

# Analysis of the Static Spherically Symmetric $SU(n)$ -Einstein-Yang-Mills Equations

H. P. Künzle

Department of Mathematics, University of Alberta, Edmonton, Canada T6G 2G1.  
E-mail: hkunzle@vega.math.ualberta.ca

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**Abstract:** The singular boundary value problem that arises for the static spherically symmetric  $SU(n)$ -Einstein-Yang-Mills equations in the so-called magnetic case is analyzed. Among the possible actions of  $SU(2)$  on a  $SU(n)$ -principal bundles over space-time there is one which appears to be the most natural. If one assumes that no electrostatic type component is present in the Yang-Mills fields and the gauge is suitably fixed a set of  $n - 1$  second order and two first order differential equations is obtained for  $n - 1$  gauge potentials and two metric components as functions of the radial distance. This system generalizes the one for the case  $n = 2$  that leads to the discrete series of the Bartnik-Mckinnon and the corresponding black hole solutions. It is highly nonlinear and singular at  $r = \infty$  and at  $r = 0$  or at the black hole horizon but it is known to admit at least one series of discrete solutions which are scaled versions of the  $n = 2$  case. In this paper local existence and uniqueness of solutions near these singular points is established which turns out to be a nontrivial problem for general  $n$ . Moreover, a number of new numerical soliton (i.e. globally regular) numerical solutions of the  $SU(3)$ -EYM equations are found that are not scaled  $n = 2$  solutions.

## 1. Introduction

The coupling of Einstein's general relativity with Yang-Mills gauge theories leads to complicated nonlinear systems of equations even in the static spherically symmetric case. If the gauge group is  $SU(2)$  and the "Coulomb" part of the gauge potential is set to zero and asymptotical flatness is imposed the resulting singular boundary value problem admits a sequence of regular solutions parametrized by the number of zeros of a convenient gauge potential component. These solutions were numerically discovered by Bartnik and Mckinnon [3] and their existence was proved analytically by Smoller et al. [18–20] for some range of the initial conditions for a suitable gauge potential at the center or at the black hole horizon. Such discrete sequences of solutions have since also been found numerically for a number of other field