

**149. A Necessary Unitary Field Theory as a Non-Holonomic
Parabolic Lie Geometry Realized in the Three-
Dimensional Cartesian Space. II**

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The purpose of the present paper consists in the following five points: to deduce (i), (ii), (ii'), (iii), (iii') mentioned below from Part I (these Proc., 29 (1953)).

Since the three-dimensional non-holonomic Laguerre (parabolic Lie) geometry is in law a four-dimensional teleparallelism geometry keeping the Riemann (Weyl) metric, it is remarkable that the following conjecture of Prof. Einstein of 1928, which seems to be now scarcely considered, must acquire a renaissance: "Es ist denkbar, dass diese Theorie die ursprüngliche Fassung der allgemeinen Relativitätstheorie verdrängen wird".

(i) A Unitary Field Theory of a Single Particle

6. *A Necessary Unitary Field Theory of a Single Particle Charged with Rest-mass m_0 and Constant Electricity $-e$.* In Art. 4, we have solved a two particles problem stated in Art. 2 and the resulting generalizations of the Maxwell's equations were (4.24), (4.25), (4.26) and (4.27). Thereby the continuity condition (4.23) was assumed. Now in the case of a single particle P , we have

$$(6.1) \quad \bar{m}_0 = \bar{e} = \bar{\epsilon}^t = \bar{\epsilon}^4 = \bar{\sigma}^t = \bar{\sigma}^4 = \bar{\mathcal{X}}^t = \bar{X}^t = \bar{a}^t = \bar{a}^4 = 0.$$

Hence (4.24), (4.25), (4.26) and (4.27) become the *necessary-unitary-field-theoretical generalization of the Maxwell's equations*:

$$(6.2) \quad \frac{\partial}{\omega^i} (\mathcal{X}^i + eX^i) = \epsilon^i + \sigma^i, \frac{\partial}{\omega^4} (a^t + ea^t) + \frac{\partial}{\omega^j} (\mathcal{X}^k + eX^k) - \frac{\partial}{\omega^k} (\mathcal{X}^j + eX^j) = 0,$$

$$(6.2') \quad \frac{\partial}{\omega^j} (a^k + ea^k) - \frac{\partial}{\omega^k} (a^j + ea^j) - \frac{\partial}{\omega^4} (\mathcal{X}^i + eX^i) = \epsilon^i + \sigma^i, \frac{\partial}{\omega^i} (a^t + ea^t) = 0.$$

7. *General-Relativistic and Necessary-Unitary-Field-Theoretical Generalization of the Dirac Equation in the Case of a Single Particle.* In the case of a single particle P , the *general-relativistic and necessary-unitary-field-theoretical generalizations*

$$(5.3) \quad \left[\gamma_t \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^t} + e\phi^t + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^t} + \bar{e}\bar{\phi}^t \right) + \gamma_4 \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + e\phi^4 + \bar{m}_0 \bar{E}^2 \right) \right. \\ \left. + \gamma_5 \left(\frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^5} + \bar{e}\bar{\phi}^5 + m_0 E^2 \right) \right] \psi = 0,$$

$$(7.1) \quad \left[\gamma_i \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^i} + \bar{e}\bar{\phi}^i \right) + \gamma_5 \left[\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^5} + e\phi^5 + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^5} + \bar{e}\bar{\phi}^5 \right] \right] \psi = 0$$

of the Dirac equation become

$$(7.2) \quad \left[\gamma_i \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i \right) + \gamma_5 m_0 E^2 \right] \psi = 0,$$

$$(7.3) \quad \left[\gamma_i \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i \right) + \gamma_5 \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^5} + e\phi^5 \right) \right] \psi = 0,$$

$$(7.4) \quad \psi \equiv -\gamma_i \gamma_i (\mathcal{X}^i + eX^i) + \sum \gamma_j \gamma_k (a^j + ea^k).$$

8. *General-Relativistic and Necessary-Unitary-Field-Theoretical Generalization of the Schrödinger Equation in the Case of a Single Particle.* In the case of a single particle P , the general-relativistic and necessary-unitary-field-theoretical generalizations

$$(8.1) \quad \left[\left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^i} + \bar{e}\bar{\phi}^i \right)^2 - \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + e\phi^4 + \bar{m}_0 \bar{E}^2 \right)^2 + \left(\frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^5} + \bar{e}\bar{\phi}^5 + m_0 E^2 \right)^2 \right] \psi = 0,$$

$$(8.2) \quad \left[\left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^i} + \bar{e}\bar{\phi}^i \right)^2 - \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + e\phi^4 + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^4} + \bar{e}\bar{\phi}^4 \right)^2 + \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^5} + e\phi^5 + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^5} + \bar{e}\bar{\phi}^5 \right)^2 \right] \psi = 0$$

become

$$(8.3) \quad \left[\left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i \right)^2 + (m_0 E^2)^2 \right] \psi = 0,$$

$$(8.4) \quad \left[\left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + e\phi^4 \right)^2 - \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + e\phi^4 \right)^2 + \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^5} + e\phi^5 \right)^2 \right] \psi = 0.$$

(ii) An Exact Gravitational (ii') An Electromagnetic
Wave Theory of Two Particles

9. Gravitational Wave Theories

A. S. Eddington [1]¹⁾

Prof. A. Einstein [4]

has given an approximative wave theory of gravity within the general relativity. teleparallelism geometry.

10. *Problem Formulation.* Consider two particles O and P charged with constant

1) The ciphers in the square brackets refer throughout this paper to the References at the end of this paper.

rest-masses \bar{m}_0 and m_0 electricity $-\bar{e}$ and $-e$ respectively, which move relative to each other. Then both O and P emit gravitational P emit electromagnetic energy radially in such a manner that the action is non-holonomic, the energy levels being spherical. The law of motion is required.

Solution. Introducing our conditions:

$$(10.1) \quad \begin{cases} \phi^t=0, \phi^4=0, \phi^5 \neq 0, \bar{\phi}^t=0, & \phi^t \neq 0, -\phi^5=0, -\phi^4 \neq 0, \bar{\phi}^t \neq 0, \\ \bar{\phi}^4 \neq 0, \bar{\phi}^5=0, E\phi^5 \neq 0, & \bar{\phi}^4=0, \bar{\phi}^5 \neq 0, Ep^4 \neq 0, \\ \bar{E}\bar{p}^4 \neq 0, Ep^5 \neq 0, \bar{E}\bar{p}^5=0, & \bar{E}\bar{p}^4=0, Ep^5=0, \bar{E}\bar{p}^5 \neq 0, \end{cases}$$

$$(10.2) \quad (Ep^t + \bar{E}\bar{p}^t) \quad (Ep^t + e\phi^t + \bar{E}\bar{p}^t + \bar{e}\bar{\phi}^t) \\ = (mE + \bar{m}\bar{E}) \frac{d\xi^t}{dt}, \quad \left(m = m_0 \frac{dr}{dS}, \quad \bar{m} = \bar{m}_0 \frac{d\bar{S}}{dr} \right),$$

$$(10.3) \quad (\bar{E}\bar{p}^4 + \bar{e}\bar{\phi}^4) = (mE + \bar{m}\bar{E}) \frac{d\xi^4}{dt}, \quad (Ep^4 + e\phi^4) = (mE + \bar{m}\bar{E}) \frac{d\xi^4}{dt},$$

$$(10.4) \quad (Ep^5 + e\phi^5) = (mE + \bar{m}\bar{E}) \frac{d\xi^5}{dt}, \quad (\bar{E}\bar{p}^5 + \bar{e}\bar{\phi}^5) = (mE + \bar{m}\bar{E}) \frac{d\xi^5}{dt}.$$

Introducing these values into $(mE + \bar{m}\bar{E})$ -times of

$$(10.5) \quad -i\gamma_5\omega^5/dt = \gamma_t\omega^t/dt,$$

$$(10.6) \quad \gamma_i(Ep^t + \bar{E}\bar{p}^t) + \gamma_4(\bar{E}\bar{p}^4 + \bar{e}\bar{\phi}^4) \quad \gamma_i(Ep^t + e\phi^t + \bar{E}\bar{p}^t + \bar{e}\bar{\phi}^t) + \gamma_4(Ep^4 \\ = -i\gamma_5(Ep^5 + e\phi^5), \quad + e\phi^4) = -i\gamma_5(\bar{E}\bar{p}^5 + \bar{e}\bar{\phi}^5),$$

which becomes

$$(10.7) \quad EP + e\psi + \bar{E}\bar{P} + \bar{e}\bar{\psi} = 0$$

for

$$(10.8) \quad \begin{cases} \gamma_i p^t + i\gamma_5 p^5 = P, \gamma_i \bar{p}^t + \gamma_4 \bar{p}^4 = \bar{P}, \\ i\gamma_5 \phi^5 = \psi, \gamma_4 \bar{e} \bar{p}^4 = \bar{\psi}. \end{cases} \quad \left| \begin{cases} \gamma_i p^t + \gamma_4 p^4 = P, \gamma_i \bar{p}^t + i\gamma_5 \bar{p}^5 = \bar{P}, \\ \gamma_i \phi^t + \gamma_4 \phi^4 = \psi, \gamma_i \bar{\phi}^t + i\gamma_5 \bar{\phi}^5 = \bar{\psi}. \end{cases} \right.$$

Applying the operator (4.17) to (10.7), we obtain

$$(10.9) \quad 2\gamma_5 \frac{\partial}{\omega^5} (EP + e\psi + \bar{E}\bar{P} + \bar{e}\bar{\psi}) = \frac{\partial}{\omega^t} (Ep^t + e\phi^t + \bar{E}\bar{p}^t + \bar{e}\bar{\phi}^t) \\ - \gamma_4 \gamma_i (\mathcal{X}^t + eX^t + \bar{\mathcal{X}}^t + \bar{e}\bar{X}^t) + \sum \gamma_j \gamma_k (a^t + ea^t + \bar{a}^t + \bar{e}\bar{a}^t) \\ + 2i \frac{\partial}{\omega^5} (Ep^5 + e\phi^5 + \bar{E}\bar{p}^5 + \bar{e}\bar{\phi}^5) = 0.$$

In the present case, we have

$$(10.10) \quad \mathcal{X}^t = \frac{\partial(Ep^t)}{\omega^4}, \quad \mathcal{X}^t = \frac{\partial(Ep^4)}{\omega^t} + \frac{\partial(Ep^t)}{\omega^4} = 0,$$

$$(10.11) \quad \bar{\mathcal{X}}^t = i \frac{\partial(\bar{E}\bar{p}^t)}{\omega^5}, \quad \bar{\mathcal{X}}^t = \frac{\partial(\bar{E}\bar{p}^5)}{\omega^t} + i \frac{\partial(\bar{E}\bar{p}^t)}{\omega^5} = 0,$$

$$(10.12) \quad \alpha^i = \frac{\partial(Ep^k)}{\omega^j} - \frac{\partial(Ep^j)}{\omega^k},$$

$$(10.13) \quad \bar{\alpha}^i = \frac{\partial(\bar{E}\bar{p}^k)}{\omega^j} - \frac{\partial(\bar{E}\bar{p}^j)}{\omega^k},$$

$$(10.14) \quad X^i = 0, \quad \bar{X}^i = 0,$$

$$(10.15) \quad \mathcal{X}^i = \frac{\partial(Ep^i)}{\omega^4}, \quad \bar{\mathcal{X}}^i = i \frac{\partial(\bar{E}\bar{p}^i)}{\omega^5},$$

$$(10.16) \quad \alpha^i = 0, \quad \bar{\alpha}^i = 0$$

$$\alpha^i = \frac{\partial(Ep^k)}{\omega^j} - \frac{\partial(Ep^j)}{\omega^k},$$

$$\bar{\alpha}^i = \frac{\partial(\bar{E}\bar{p}^k)}{\omega^j} - \frac{\partial(\bar{E}\bar{p}^j)}{\omega^k} = 0,$$

$$X^i = \frac{\partial\phi^4}{\omega^i} + \frac{\partial\phi^i}{\omega^4}, \quad \bar{X}^i = i \frac{\partial\bar{\phi}^i}{\omega^5},$$

$$\mathcal{X}^i = 0, \quad \bar{\mathcal{X}}^i = 0,$$

$$\alpha^i = \frac{\partial\phi^k}{\omega^j} - \frac{\partial\phi^j}{\omega^k}, \quad \bar{\alpha}^i = \frac{\partial\bar{\phi}^k}{\omega^j} - \frac{\partial\bar{\phi}^j}{\omega^k}$$

and (10.9):

$$(10.17) \quad \begin{aligned} & 2\gamma_5 \frac{\partial}{\omega^5} (EP + e\bar{\Psi} + \bar{E}\bar{P} + \bar{e}\bar{\Psi}) \\ &= \frac{\partial}{\omega^i} (Ep^i + \bar{E}\bar{p}^i + \bar{e}\bar{\phi}^i) - \gamma_4 \gamma_i (\mathcal{X}^i + \bar{\mathcal{X}}^i) + \sum \gamma_j \gamma_k (\alpha^i + \bar{\alpha}^i) + 2i \frac{\partial}{\omega^5} (Ep^5 + ep^5) = 0. \end{aligned} \quad \left| \begin{aligned} &= \frac{\partial}{\omega^i} (Ep^i + e\phi^i + \bar{E}\bar{p}^i + \bar{e}\bar{\phi}^i) - \gamma_4 \gamma_i (eX^i + \bar{e}\bar{X}^i) + \sum \gamma_j \gamma_k (e\alpha^i + \bar{e}\bar{\alpha}^i) + 2i \frac{\partial}{\omega^5} (\bar{E}\bar{p}^5 + \bar{e}\bar{\phi}^5) = 0. \end{aligned} \right.$$

Introducing the continuity condition

$$(10.18) \quad \begin{aligned} & \frac{\partial}{\omega^i} (Ep^i + \bar{E}\bar{p}^i) + \frac{\partial}{\omega^4} (\bar{E}\bar{p}^4 + \bar{e}\bar{\phi}^4) + 2i \frac{\partial}{\omega^5} (Ep^5 + ep^5) = 0 \end{aligned} \quad \left| \begin{aligned} & \frac{\partial}{\omega^i} (Ep^i + ep^i) + \frac{\partial}{\omega^i} (\bar{E}\bar{p}^i + \bar{e}\bar{\phi}^i) + 2i \frac{\partial}{\omega^5} (\bar{E}\bar{p}^5 + \bar{e}\bar{\phi}^5) = 0 \end{aligned} \right.$$

and then applying the operator (4.17) to (10.18) once more, we obtain the *generalized Maxwell's equations*

$$(10.19) \quad \frac{\partial}{\omega^i} (\mathcal{X}^i + \bar{\mathcal{X}}^i) = \varepsilon^4 + \bar{\varepsilon}^4, \quad \left| \frac{\partial}{\omega^i} (eX^i + \bar{e}\bar{X}^i) = \sigma^4 + \bar{\sigma}^4, \right.$$

$$(10.20) \quad \begin{aligned} & \frac{\partial}{\omega^i} (\alpha^i + \bar{\alpha}^i) + \frac{\partial}{\omega^j} (\mathcal{X}^k + \bar{\mathcal{X}}^k) - \frac{\partial}{\omega^k} (\mathcal{X}^j + \bar{\mathcal{X}}^j) = 0, \end{aligned} \quad \left| \begin{aligned} & \frac{\partial}{\omega^i} (e\alpha^i + \bar{e}\bar{\alpha}^i) + \frac{\partial}{\omega^j} (eX^k + \bar{e}\bar{X}^k) - \frac{\partial}{\omega^k} (eX^j + \bar{e}\bar{X}^j) = 0, \end{aligned} \right.$$

$$(10.21) \quad \begin{aligned} & \frac{\partial}{\omega^j} (\alpha^k + \bar{\alpha}^k) - \frac{\partial}{\omega^4} (\alpha^j + \bar{\alpha}^j) - \frac{\partial}{\omega^k} (\mathcal{X}^i + \bar{\mathcal{X}}^i) = \varepsilon^i + \bar{\varepsilon}^i, \end{aligned} \quad \left| \begin{aligned} & \frac{\partial}{\omega^i} (e\alpha^k + \bar{e}\bar{\alpha}^k) - \frac{\partial}{\omega^k} (e\alpha^j + \bar{e}\bar{\alpha}^j) - \frac{\partial}{\omega^4} (eX^i + \bar{e}\bar{X}^i) = \sigma^i + \bar{\sigma}^i, \end{aligned} \right.$$

$$(10.22) \quad \frac{\partial}{\omega^i} (\alpha^i + \bar{\alpha}^i) = 0 \quad \left| \frac{\partial}{\omega^i} (e\alpha^i + \bar{e}\bar{\alpha}^i) = 0 \right.$$

for the two particles *O* and *P*.

11. *General-Relativistic Analogue to the Dirac Equation for the Case of Gravitation* | *Case of Electromagnetism*
for Two Particles. The (5.1) becomes

$$\begin{aligned}
 (11.1) \quad & \psi \\
 & = -\gamma_4 \gamma_i (\mathcal{X}^i + \bar{\mathcal{X}}^i) + \sum \gamma_j \gamma_k (a^i + \bar{a}^i) \quad \left| \quad = -\gamma_4 \gamma_i (eX^i + \bar{e}\bar{X}^i) \right. \\
 & \text{and (5.3) and (7.1) to} \quad \left. + \sum \gamma_j \gamma_k (ea^i + \bar{e}\bar{a}^i) \right. \\
 (11.2) \quad & \left[\gamma_i (E + \bar{E}) \frac{\hbar}{2\pi i} \frac{\partial}{\omega^i} \right. \quad \left. \left[\gamma_i \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^i} + \bar{e}\bar{\phi}^i \right) \right. \right. \\
 & + \gamma_4 \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + \bar{m}_0 \bar{E}^2 \right) \quad \left. + \gamma_4 \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + e\phi^4 \right) + \gamma_5 \left(\frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^5} \right. \right. \\
 & \left. \left. + \gamma_5 \left(\frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^5} + m_0 E^2 \right) \right] \psi = 0, \quad \left. + \bar{e}\bar{\phi}^5 \right] \psi = 0, \\
 (11.3) \quad & \gamma_i \left[\frac{\hbar}{2\pi i} (E + \bar{E}) \frac{\partial}{\omega^i} \right. \quad \left. \left[\gamma_i \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^i} + \bar{e}\bar{\phi}^i \right) \right. \right. \\
 & + \gamma_4 \left\{ \frac{\hbar}{2\pi i} (E + \bar{E}) \frac{\partial}{\omega^4} + \bar{e}\bar{\phi}^4 \right\} \quad \left. + \gamma_4 \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + e\phi^4 + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^4} \right) \right. \\
 & \left. + \gamma_5 \left\{ \frac{\hbar}{2\pi i} (E + \bar{E}) \frac{\partial}{\omega^5} + e\phi^4 \right\} \right] \psi = 0. \quad \left. + \gamma_5 \left\{ (E + \bar{E}) \frac{\hbar}{2\pi i} \frac{\partial}{\omega^5} + \bar{e}\bar{\phi}^5 \right\} \right] \psi = 0.
 \end{aligned}$$

12. *General Relativistic Analogue to the Schrödinger Equation for the Gravitation* | *for the Electromagnetism*
for Two Particles. (8.1) and (8.2) become our *general-relativistic analogues to the Schrödinger equation*:

$$\begin{aligned}
 (12.1) \quad & \left[-\frac{\hbar^2}{4\pi^2} (E + \bar{E})^2 \left(\frac{\partial}{\omega^i} \right)^2 \right. \quad \left. \left[\left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^i} + \bar{e}\bar{\phi}^i \right)^2 \right. \right. \\
 & - \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + \bar{m}_0 \bar{E}^2 \right)^2 \quad \left. - \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + e\phi^4 \right)^2 + \left(\frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^5} \right. \right. \\
 & \left. \left. + \left(\frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^5} + m_0 E^2 \right)^2 \right] \psi = 0, \quad \left. + \bar{e}\bar{\phi}^5 \right)^2 \right] \psi = 0, \\
 (12.2) \quad & \left[-\frac{\hbar^2}{4\pi^2} (E + \bar{E})^2 \left(\frac{\partial}{\omega^i} \right)^2 \right. \quad \left. \left[\left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^i} + \bar{e}\bar{\phi}^i \right)^2 \right. \right. \\
 & - \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^4} + \bar{e}\bar{\phi}^4 \right)^2 \quad \left. - \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + e\phi^4 + \frac{\hbar}{2\pi i} \bar{E} \frac{\partial}{\omega^4} \right)^2 \right. \\
 & \left. + \left(\frac{\hbar}{2\pi i} (E + \bar{E}) \frac{\partial}{\omega^5} + e\phi^5 \right)^2 \right] \psi = 0. \quad \left. + \left(\frac{\hbar}{2\pi i} (E + \bar{E}) \frac{\partial}{\omega^5} + \bar{e}\bar{\phi}^5 \right)^2 \right] \psi = 0.
 \end{aligned}$$

(iii) An Exact Gravitational | (iii') An Electromagnetic
 Wave Theory of a Single Particle

13. *Problem Formulation.* Consider a single particle P charged with constant rest-mass m_0 , | with constant electricity $-e$,

which makes a motion emitting
gravitational

electromagnetic

energy radially in such a manner, that the action is non-holonomic,
the energy level being spherical. The law of motion is required.

Solution. Introducing the condition that the particle O has no
rest-mass ($\bar{m}_0=0$)

electric charge ($e=0$)

into (10.19), (10.20), (10.21) and (10.22), we obtain

$$(13) \quad \left. \begin{aligned} \frac{\partial \mathcal{X}^i}{\omega^i} = \varepsilon^4, \quad \frac{\partial \alpha^i}{\omega^i} + \frac{\partial \mathcal{X}^k}{\omega^j} \\ - \frac{\partial \mathcal{X}^j}{\omega^k} = 0, \quad \frac{\partial \alpha^k}{\omega^j} - \frac{\partial \alpha^j}{\omega^k} \\ - \frac{\partial \mathcal{X}^i}{\omega^4} = 0, \quad \frac{\partial \alpha^i}{\omega^i} = 0. \end{aligned} \right| \begin{aligned} \frac{\partial X^i}{\omega^i} = \frac{\sigma^4}{e}, \quad \frac{\partial \alpha^i}{\omega^i} + \frac{\partial X^k}{\omega^j} \\ - \frac{\partial X^j}{\omega^k} = 0, \quad \frac{\partial \alpha^k}{\omega^j} - \frac{\partial \alpha^j}{\omega^k} \\ - \frac{\partial X^i}{\omega^4} = \frac{\sigma^i}{e}, \quad \frac{\partial \alpha^i}{\omega^i} = 0. \end{aligned}$$

The lefthand side is

The righthand side is

*gravitational analogues
of the Maxwell's equations.*

general-relativistic generalizations

14. *General-Relativistic Analogues to the Dirac Equations for
the Case of Gravitation* | *the Case of Electromagnetism*
for a Single Particle. For the case of a single particle P , (11.2)
and (11.3) becomes respectively to

$$(14.1) \quad \left[\gamma_t E \frac{\hbar}{2\pi i} \frac{\partial}{\omega^i} + \gamma_s m_0 E^2 \right] \psi = 0, \quad \left[\gamma_t \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i \right) \right] \psi = 0,$$

$$(14.2) \quad \left[\gamma_i \frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + \gamma_4 \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} \right. \right. \left. \left. \left[\gamma_t \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i \right) + \gamma_5 \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^5} \right) \right] \right) \right. \right. \\ \left. \left. + \bar{e}\bar{\phi}^4 \right) + \gamma_5 \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^5} + e\phi^5 \right) \right] \psi = 0. \quad \psi = 0.$$

15. *General-Relativistic Analogue to the Schrödinger Equations
for the Case of* | *Electromagnetism*
Gravitation

for a Single Particle. For the case of a single particle P , (12.1)
and (12.2) become respectively to

$$(15.1) \quad \left(-\frac{\hbar^2}{4\pi^2} E^2 \left(\frac{\partial}{\omega^i} \right)^2 - \frac{\hbar^2}{4\pi^2} E^2 \left(\frac{\partial}{\omega^t} \right)^2 \right. \left. \left[\left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i \right)^2 - \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} \right. \right. \right. \\ \left. \left. \left. + e\phi^4 \right)^2 \right] \right) \psi = 0, \quad \left[\left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^i} + e\phi^i \right)^2 - \left(\frac{\hbar}{2\pi i} E \frac{\partial}{\omega^4} + e\phi^4 \right)^2 \right. \\ \left. - \frac{\hbar^2}{4\pi^2} E^2 \left(\frac{\partial}{\omega^5} \right)^2 \right] \psi = 0.$$

References

- [1] A. S. Eddington: *The Mathematical Theory of Relativity*. Cambridge (1923).
- [2] T. Takasu: A necessary unitary field theory as a non-holonomic parabolic Lie geometry realized in the three-dimensional Cartesian space and its quantum mechanics, *Yokohama Mathematical Journal*, **1**, No. 2, 263-273 (1933).
- [3] T. Takasu: A necessary unitary field theory as a non-holonomic parabolic Lie geometry realized in the three-dimensional Cartesian space, *Proc. Japan Acad*, **29**, No. 10, 533-536 (1953).
- [4] A. Einstein: Neue Möglichkeit für eine einheitliche Feldtheorie von Gravitation und Elektrizität, *Sitzungsber. d. preuss. Akad. d. Wiss. Phys. -Math. Kl.* (1928)