Commun. Math. Phys. 84, 329-331 (1982)

## Local Extensions in Singular Space-Times II

C. J. S. Clarke

Department of Mathematics, University of York, Heslington, York YOI 5DD, England

Abstract. Previous results of the author are corrected by reformulating them in space-times whose Riemann tensor satisfies a Hölder condition.

## 1. Introduction

In an earlier paper with this title [2] I showed the existence of local extensions through quasiregular singularities (in the terminology of [5]) by (implicitly) assuming that a spacetime with a  $C^{k-2}$  Riemann tensor had a  $C^k$  metric. This assumption may not be correct (the alleged proof which I gave in [3] being invalid). The basic results do hold, however, if one uses  $C^{k,\alpha}$  conditions (a Hölder condition with exponent  $\alpha$ ,  $0 < \alpha < 1$ , on the  $k^{\text{th}}$  derivative). The technical tools needed to modify the proof are given in detail in [4]; the aim of the present paper is to outline their application to local extensions.

We first clarify the term "local extension" of a spacetime (M, g), of which there are two definitions in the literature. Here, and in [5], it means an isometry  $\phi: U \to M'$ , where  $U \subset M$  and (M', g') is a spacetime, such that

(i) U contains a curve  $\gamma$  which is incomplete with respect to a generalised affine parameter and inextendible in M.

(ii)  $\phi \circ \gamma$  is extendible in M'.

Hawking and Ellis [6], on the other hand, replace (i) by the condition that  $\overline{U}$  is not compact in M, and (ii) by the condition that  $\phi(U)$  is compact in M'. This has the undesirable consequence that Minkowski space is locally extendible [1]. With the author's definition, certain compact space-times having trapped geodesics may be locally extendible.

## 2. Results

The theorem will be formulated for the case where  $\gamma$  in the definition above is a broken geodesic. Since any rectifiable curve can be approximated by a broken geodesic this is no loss of generality, and it enables us to give a concrete description of the set U that can be extended. In addition we impose a restriction ((iv) below) that corresponds to the non-spiral condition imposed in the earlier paper [2]. The theorem will only be proved for  $C^{0,\alpha}$  Riemann tensors; but it is clear that the procedure extends to  $C^{k,\alpha}$ .