{\eta}\right|_{\mathbb{R}_{t} \times \partial \Omega}=h_{\eta}$. From the Theorems $5.3,5.8$ applied to the boundary data $\left(g, h_{\eta}\right)$ we deduce the existence of a $T$ periodic solution for the perturbed VlasovMaxwell system $\left(f_{\eta} \geq 0, E_{\eta}, B_{\eta}\right)$. By the Proposition 6.1 we deduce that for all $s \in \mathbb{R}_{t}$ :

$$
\begin{array}{rl}
\int_{\Omega} \int_{\mathbb{R}_{p}^{3}} & \mathcal{E}(p) f_{\eta}(s, x, p) d x d p+\int_{\Omega}\left\{\varepsilon_{0}\left|E_{\eta}(s, x)\right|^{2}+\frac{1}{\mu_{0}}\left|B_{\eta}(s, x)\right|^{2}\right\} d x \\
& +\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x))(1+\mathcal{E}(p)) \gamma^{+} f_{\eta}(t, x, p) d t d \sigma d p \\
& +\int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}\left[\left(n \cdot E_{\eta}\right)^{2}+\left|n \wedge E_{\eta}\right|^{2}\right]+\frac{1}{\mu_{0}}\left[\left(n \cdot B_{\eta}\right)^{2}+\left|n \wedge B_{\eta}\right|^{2}\right]\right\} d t d \sigma \\
& \leq C \cdot W_{0, \eta}+C \cdot \eta^{1 / 2}, \tag{7.1}
\end{array}
$$

where $W_{0, \eta}=\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))|(1+\mathcal{E}(p)) g(t, x, p) d t d \sigma d p+\int_{0}^{T} \int_{\partial \Omega}\left|h_{\eta}(t, x)\right|^{2} d t d \sigma$. After extraction of subsequences we can suppose that there is $f, \gamma^{+} f, E, B,(n$. $E),(n \cdot B), n \wedge E, n \wedge B$ such that $f_{k}:=f_{\eta_{k}} \rightharpoonup f$ weakly $\star$ in $L^{\infty}\left(\mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}\right)$, $\gamma^{+} f_{k}:=\gamma^{+} f_{\eta_{k}} \rightharpoonup \gamma^{+} f$ weakly $\star$ in $L^{\infty}\left(\mathbb{R}_{t} \times \Sigma^{+}\right), \quad\left(E_{k}, B_{k}\right):=\left(E_{\eta_{k}}, B_{\eta_{k}}\right) \rightharpoonup(E, B)$ weakly in $L^{2}(] 0, T\left[; L^{2}(\Omega)^{3}\right)^{2},\left(n \cdot E_{k}, n \cdot B_{k}\right):=\left(n \cdot E_{\eta_{k}}, n \cdot B_{\eta_{k}}\right) \rightharpoonup(n \cdot E, n \cdot B)$ weakly in $L^{2}(] 0, T[\times \partial \Omega)^{2}, \quad\left(n \wedge E_{k}, n \wedge B_{k}\right):=\left(n \wedge E_{\eta_{k}}, n \wedge B_{\eta_{k}}\right) \rightharpoonup(n \wedge E, n \wedge B)$ weakly in $L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)^{2}$, where $\eta_{k} \rightarrow 0$, as $k \rightarrow+\infty$. We note also by $\alpha_{k}, \varepsilon_{1, k}, \varepsilon_{2, k}$ the
values given by $\alpha_{k}^{q}=\eta_{k}^{1 / 4}, \varepsilon_{1, k}^{r}=\eta_{k}^{1 / 4}, \varepsilon_{2, k}^{s}=\eta_{k}$. Obviously, by weak $\star$ convergence we have:

$$
\max \left(\|f\|_{\infty},\left\|\gamma^{+} f\right\|_{\infty}\right) \leq \max \left(\liminf _{k \rightarrow+\infty}\left\|f_{k}\right\|_{\infty}, \liminf _{k \rightarrow+\infty}\left\|\gamma^{+} f_{k}\right\|_{\infty}\right) \leq\|g\|_{\infty}
$$

For $R>0, \theta \in C^{0}\left(\mathbb{R}_{t}\right), \theta \geq 0$ we have:

$$
\begin{aligned}
& \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \theta(t) \mathcal{E}(p) \cdot \mathbf{1}_{\{|p| \leq R\}} f(t, x, p) d t d x d p \\
& +\int_{0}^{T} \int_{\Omega} \theta(t)\left\{\varepsilon_{0}|E(t, x)|^{2}+\frac{1}{\mu_{0}}|B(t, x)|^{2}\right\} d t d x \\
\leq & \liminf _{k \rightarrow+\infty}\left\{\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \theta(t) \mathcal{E}(p) \cdot \mathbf{1}_{\{|p| \leq R\}} f_{k} d t d x d p+\int_{0}^{T} \int_{\Omega} \theta(t)\left\{\varepsilon_{0}\left|E_{k}\right|^{2}+\frac{1}{\mu_{0}}\left|B_{k}\right|^{2}\right\} d t d x\right\} \\
\leq & \liminf _{k \rightarrow+\infty} \int_{0}^{T} \theta(t)\left(C \cdot W_{0, \eta_{k}}+C \cdot \eta_{k}^{1 / 2}\right) d t \\
= & C \cdot W_{0} \int_{0}^{T} \theta(t) d t
\end{aligned}
$$

By letting $R \rightarrow+\infty$ we deduce that:

$$
\int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f(s, x, p) d x d p+\int_{\Omega}\left\{\varepsilon_{0}|E(s, x)|^{2}+\frac{1}{\mu_{0}}|B(s, x)|^{2}\right\} d x \leq C \cdot W_{0}, \text { a.e. } s \in \mathbb{R}_{t}
$$

Similarly we have:

$$
\begin{aligned}
\int_{0}^{T} \int_{\Sigma^{+}} & (v(p) \cdot n(x))(1+\mathcal{E}(p)) \gamma^{+} f(t, x, p) \mathbf{1}_{\{|p| \leq R\}} d t d \sigma d p \\
& +\int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}\left[(n \cdot E)^{2}+|n \wedge E|^{2}\right]+\frac{1}{\mu_{0}}\left[(n \cdot B)^{2}+|n \wedge B|^{2}\right]\right\} d t d \sigma \\
\leq & \liminf _{k \rightarrow+\infty}\left(\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x))(1+\mathcal{E}(p)) \gamma^{+} f_{k}(t, x, p) \mathbf{1}_{\{|p| \leq R\}} d t d \sigma d p\right. \\
& \left.+\int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}\left[\left(n \cdot E_{k}\right)^{2}+\left|n \wedge E_{k}\right|^{2}\right]+\frac{1}{\mu_{0}}\left[\left(n \cdot B_{k}\right)^{2}+\left|n \wedge B_{k}\right|^{2}\right]\right\} d t d \sigma\right) \\
\leq & C \cdot W_{0}, \forall R>0
\end{aligned}
$$

We check that $(f, E, B)$ is a $T$ periodic solution of the Vlasov-Maxwell system. Indeed, by the Green formula we have for all functions $\varphi \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega} \times \mathbb{R}_{p}^{3}\right) T$ periodic, with compact support in momentum:

$$
\begin{aligned}
-\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} & f_{k}\left(-\alpha_{k} \varphi(t, x, p)+\partial_{t} \varphi+v(p) \cdot \nabla_{x} \varphi+q\left(\left(\bar{E}_{k} \star \zeta_{\varepsilon_{k}}\right)\right.\right. \\
& \left.\left.+v(p) \wedge\left(\bar{B}_{k} \star \zeta_{\varepsilon_{k}}\right)\right) \cdot \nabla_{p} \varphi\right) d t d x d p \\
= & -\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f_{k} \varphi d t d \sigma d p-\int_{0}^{T} \int_{\Sigma^{-}}(v(p) \cdot n(x)) g \varphi d t d \sigma d p
\end{aligned}
$$

Observe that $\alpha_{k} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \varphi f_{k} d t d x d p \rightarrow 0$ since $\left\|f_{k}\right\|_{\infty} \leq\|g\|_{\infty}$, and $\left(\bar{E}_{k} \star \zeta_{\varepsilon_{k}}, \bar{B}_{k} \star\right.$ $\left.\zeta_{\varepsilon_{k}}\right) \rightharpoonup(E, B)$ weakly in $L^{2}(] 0, T\left[; L^{2}(\Omega)^{3}\right)$, as $k \rightarrow+\infty$. Note that the compactness
average result of DiPerna and Lions [14] adapts easily in the time periodic case and still holds true for bounded spatial domains (use a cut-off function $\eta \in C_{c}^{\infty}(\Omega)$ and write $f_{k}=\eta f_{k}+(1-\eta) f_{k}$ as it was done in [23] page 256). For test functions $\varphi(t, x, p)=\varphi_{1}(t, x) \cdot \varphi_{2}(p), \varphi_{1} \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right) T$ periodic, $\varphi_{2} \in C_{c}^{1}\left(\mathbb{R}_{p}^{3}\right)$, by using the velocity average result we have:

$$
\begin{gathered}
\lim _{k \rightarrow+\infty} \int_{\mathbb{R}_{p}^{3}} \nabla_{p} \varphi_{2} f_{k} d p=\int_{\mathbb{R}_{p}^{3}} \nabla_{p} \varphi_{2} f d p, \text { in } L^{2}(] 0, T\left[; L^{2}(\Omega)^{3}\right) \\
\lim _{k \rightarrow+\infty} \int_{\mathbb{R}_{p}^{3}}\left(\nabla_{p} \varphi_{2} \wedge v(p)\right) f_{k} d p=\int_{\mathbb{R}_{p}^{3}}\left(\nabla_{p} \varphi_{2} \wedge v(p)\right) f d p, \text { in } L^{2}(] 0, T\left[; L^{2}(\Omega)^{3}\right),
\end{gathered}
$$

and thus by combining strong and weak convergence we can pass to the limit in the nonlinear term:

$$
\begin{align*}
\lim _{k \rightarrow+\infty} & \int_{0}^{T} \int_{\Omega}\left\{\varphi_{1}\left(\bar{E}_{k} \star \zeta_{\varepsilon_{k}}\right) \cdot \int_{\mathbb{R}_{p}^{3}} \nabla_{p} \varphi_{2} f_{k} d p+\varphi_{1}\left(\bar{B}_{k} \star \zeta_{\varepsilon_{k}}\right) \cdot \int_{\mathbb{R}_{p}^{3}}\left(\nabla_{p} \varphi_{2} \wedge v(p)\right) f_{k} d p\right\} d t d x \\
= & \int_{0}^{T} \int_{\Omega}\left\{\varphi_{1} E \cdot \int_{\mathbb{R}_{p}^{3}} \nabla_{p} \varphi_{2} f d p+\varphi_{1} B \cdot \int_{\mathbb{R}_{p}^{3}}\left(\nabla_{p} \varphi_{2} \wedge v(p)\right) f d p\right\} d t d x \tag{7.2}
\end{align*}
$$

Finally we deduce that:

$$
\begin{aligned}
& -\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f(t, x, p)\left(\partial_{t} \varphi+v(p) \cdot \nabla_{x} \varphi+q(E(t, x)+v(p) \wedge B(t, x)) \cdot \nabla_{p} \varphi\right) d t d x d p \\
& \quad=-\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f \varphi d t d \sigma d p-\int_{0}^{T} \int_{\Sigma^{-}}(v(p) \cdot n(x)) g \varphi d t d \sigma d p, \forall \varphi=\varphi_{1} \cdot \varphi_{2}
\end{aligned}
$$

and by density the previous equality holds for all $\varphi \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega} \times \mathbb{R}_{p}^{3}\right) T$ periodic and compactly supported in momentum. Now take $\varphi, \psi \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right)^{3}, T$ periodic with $n \wedge \varphi-c_{0} n \wedge(n \wedge \psi)=0$ on $\mathbb{R}_{t} \times \partial \Omega$. We have:

$$
\begin{gathered}
\int_{0}^{T} \int_{\Omega}\left\{E_{k} \cdot\left(\alpha_{k} \varphi-\partial_{t} \varphi\right)-c_{0}^{2} B_{k} \cdot \operatorname{rot} \varphi\right\} d t d x+c_{0}^{2} \int_{0}^{T} \int_{\Omega}\left\{B_{k} \cdot\left(\alpha_{k} \psi-\partial_{t} \psi\right)+E_{k} \cdot \operatorname{rot} \psi\right\} d t d x \\
=c_{0} \int_{0}^{T} \int_{\partial \Omega}(n \wedge \varphi) \cdot h_{k} d t d \sigma-\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega}\left(\bar{j}_{k} \star \zeta_{\varepsilon_{k}}^{\vee}\right) \cdot \varphi d t d x
\end{gathered}
$$

Since $\left(E_{k}, B_{k}\right)_{k}$ is bounded in $L^{2}(] 0, T\left[; L^{2}(\Omega)^{3}\right)$ we have:

$$
\lim _{k \rightarrow+\infty} \alpha_{k} \int_{0}^{T} \int_{\Omega} E_{k}(t, x) \cdot \varphi(t, x) d t d x=\lim _{k \rightarrow+\infty} \alpha_{k} \int_{0}^{T} \int_{\Omega} B_{k}(t, x) \cdot \psi(t, x) d t d x=0
$$

We have also:

$$
\begin{aligned}
\left|\int_{0}^{T} \int_{\Omega}\left(\bar{j}_{k} \star \zeta_{\varepsilon_{k}}^{\sqrt{l}}\right) \cdot \varphi d t d x-\int_{0}^{T} \int_{\Omega} j \cdot \varphi d t d x\right| \leq & \left|\int_{0}^{T} \int_{\Omega} j_{k} \cdot\left(\bar{\varphi} \star \zeta_{\varepsilon_{k}}-\varphi\right) d t d x\right| \\
& +\left|\int_{0}^{T} \int_{\Omega}\left(j_{k}-j\right) \cdot \varphi d t d x\right|=I_{1}^{k}+I_{2}^{k}
\end{aligned}
$$

For the first term $I_{1}^{k}$ we write:

$$
\begin{aligned}
|q|^{-1} \cdot I_{1}^{k} \leq & \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}|v(p)| \cdot f_{k}(t, x, p) \cdot\left|\left(\bar{\varphi} \star \zeta_{\varepsilon_{k}}\right)(t, x)-\varphi(t, x)\right| \cdot \mathbf{1}_{\{|p| \leq R\}} d t d x d p \\
& +\frac{1}{R} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}|v(p)| \cdot|p| \cdot f_{k}(t, x, p) \cdot 2 \cdot\|\varphi\|_{\infty} \cdot \mathbf{1}_{\{|p|>R\}} d t d x d p \\
\leq & \int_{0}^{T} \int_{\Omega}\left|\left(\bar{\varphi} \star \zeta_{\varepsilon_{k}}\right)(t, x)-\varphi(t, x)\right|\left(\int_{\mathbb{R}_{p}^{3}}|v(p)| \cdot\|g\|_{\infty} \cdot \mathbf{1}_{\{|p| \leq R\}} d p\right) d t d x \\
& +\frac{4\|\varphi\|_{\infty}}{R} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f_{k}(t, x, p) d t d x d p
\end{aligned}
$$

For $\delta>0$ we take $R_{\delta}$ large enough such that $\frac{4\|\varphi\|_{\infty}}{R_{\delta}} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f_{k}(t, x, p) d t d x d p<\frac{\delta}{2}$, uniformly in $k$. By using the dominated convergence theorem we have also:

$$
\lim _{k \rightarrow+\infty} \int_{0}^{T} \int_{\Omega}\left|\left(\bar{\varphi} \star \zeta_{\varepsilon_{k}}\right)(t, x)-\varphi(t, x)\right|\left(\int_{\mathbb{R}_{p}^{3}}|v(p)| \cdot\|g\|_{\infty} \cdot \mathbf{1}_{\left\{|p| \leq R_{\delta}\right\}} d p\right) d t d x=0
$$

and therefore $\lim _{k \rightarrow+\infty} I_{1}^{k}=0$. For the second term we have:

$$
\begin{aligned}
& |q|^{-1} I_{2}^{k} \leq\left|\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} v(p) \varphi\left(f_{k}-f\right) \mathbf{1}_{\{|p| \leq R\}} d t d x d p\right| \\
+ & \left|\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} v(p) \varphi\left(f_{k}-f\right) \mathbf{1}_{\{|p|>R\}} d t d x d p\right|=I_{3}^{k}+I_{4}^{k}
\end{aligned}
$$

By using the inequality $|p| \cdot|v(p)|=(p \cdot v(p)) \leq 2 \mathcal{E}(p)$ we have for $R$ large enough:

$$
\begin{aligned}
I_{4}^{k} & \leq \frac{1}{R} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f_{k} \cdot\|\varphi\|_{\infty} \cdot|p| \cdot|v(p)| d t d x d p+\frac{1}{R} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f \cdot\|\varphi\|_{\infty} \cdot|p| \cdot|v(p)| d t d x d p \\
& \leq \frac{2\|\varphi\|_{\infty}}{R} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f_{k} \mathcal{E}(p) d t d x d p+\frac{2\|\varphi\|_{\infty}}{R} \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f \mathcal{E}(p) d t d x d p<\frac{\delta}{2}
\end{aligned}
$$

for $R \geq R_{\delta}$, since $\left(\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \mathcal{E}(p) f_{k} d t d x d p\right)_{k}$ is bounded. Now, for $R=R_{\delta}$ we can use the weak $\star$ convergence $f_{k} \rightharpoonup f$ in order to obtain $I_{3}^{k}<\frac{\delta}{2}$, for $k \geq k_{\delta}$. Finally we proved the convergence $\lim _{k \rightarrow+\infty} \int_{0}^{T} \int_{\Omega}\left(\bar{j}_{k} \star \zeta_{\varepsilon_{k}}^{\sqrt{ }}\right) \cdot \varphi d t d x=\int_{0}^{T} \int_{\Omega} j \cdot \varphi d t d x$. By using the weak convergence $\left(E_{k}, B_{k}\right) \rightharpoonup(E, B)$ in $L^{2}(] 0, T\left[; L^{2}(\Omega)^{3}\right)^{2}$ and $h_{k} \rightarrow h$ in $L^{2}(] 0, T\left[; L^{2}(\partial \Omega)^{3}\right)$ we deduce that:

$$
\begin{aligned}
& \int_{0}^{T} \int_{\Omega}\left\{-E(t, x) \cdot \partial_{t} \varphi-c_{0}^{2} B(t, x) \cdot \operatorname{rot} \varphi\right\} d t d x \\
& +c_{0}^{2} \int_{0}^{T} \int_{\Omega}\left\{-B(t, x) \cdot \partial_{t} \psi+E(t, x) \cdot \operatorname{rot} \psi\right\} d t d x \\
= & c_{0} \int_{0}^{T} \int_{\partial \Omega}(n \wedge \varphi) \cdot h(t, x) d t d \sigma-\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega} j(t, x) \cdot \varphi(t, x) d t d x .
\end{aligned}
$$

In fact, by using the Remark 4.7, we can prove that:

$$
\begin{array}{r}
\int_{0}^{T} \int_{\Omega}\left\{-E(t, x) \cdot \partial_{t} \varphi-c_{0}^{2} B(t, x) \cdot \operatorname{rot} \varphi\right\} d t d x-c_{0}^{2} \int_{0}^{T} \int_{\partial \Omega}(n \wedge B) \cdot \varphi(t, x) d t d \sigma \\
=-\frac{1}{\varepsilon_{0}} \int_{0}^{T} \int_{\Omega} j \cdot \varphi d t d x
\end{array}
$$

and:

$$
\int_{0}^{T} \int_{\Omega}\left\{-B(t, x) \cdot \partial_{t} \psi+E(t, x) \cdot \operatorname{rot} \psi\right\} d t d x+\int_{0}^{T} \int_{\partial \Omega}(n \wedge E) \cdot \psi(t, x) d t d \sigma=0
$$

for all $\varphi, \psi \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right)^{3}, T$ periodic. By the Proposition 6.1 and by using that $\operatorname{div} E_{k}=\frac{1}{\varepsilon_{0}}\left(\bar{\rho}_{k} \star \zeta_{\varepsilon_{k}}^{\sqrt{ }}\right)$ we have for $\theta \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right)$ :

$$
\int_{0}^{T} \int_{\Omega} \frac{1}{\varepsilon_{0}}\left(\bar{\rho}_{k} \star \zeta_{\varepsilon_{k}}^{\sqrt{ }}\right) \theta(t, x) d t d x+\int_{0}^{T} \int_{\Omega} E_{k}(t, x) \cdot \nabla_{x} \theta d t d x=\int_{0}^{T} \int_{\partial \Omega}\left(n \cdot E_{k}\right) \theta(t, x) d t d \sigma
$$

As previous we have $\lim _{k \rightarrow+\infty} \int_{0}^{T} \int_{\Omega}\left(\bar{\rho}_{k} \star \zeta_{\varepsilon_{k}}^{\checkmark}\right) \theta d t d x=\int_{0}^{T} \int_{\Omega} \rho \theta d t d x$ and thus, by passing to the limit for $k \rightarrow+\infty$ one gets:

$$
\int_{0}^{T} \int_{\Omega} \frac{1}{\varepsilon_{0}} \rho \theta d t d x+\int_{0}^{T} \int_{\Omega} E \cdot \nabla_{x} \theta d t d x=\int_{0}^{T} \int_{\partial \Omega}(n \cdot E) \theta d t d \sigma
$$

or div $E=\frac{1}{\varepsilon_{0}} \rho$ and $(n \cdot E)$ is the normal trace of $E$. Similarly we deduce that div $B=0$ and $(n \cdot B)$ is the normal trace of $B$. By using the Green formula (2.9) with the test function $\theta=\theta(t, x), \theta \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega}\right), T$ periodic, we have:

$$
\begin{aligned}
\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} \alpha_{k} f_{k} \theta d t d x d p- & \int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}} f_{k}\left(\partial_{t} \theta+v(p) \cdot \nabla_{x} \theta\right) d t d x d p \\
& +\int_{0}^{T} \int_{\Sigma}(v(p) \cdot n(x)) \theta \gamma f_{k} d t d \sigma d p=0
\end{aligned}
$$

By using that
$\left(\int_{0}^{T} \int_{\Omega} \int_{\mathbb{R}_{p}^{3}}(1+\mathcal{E}(p)) f_{k} d t d x d p\right)_{k}, \quad\left(\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x))(1+\mathcal{E}(p)) \gamma^{+} f_{k} d t d \sigma d p\right)_{k} \quad$ are bounded, after passing to the limit for $k \rightarrow+\infty$ we deduce that
$-\int_{0}^{T} \int_{\Omega} \rho \partial_{t} \theta d t d x-\int_{0}^{T} \int_{\Omega} j \cdot \nabla_{x} \theta d t d x+\int_{0}^{T} \int_{\Sigma}(v(p) \cdot n(x)) \theta q \gamma f d t d \sigma d p=0$ and in particular $\partial_{t} \rho+\operatorname{div} j=0$ in $\mathcal{D}^{\prime}$.

## 8. Final remarks

By using basically the same arguments, it is possible to analyze also the time periodic Vlasov-Maxwell system when the boundary condition (2.2) is replaced by:

$$
\begin{equation*}
f(t, x, p)=g(t, x, p)+a(t, x, p) \cdot f(t, x, R(t, x) p),(t, x, p) \in \mathbb{R}_{t} \times \Sigma^{-} \tag{8.1}
\end{equation*}
$$

with $\quad R(t, x): \mathbb{R}_{p}^{3} \rightarrow \mathbb{R}_{p}^{3}, R(t, x) p=p-2(p \cdot n(x)) n(x), \forall(t, x, p) \in \mathbb{R}_{t} \times \Sigma \quad$ and $0 \leq a(t, x, p) \leq a_{0}<1, \forall(t, x, p) \in \mathbb{R}_{t} \times \Sigma^{-} \quad$ (for the definition of $T$ periodic weak solution for the problem (2.1), (8.1) consider in the weak formulation (2.4) test functions $\varphi \in C^{1}\left(\mathbb{R}_{t} \times \bar{\Omega} \times \mathbb{R}_{p}^{3}\right)$, $T$ periodic, compactly supported in momentum such
that $\left.\varphi(t, x, R p)=a(t, x, p) \varphi(t, x, p), \forall(t, x, p) \in \mathbb{R}_{t} \times \Sigma^{-}\right)$. We assume that $g, h$ are $T$ periodic, $g \geq 0, g \in L^{\infty}\left(\mathbb{R}_{t} \times \Sigma^{-}\right),\left.(n \cdot h)\right|_{\mathbb{R}_{t} \times \partial \Omega}=0$ such that:

$$
W_{0}=\int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))|(1+\mathcal{E}(p)) g(t, x, p) d t d \sigma d p+\int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma<+\infty
$$

The proofs are quite similar and we don't detail them. We only indicate how a priori estimates can be obtained in this case by formal computations. For example, by using the weak formulation of the Vlasov problem with the test function 1 we get:

$$
\int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdot n(x)) \gamma^{+} f(t, x, p) d t d \sigma d p=-\int_{0}^{T} \int_{\Sigma^{-}}(v(p) \cdot n(x)) \gamma^{-} f(t, x, p) d t d \sigma d p
$$

and by using the boundary condition (8.1) we deduce that:

$$
\begin{aligned}
& \int_{0}^{T} \int_{\Sigma^{+}}(v(p) \cdotn(x))(1-a(t, x, R p)) \gamma^{+} f(t, x, p) d t d \sigma d p \\
&=-\int_{0}^{T} \int_{\Sigma^{-}}(v(p) \cdot n(x)) g(t, x, p) d t d \sigma d p
\end{aligned}
$$

Since $a(t, x, p) \leq a_{0}<1, \forall(t, x, p) \in \mathbb{R}_{t} \times \Sigma^{-}$, finally one gets:

$$
\int_{0}^{T} \int_{\Sigma^{ \pm}}|(v(p) \cdot n(x))| \gamma^{ \pm} f(t, x, p) d t d \sigma d p \leq \frac{1}{1-a_{0}} \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| g(t, x, p) d t d \sigma d p
$$

Similarly, by using the weak formulation of the Vlasov problem with the test function $\mathcal{E}(p)$ and by combining with the conservation of the electro-magnetic energy (obtained by multiplication of the Maxwell equations by $(E, B)$ ), we deduce as before that:

$$
\begin{gathered}
\int_{0}^{T} \int_{\Sigma^{+}} \quad(v(p) \cdot n(x)) \mathcal{E}(p) \gamma^{+} f(t, x, p) d t d \sigma d p+\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}|n \wedge E|^{2}+\frac{1}{\mu_{0}}|n \wedge B|^{2}\right\} d t d \sigma \\
=-\int_{0}^{T} \int_{\Sigma^{-}}(v(p) \cdot n(x)) \mathcal{E}(p) \gamma^{-} f(t, x, p) d t d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma,
\end{gathered}
$$

and by using the boundary condition (8.1) we obtain as above that:

$$
\begin{aligned}
\left(1-a_{0}\right) \int_{0}^{T} \int_{\Sigma^{+}} & |(v(p) \cdot n(x))| \mathcal{E}(p) \gamma^{+} f(t, x, p) d t d \sigma d p \\
& +\frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}|n \wedge E|^{2}+\frac{1}{\mu_{0}}|n \wedge B|^{2}\right\} d t d \sigma \\
\leq & -\int_{0}^{T} \int_{\Sigma^{-}}(v(p) \cdot n(x)) \mathcal{E}(p) g(t, x, p) d t d \sigma d p+\frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma
\end{aligned}
$$

and:

$$
\begin{aligned}
& \left(1-a_{0}\right) \int_{0}^{T} \int_{\Sigma^{-}}|(v(p) \cdot n(x))| \mathcal{E}(p) \gamma^{-} f(t, x, p) d t d \sigma d p \\
& \quad+a_{0} \frac{c_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}\left\{\varepsilon_{0}|n \wedge E|^{2}+\frac{1}{\mu_{0}}|n \wedge B|^{2}\right\} d t d \sigma \\
& \leq-\int_{0}^{T} \int_{\Sigma^{-}}(v(p) \cdot n(x)) \mathcal{E}(p) g(t, x, p) d t d \sigma d p+a_{0} \frac{c_{0} \varepsilon_{0}}{2} \int_{0}^{T} \int_{\partial \Omega}|h(t, x)|^{2} d t d \sigma .
\end{aligned}
$$

From this point the computations follow exactly as for the case of absorbing conditions.
Note also that all these results apply for the Vlasov-Maxwell system with several species of particles. We point out that similar a priori estimates can be established by formal computations for the $T$ periodic solutions of the classical Vlasov-Maxwell-Fokker-Planck system, which is obtained from the classical Vlasov-Maxwell system by replacing the Vlasov equation by:

$$
\begin{aligned}
& \partial_{t} f+v(p) \cdot \nabla_{x} f+q(E(t, x)+v(p) \wedge B(t, x)) \cdot \nabla_{p} f \\
& \quad=\operatorname{div}_{p}\left(\sigma \nabla_{p} f+\beta v(p) f\right),(t, x, p) \in \mathbb{R}_{t} \times \Omega \times \mathbb{R}_{p}^{3}
\end{aligned}
$$

where $\beta \geq 0, \sigma>0$ are fixed parameters.

## REFERENCES

[1] A. Arseneev, Global existence of a weak solution of the Vlasov system of equations, U. R. S. S. Comp. and Math. Phys., 15, 131-143, 1975.
[2] J. Batt, Global symmetric solutions of the initial value problem of stellar dynamics, J. Diff. Eq., 25, 342-364, 1977.
[3] V. Barbu, Nonlinear Semigroups and Differential Equations in Banach Spaces, Noordhoff, 1976.
[4] C. Bardos, Problèmes aux limites pour les équations aux dérivées partielles du premier ordre, Ann. Sci. Ecole Norm. Sup., 3, 185-233, 1969.
[5] C. Bardos and P. Degond, Global existence for the Vlasov-Poisson equation in three space variables with small initial data, Ann. Inst. H. Poincaré, Anal. non linéaire 2, 101-118, 1985.
[6] N. Ben Abdallah, Weak solutions of the initial-boundary value problem for the Vlasov-Poisson system, Math. Meth. Appl. Sci., 17, 451-476, 1994.
[7] M. Bostan and F. Poupaud, Periodic solutions of the $1 D$ Vlasov-Maxwell system with boundary conditions, Math. Methods Appl. Sci., 23, 1195-1221, 2000.
[8] M. Bostan, Permanent regimes for the $1 D$ Vlasov-Poisson system with boundary conditions, SIAM J. Math. Anal., 35, 922-948, 2003.
[9] M. Bostan, Convergence des solutions faibles du système de Vlasov-Maxwell stationnaire vers des solutions faibles du système de Vlasov-Poisson stationnaire quand la vitesse de la lumière tend vers l'infini, C. R. Acad. Sci. Paris, Ser. I, 340, 803-808, 2005.
[10] F. Bouchut, F. Golse and C. Pallard, Classical solutions and the Glassey-Strauss theorem for the 3D Vlasov-Maxwell system, Arch. Ration. Mech. Anal., 170, 1-15, 2003.
[11] H. Brezis, Analyse Fonctionnelle, Masson, 1998.
[12] P. Degond, Local existence of solutions of the Vlasov-Maxwell equations and convergence to the Vlasov-Poisson equations for infinite light velocity, Math. Methods Appl. Sci., 8, 533-558, 1986.
[13] P. Degond and P.-A. Raviart, An asymptotic analysis of the one-dimensional Vlasov-Poisson system: the Child-Langmuir law, Asymptotic Anal., 4, 187-214, 1991.
[14] R. J. Diperna and P. L. Lions, Global weak solutions of the Vlasov-Maxwell system, Comm. Pure Appl. Math. XVII, 729-757, 1989.
[15] R. J. Diperna and P.-L. Lions, Ordinary differential equations, transport theory and Sobolev spaces, Invent. Math., 98, 511-547, 1989.
[16] G. Duvaut and J.-L. Lions, Les Inéquations en Mécanique et en Physique, Dunod, 1972.
[17] R. Glassey and J. Schaeffer, Global existence for the relativistic Vlasov-Maxwell system with nearly neutral initial data, Comm. Math. Phy., 119, 353-384, 1988.
[18] R. Glassey and J. Schaeffer, On the 'one and one-half dimensional' relativistic Vlasov-Maxwell system, Math. Meth. Appl. Sci., 13, 169-179, 1990.
[19] R. Glassey and W. Strauss, Singularity formation in a collisionless plasma could only occur at high velocities, Arch. Rat. Mech. Anal., 92, 56-90, 1986.
[20] R. Glassey and W. Strauss, Large velocities in the relativistic Vlasov-Maxwell equations, J. Fac. Sci. Tokyo, 36, 615-627, 1989.
[21] F. Golse, P.-L. Lions, B. Perthame and R. Sentis, Regularity of the moments of the solution of a transport equation, J. Funct. Anal., 88, 110-125, 1988.
[22] C. Greengard and P.-A. Raviart, A boundary value problem for the stationary Vlasov-Poisson equations: the plane diode, Comm. Pure and Appl. Math. vol. XLIII, 473-507, 1990.
[23] Y. Guo, Global weak solutions of the Vlasov-Maxwell system with boundary conditions, Commun. Math. Phys. 154, 245-263, 1993.
[24] Y. Guo, Singular solutions to the Vlasov-Maxwell system in a half line, Arch. Rational Mech. Anal., 131, 241-304, 1995.
[25] E. Horst, On the classical solutions of the initial value problem for the unmodified nonlinear Vlasov equation, Math. Meth. Appl. Sci., 3, 229-248, 1981.
[26] E. Horst and R. Hunze, Weak solutions of the initial value problem for the unmodified nonlinear Vlasov equation, Math. Meth. Appl. Sci., 6, 262-279, 1984.
[27] S. Klainerman and G. Staffilani, A new approach to study the Vlasov-Maxwell system, Commun. Pure Appl. Anal., 1, 103-125, 2002.
[28] V. Komornik, Exact Controllability and Stabilization: the Multiplier Method, RAM, Masson, Paris, John Wiley \& Sons, Ltd., Chichester, 1994.
[29] J. E. Lagnese, A singular perturbation problem in exact controllability of the Maxwell system, ESAIM, Control. Optim. Calc. Var., 6, 275-289, 2001.
[30] P.-L. Lions and B. Perthame, Propagation of moments and regularity for the 3-dimensional Vlasov-Poisson system, Invent. Math., 105, 415-430, 1991.
[31] K. Pfaffelmoser, Global classical solutions of the Vlasov-Poisson system in three dimensions for general initial data, J. Diff. Eq., 95, 281-303, 1992.
[32] F. Poupaud, Boundary value problems for the stationary Vlasov-Maxwell system, Forum Math., 4, 499-527, 1992.
[33] J. Schaeffer, Global existence of smooth solutions to the Vlasov-Poisson system in three dimensions, Comm. P.D.E., 16, 1313-1335, 1991.
[34] J. Schaeffer, Global existence for the Vlasov-Poisson system with nearly symetric data, J. Diff. Eq., 69, 111-148, 1987.
[35] T. Ukai and S. Okabe, On the classical solution in the large time of the two dimensional Vlasov equations, Osaka J. Math., 15, 245-261, 1978.


[^0]:    *Received: June 17; accepted (in revised version); October 24, 2005. Communicated by Francois Golse.
    † Université de Franche-Comté, 16 route de Gray F-25030 Besançon Cedex, tel: (33)(0)381666338, fax: $(33)(0) 381666623$, France (mbostan@math.univ-fcomte.fr).

