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On the *b*-Boundary of the Closed Friedman-Model

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Abstract. Some points of the past Big Bang in the closed fourdimensional Friedman-model are found to be identical with points of the future collapse according to the bundle-boundary definition.

1. Introduction

Consider the closed Friedman-model (M, g) with metric

$$ds_g^2 = R^2(\psi) \{ d\psi^2 - d\sigma^2 - \sin^2 \sigma (d\vartheta^2 + \sin^2 \vartheta d\varphi^2) \}$$

with $R(\psi) = 1 - \cos \psi$,

with singularities at $\psi = 0$ and $\psi = 2\pi$. We shall investigate the structure of the *b*-boundary for this space-time by working with, rather than the ten-dimensional orthonormal bundle O(M) (see [1, 2]), a certain three-dimensional subbundle. The construction is as follows. Consider the timelike and totally geodesic two-dimensional submanifolds NcM with induced metric γ , given by

 $\vartheta = \text{const}$ and $\varphi = \text{const}$.

Moreover, there exists an orthonormal dyad field

$$W_{\alpha}, \quad \alpha = 2, 3$$

which is parallel along and orthogonal to N. Therefore we can construct a threedimensional submanifold $\tilde{N}cO(M)$, consisting of every orthonormal tetrad Y_i , i=0,...3 with

$Y_A \in T(N)$	A = 0, 1
$Y_{\alpha} = W_{\alpha}$	$\alpha = 2, 3$

at every point of *N*. \tilde{N} is isomorphic to O(N). Furthermore the induced metric in \tilde{N} is equal to the bundle metric $\tilde{\gamma}$ in O(N), because any curve in *N*, which is horizontal with respect to γ is horizontal with respect to *g* as well. The metric $\tilde{\gamma}$ can be easily computed. This reduction method can be applied also to other space-times, e.g. the Schwarzschild and Reissner-Nordström space-times. If we now find curves, which connect two points in the fibres of the two singularities with arbitrarily small length in $\overline{O(N)}^{-1}$, the Cauchy completion of $O(N)^{-1}$ [1],

¹ The prime denotes the connected component, i.e. here the manifolds consisting of every positively oriented orthonormal dyad resp. tetrad in every point of N res. M.