# The General Stationary Gravitational Vacuum Field of Cylindrical Symmetry 

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#### Abstract

The general stationary vacuum gravitational field of cylindrical symmetry as recently found by Davies and Caplan is even static. The possible Petrov types of the Riemann tensor are $I, D$ or $O$. In spacelike infinity the spacetime becomes necessarily flat.


## 1. Introduction

Levy and Robinson [1] have argued that for axisymmetric stationary systems and modulo the vacuum field equations

$$
\begin{equation*}
R_{\mu \nu}=0 \tag{1.1}
\end{equation*}
$$

there exists a canonical (cylindrical) coordinate system in which the line element takes the form

$$
\begin{align*}
d s^{2}=e^{2 u}(d t+a d \varphi)^{2} & -e^{2(k-u)}\left(d r^{2}+d z^{2}\right) \\
& -r^{2} e^{-2 u} d \varphi^{2} \tag{1.2}
\end{align*}
$$

$a=0$ corresponds to Weyl's canonical coordinates for the static case. If $u, k, a$ are functions of $r$ only, the line element represents the vacuum gravitational field within or outside an infinite, axially symmetric rotating cylindrical mass distribution and the field Eq. (1.1) reduce to

$$
\begin{gather*}
\frac{d^{2} u}{d r^{2}}+\frac{1}{r} \frac{d u}{d r}+\frac{1}{2 r^{2}} e^{4 u}\left(\frac{d a}{d r}\right)^{2}=0 \\
\frac{d^{2} a}{d r^{2}}-\frac{1}{r} \frac{d a}{d r}+4 \frac{d a}{d r} \frac{d u}{d r}=0  \tag{1.3}\\
\frac{2}{r} \frac{d k}{d r}-2\left(\frac{d u}{d r}\right)^{2}+\frac{1}{2 r^{2}} e^{4 u}\left(\frac{d a}{d r}\right)^{2}=0
\end{gather*}
$$

Recently Davies and Caplan [2] have found the general solution of (1.3), from which they deduced, that under the condition $u, a, k$ to be finite at $r=0$ the interior of the rotating cylinder is flat.

